



EXO

## *Status and perspectives*

*Giorgio Gratta  
Physics Dept, Stanford*

*LBL, Mar 19, 2009*

# Our knowledge of the lepton mixing matrix

$$U = \begin{pmatrix} U_{e1} & U_{e2} & U_{e3} \\ U_{\mu1} & U_{\mu2} & U_{\mu3} \\ U_{\tau1} & U_{\tau2} & U_{\tau3} \end{pmatrix} = \begin{pmatrix} 1 & 0 & 0 \\ 0 & \cos\theta_{23} & \sin\theta_{23} \\ 0 & -\sin\theta_{23} & \cos\theta_{23} \end{pmatrix} \cdot \begin{pmatrix} \cos\theta_{13} & 0 & e^{-i\delta_{CP}} \sin\theta_{13} \\ 0 & 1 & 0 \\ -e^{-i\delta_{CP}} \sin\theta_{13} & 0 & \cos\theta_{13} \end{pmatrix} \cdot \begin{pmatrix} \cos\theta_{12} & \sin\theta_{12} & 0 \\ -\sin\theta_{12} & \cos\theta_{12} & 0 \\ 0 & 0 & 1 \end{pmatrix} \cdot \begin{pmatrix} 1 & 0 & 0 \\ 0 & e^{-i\alpha/2} & 0 \\ 0 & 0 & e^{-i\alpha/2+i\beta} \end{pmatrix}$$

information on  $\delta_{CP}$  is here Neutrino factory ?

Atmospheric  $\nu$   
K2K  
Minos et al

$\theta_{23} \sim 45^\circ$

Chooz/Palo Verde  
Off axis  
Future reactors

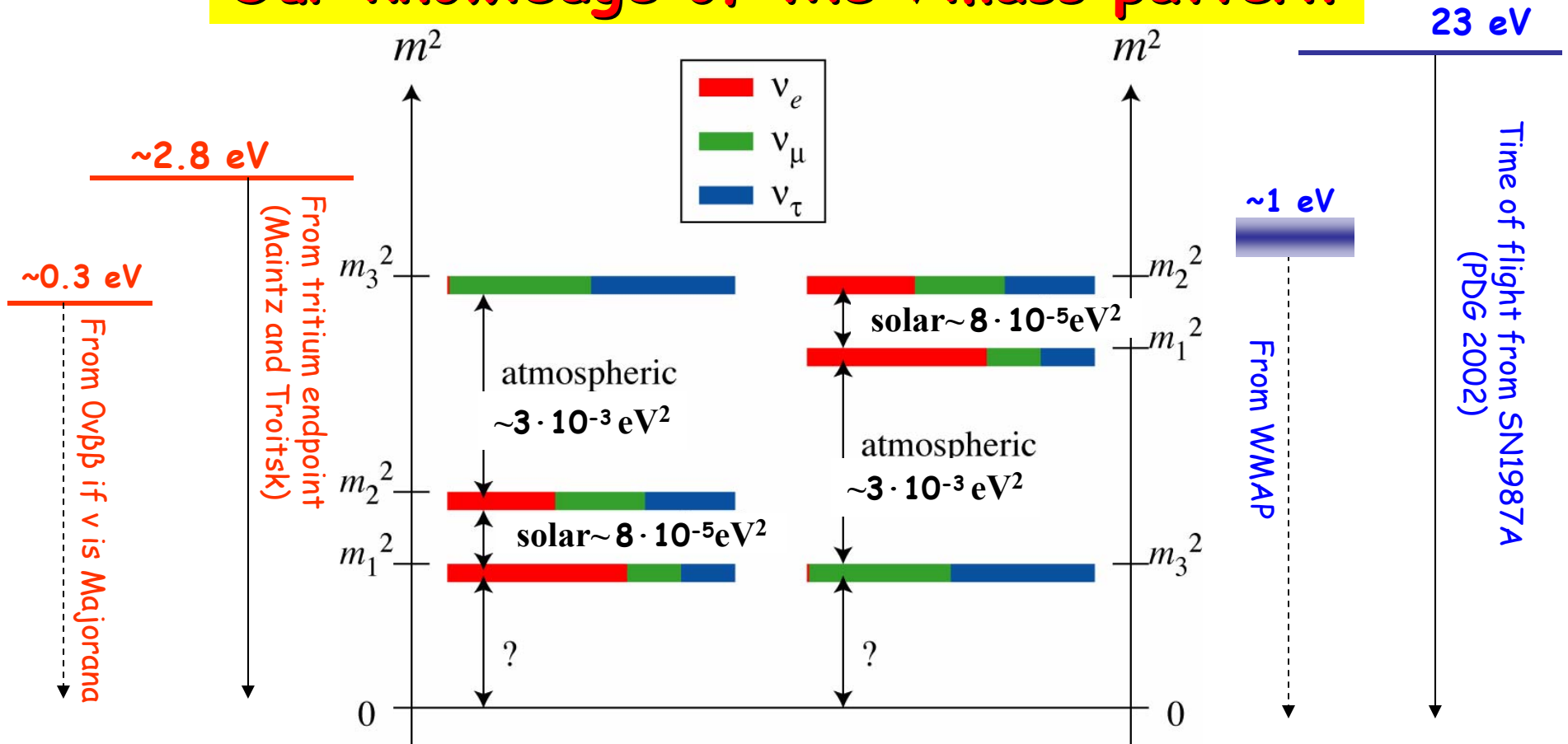
$\theta_{13} < 7^\circ @ 90\%CL$

Solar  $\nu$   
KamLAND  
Future solar  $\nu$  exp.

$\theta_{12} \sim 32^\circ$

Neutrinoless  
double beta  
decay

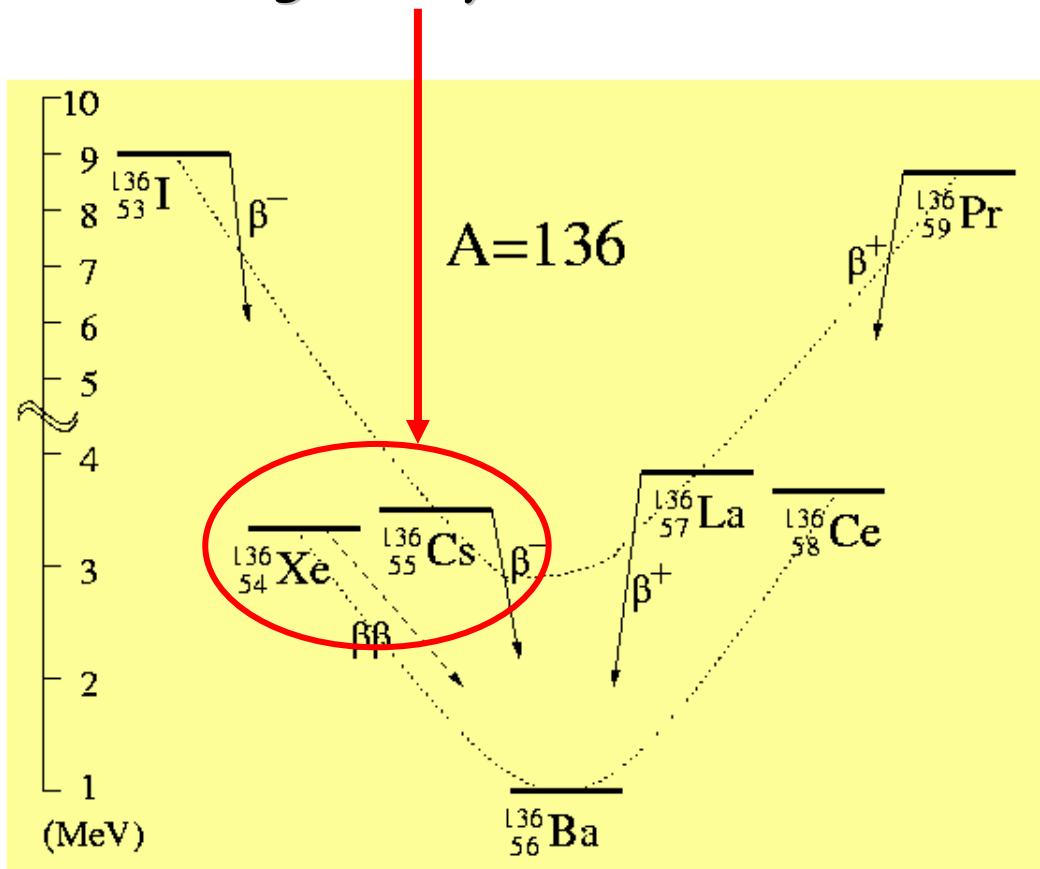
# Our knowledge of the $\nu$ mass pattern



The measurement of the **absolute mass scale**,  $\theta_{13}$  and the choice of hierarchy are the next big challenges in neutrino physics

# Double-beta decay:

*a second-order process  
only detectable if first  
order beta decay is  
energetically forbidden*



## Candidate nuclei with $Q > 2$ MeV

Candidate	Q (MeV)	Abund. (%)
$^{48}\text{Ca} \rightarrow ^{48}\text{Ti}$	4.271	0.187
$^{76}\text{Ge} \rightarrow ^{76}\text{Se}$	2.040	7.8
$^{82}\text{Se} \rightarrow ^{82}\text{Kr}$	2.995	9.2
$^{96}\text{Zr} \rightarrow ^{96}\text{Mo}$	3.350	2.8
$^{100}\text{Mo} \rightarrow ^{100}\text{Ru}$	3.034	9.6
$^{110}\text{Pd} \rightarrow ^{110}\text{Cd}$	2.013	11.8
$^{116}\text{Cd} \rightarrow ^{116}\text{Sn}$	2.802	7.5
$^{124}\text{Sn} \rightarrow ^{124}\text{Te}$	2.228	5.64
$^{130}\text{Te} \rightarrow ^{130}\text{Xe}$	2.533	34.5
$^{136}\text{Xe} \rightarrow ^{136}\text{Ba}$	2.479	8.9
$^{150}\text{Nd} \rightarrow ^{150}\text{Sm}$	3.367	5.6



# There are two varieties of $\beta\beta$ decay

$2\nu$  mode:  
a conventional  
 $2^{\text{nd}}$  order process  
in nuclear physics

$0\nu$  mode: a hypothetical  
process can happen

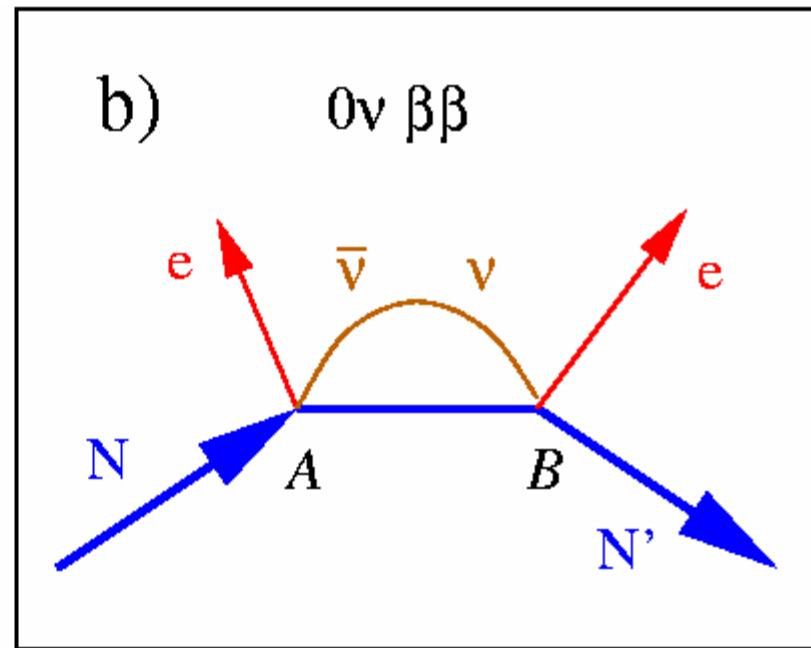
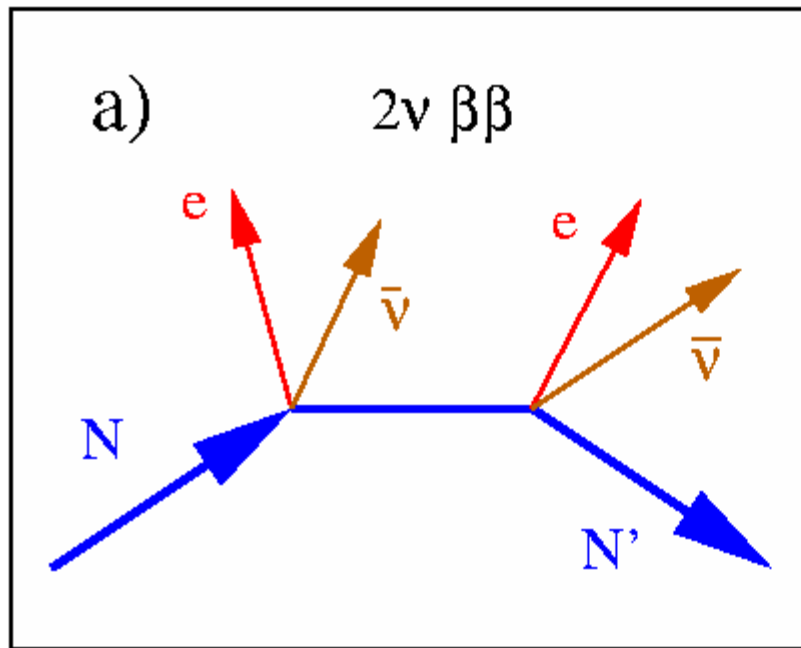
only if:  $M_\nu \neq 0$

$$\nu = \bar{\nu}$$

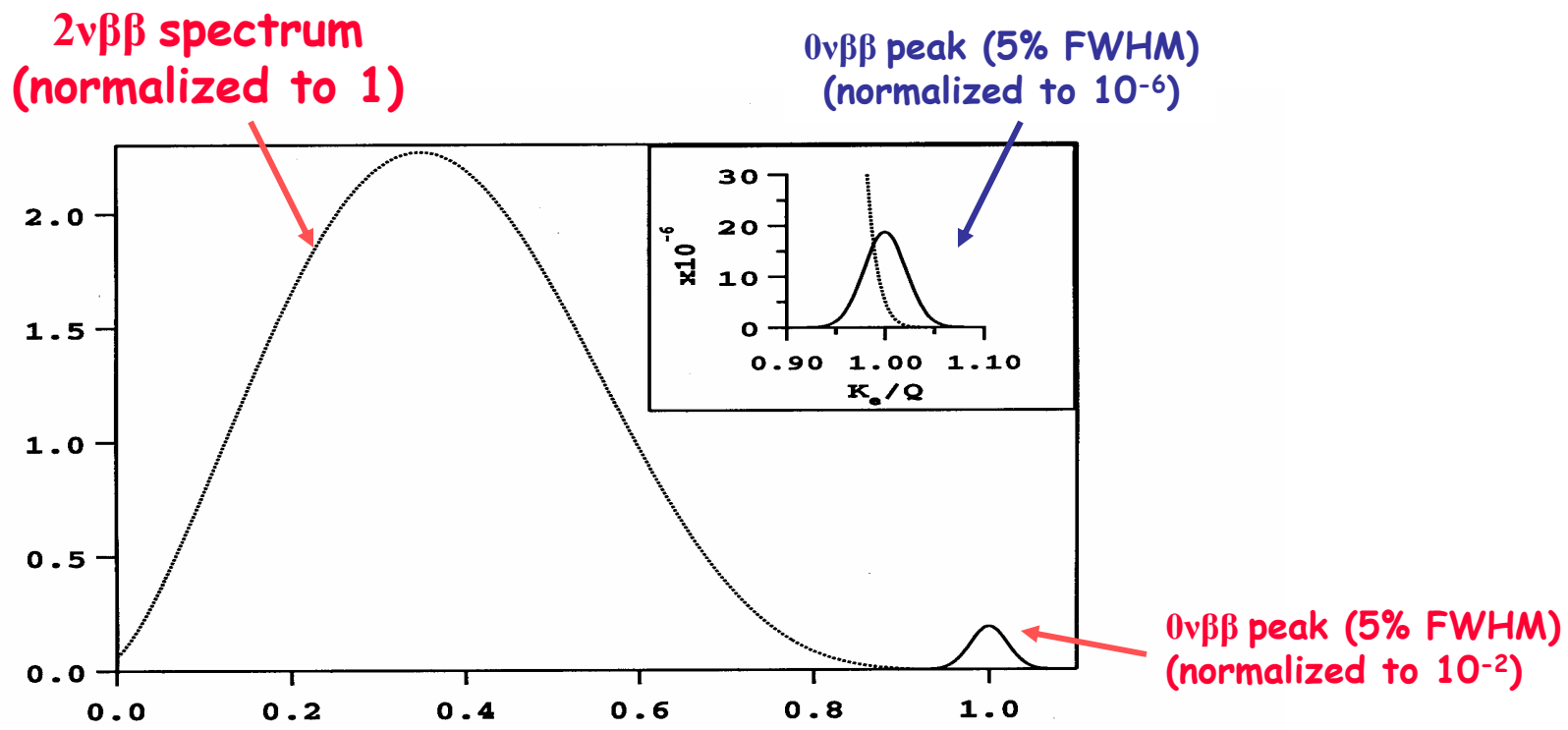
$$|\Delta L|=2$$

$$|\Delta(B-L)|=2$$

Since helicity  
has to "flip"



## Background due to the Standard Model $2\nu\beta\beta$ decay



Summed electron energy in units of the kinematic endpoint ( $Q$ )

The two can be separated in a detector with good energy resolution

## If $0\nu\beta\beta$ is due to light $\nu$ Majorana masses

$$\langle m_\nu \rangle^2 = \left( T_{1/2}^{0\nu\beta\beta} G^{0\nu\beta\beta}(E_0, Z) \left| M_{GT}^{0\nu\beta\beta} - \frac{g_V^2}{g_A^2} M_F^{0\nu\beta\beta} \right|^2 \right)^{-1}$$

$$M_F^{0\nu\beta\beta} \text{ and } M_{GT}^{0\nu\beta\beta}$$

can be calculated within particular nuclear models

$$G^{0\nu\beta\beta}$$

a known phase space factor

$$T_{1/2}^{0\nu\beta\beta}$$

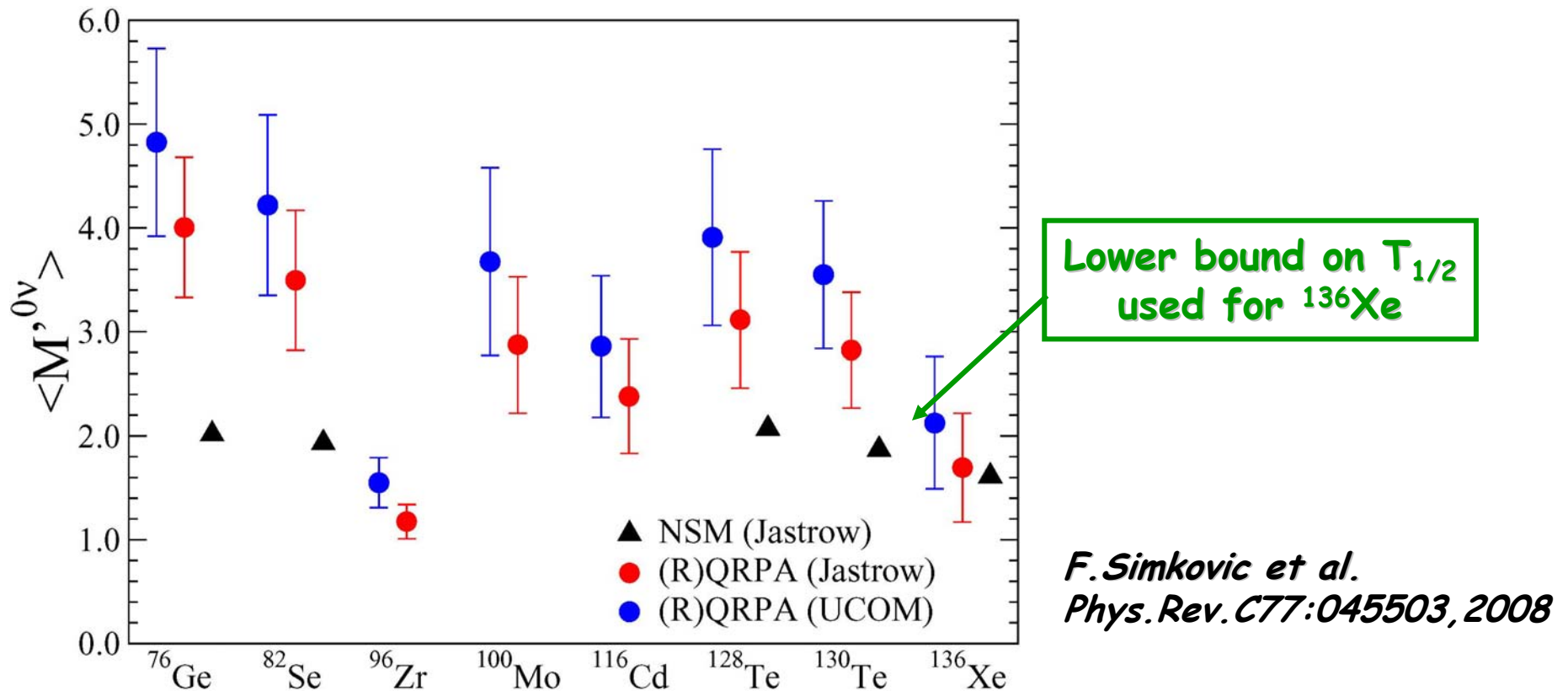
is the quantity to be measured

$$\langle m_\nu \rangle = \sum_{i=1}^3 |U_{e,i}|^2 m_i \varepsilon_i$$

**effective Majorana  $\nu$  mass**  
( $\varepsilon_i = \pm 1$  if CP is conserved)

*Cancellations are possible...*

Much progress made recently in accuracy of nuclear matrix elements.  
 (e.g. was found that main uncertainty in (R)QRPA calculations comes from the single particle space around the Fermi surface.  
 → Can use the measured  $2\nu\beta\beta$   $T_{1/2}$  to make a correction.)



Still, if/once  $0\nu\beta\beta$  decay is discovered, the  $T_{1/2}$  in more than one nucleus will be needed to pin down neutrino masses

## Present Limits for $0\nu$ double beta decay

Candidate nucleus	Detector type	(kg yr)	Present $T_{1/2}^{0\nu\beta\beta}$ (yr)	$\langle m \rangle$ (eV)
$^{48}\text{Ca}$	Ge diode	~47.7	$>1.4 \cdot 10^{22}$ (90%CL)	$<0.35^*$
$^{76}\text{Ge}$			$>1.9 \cdot 10^{25}$ (90%CL)	
$^{82}\text{Se}$			$>1 \cdot 10^{23}$ (90%CL)	
$^{100}\text{Mo}$			$>4.6 \cdot 10^{23}$ (90%CL)	
$^{116}\text{Cd}$			$>1.7 \cdot 10^{23}$ (90%CL)	
$^{128}\text{Te}$	TeO <sub>2</sub> cryo	~12	$>1.1 \cdot 10^{23}$ (90%CL)	$<0.19 - 0.68$
$^{130}\text{Te}$	TeO <sub>2</sub> cryo		$>3 \cdot 10^{24}$ (90%CL)	
$^{136}\text{Xe}$	Xe scint		$>1.2 \cdot 10^{24}$ (90%CL)	
$^{150}\text{Nd}$			$>1.2 \cdot 10^{21}$ (90%CL)	
$^{160}\text{Gd}$			$>1.3 \cdot 10^{21}$ (90%CL)	

*\* But also claim of signal by part of same group*

# There is also a discovery claim (using $^{76}\text{Ge}$ )

## EVIDENCE FOR NEUTRINOLESS DOUBLE BETA DECAY

H.V. KLAPDOR-KLEINGROTHAUS<sup>1,3</sup>,  
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<sup>2</sup>Radiophysical-Research Institute, Nishnii-Nougorod, Russia

<sup>3</sup>Spokesman of the GENIUS and HEIDELBERG-MOSCOW Collaborations,

Part of the  
Heidelberg-Moscow collaboration  
Mod. Phys Lett. A27 (2001) 2409

With the values  $T_{1/2}^{0\nu\beta\beta} = 2.23^{+0.44}_{-0.31}$   $m_\nu^{\text{eff}} = 0.32 \pm 0.03$

...but this is a controversial matter (see details in)

*C.A.Aalseth Mod. Phys. Lett. A17 (2002) 1475*

*F.Feruglio et al. Nucl.Phys. B637 (2002) 345*

*Addendum-ibid. B659 (2003) 359*

*Yu.Zdesenko et al. Phys.Lett. B 546 (2002) 206*

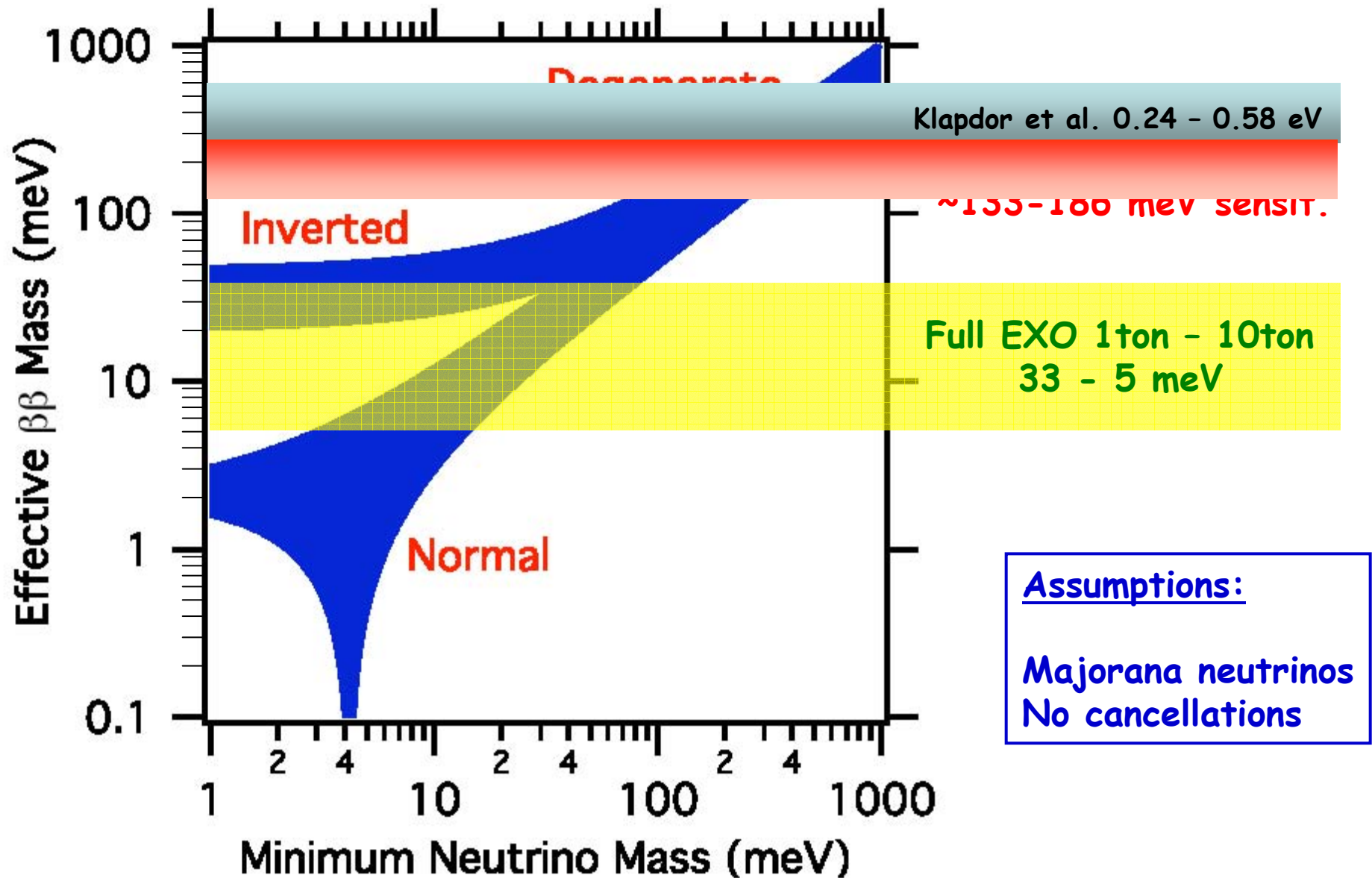
*H.L.Harney Mod.Phys.Lett. A16 (2001) 2409*

*A.M.Bakalyarov et al. hep-ex/0309016*

*H.V.Klapdor-Kleingrouthaus et al. Phys. Lett. B 586 (2004) 198*

*H.V.Klapdor-Kleingrouthaus et al. Mod. Phys. Lett. 21 (2006) 1547*

For the first time there is a clear opportunity to make an important discovery if one pushes the  $\langle m \rangle$  sensitivity to the 0.01 - 1 eV region



Plot from Avignone, Elliott, Engel arXiv:0708.1033 (2007)

In the last 10 years there has been a transition

- 1) *From a few kg detectors to 100s or 1000s kg detectors*  
→ *Think big: qualitative transition from cottage industry to large experiments*
- 2) *From "random shooting" to the knowledge that at least the inverted hierarchy will be tested*

Discovering  $0\nu\beta\beta$  decay:

- *Discovery of the neutrino mass scale*
- *Discovery of Majorana particles*
- *Discovery of lepton number violation*



To reach  $\langle m_\nu \rangle \sim 10$  meV very large fiducial mass (tons)  
(except for Te) need massive isotopic enrichment  
Need to reduce and control backgrounds in qualitatively new ways  
these are the lowest background experiment ever built

For no bkgnd  $\langle m_\nu \rangle \propto 1 / \sqrt{T_{1/2}^{0\nu\beta\beta}} \propto 1 / \sqrt{Nt}$

Scaling with bkgd  
goes like  $Nt$   $\langle m_\nu \rangle \propto 1 / \sqrt{T_{1/2}^{0\nu\beta\beta}} \propto 1 / (Nt)^{1/4}$

In addition a multi-parameter experiment, if feasible,  
would provide information for cross checks with  
more than one single variable, if a discovery is made.

**$\beta\beta$  decay experiments are at the leading edge of "low background" techniques**

- Final state ID: 1) "Geochemical": search for an abnormal abundance of  $(A, Z+2)$  in a material containing  $(A, Z)$   
2) "Radiochemical": store in a mine some material  $(A, Z)$  and after some time try to find  $(A, Z+2)$  in it
  - + Very specific signature
  - + Large live times (particularly for 1)
  - + Large masses
  - Possible only for a few isotopes (in the case of 1)
  - No distinction between  $0\nu, 2\nu$  or other modes
- "Real time": ionization or scintillation is detected in the decay
  - a) "Homogeneous": source=detector
  - b) "Heterogeneous": source $\neq$ detector
    - + Energy/some tracking available (can distinguish modes)
    - + In principle universal (b)
    - Many  $\gamma$  backgrounds can fake signature
    - Exposure is limited by human patience

# Xe is ideal for a large experiment

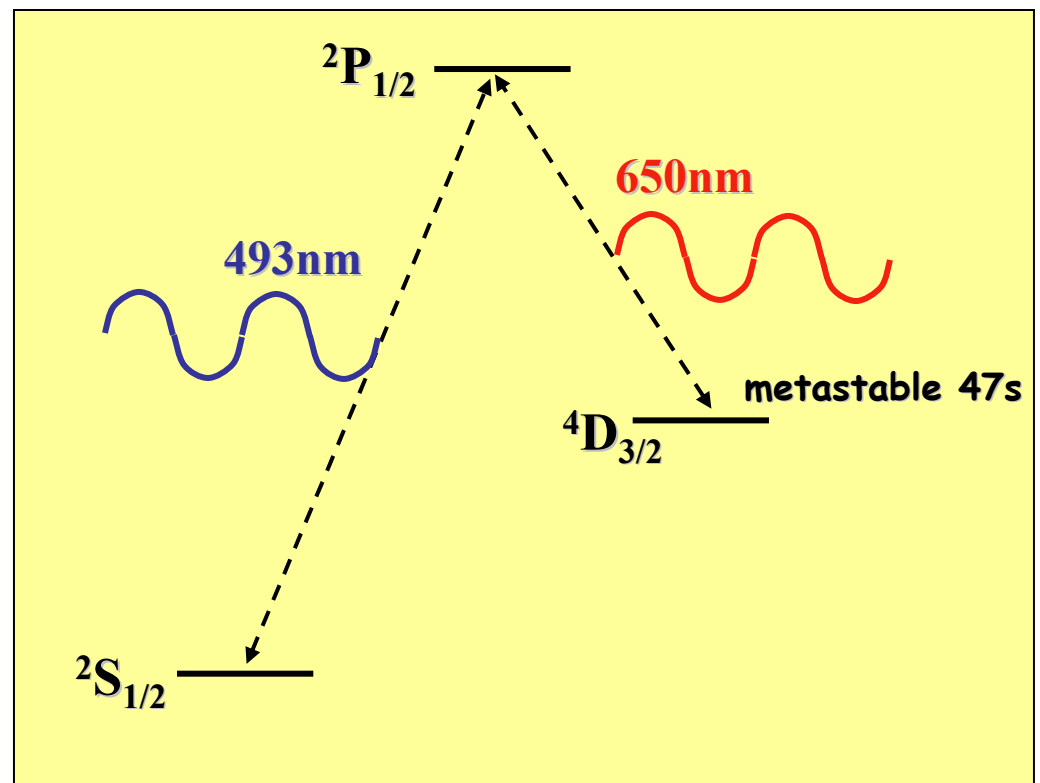
- No need to grow crystals
- Can be re-purified during the experiment
- No long lived Xe isotopes to activate
- Can be easily transferred from one detector to another if new technologies become available
- Noble gas: easy(er) to purify
- $^{136}\text{Xe}$  enrichment easier and safer:
  - noble gas (no chemistry involved)
  - centrifuge feed rate in gram/s, all mass useful
  - centrifuge efficiency  $\sim \Delta m$ . For Xe 4.7 amu
- $^{129}\text{Xe}$  is a hyperpolarizable nucleus, under study for NMR tomography... a joint enrichment program ?

Xe offers a qualitatively new tool against background:  
 $^{136}\text{Xe} \rightarrow ^{136}\text{Ba}^{++} e^- e^-$  final state can be identified  
using optical spectroscopy (M.Moe PRC44 (1991) 931)

Ba<sup>+</sup> system best studied  
(Neuhauser, Hohenstatt,  
Toshek, Dehmelt 1980)  
Very specific signature  
"shelving"

Single ions can be detected  
from a photon rate of  $10^7/\text{s}$

- Important additional constraint
- Drastic background reduction



The Ba-tagging, added to a high resolution Xe imaging detector provides the tools to develop a background-free next-generation  $\beta\beta$  experiment

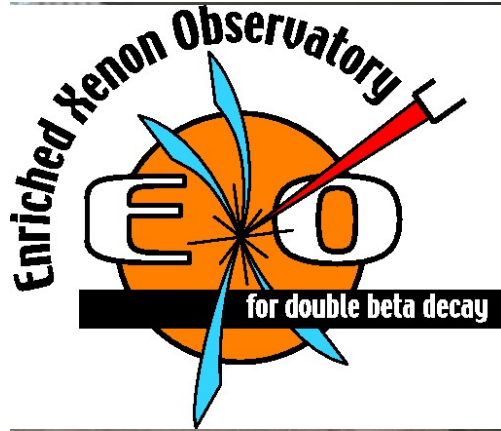
Assume an "asymptotic" fiducial mass of 10 tons of  $^{136}\text{Xe}$  at 80%

A somewhat natural scale:

- World production of Xe is  $\sim 40$  ton/yr
- Detector size
- $2 \cdot 10^3$  size increase: good match to the  $10^{-2}$  eV mass region

Mainly going in light bulbs, plasma displays and satellite propulsion







## The EXO Collaboration

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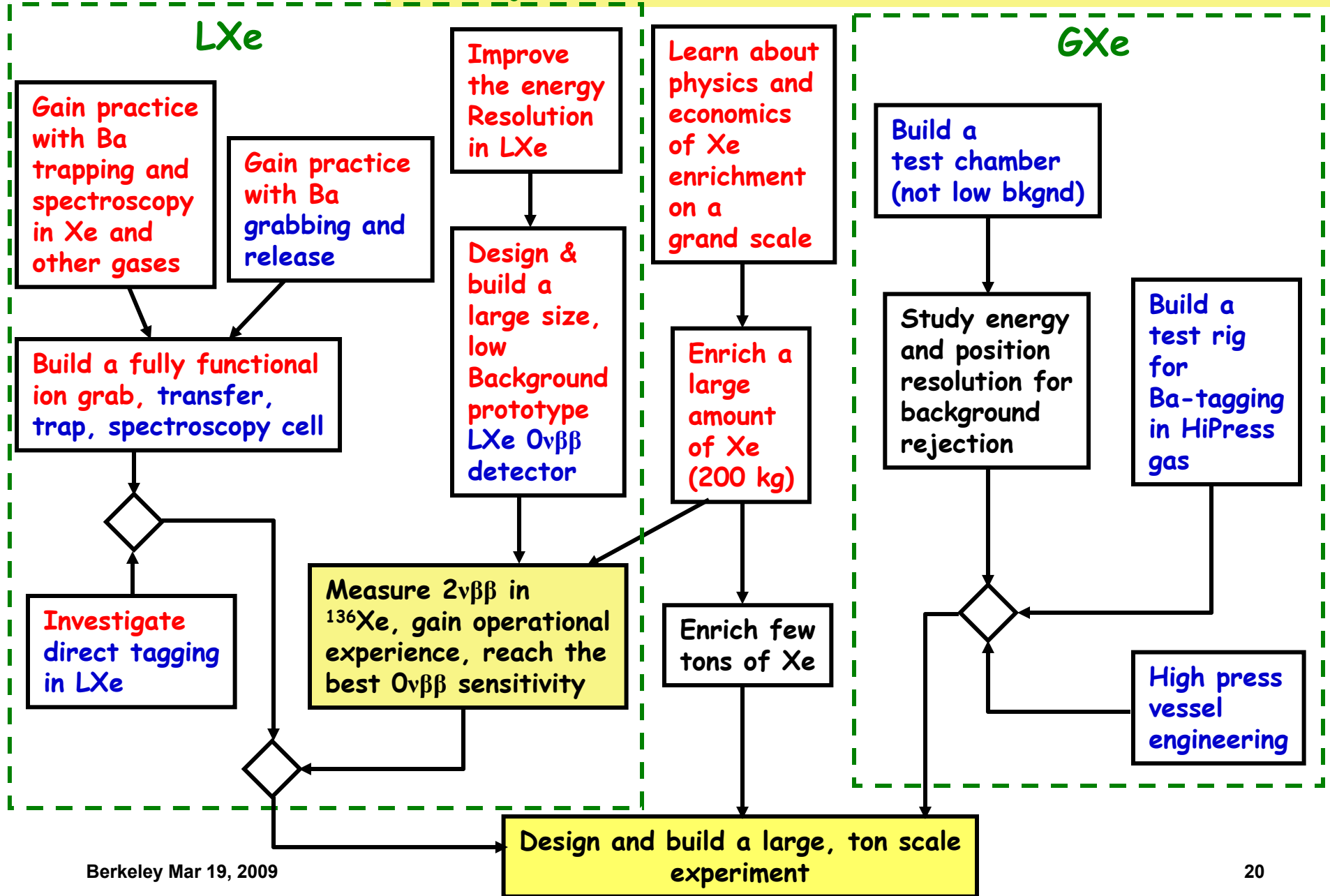
L. Bartoszek, R. Cooper, R. DeVoe, M. Dolinski, B. Flatt, G. Gratta, M. Green, F. LePort,

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# The roadmap to the background-free discovery of Majorana neutrinos and the neutrino mass scale

Done    In progress  
To do





# EXO-200kg Majorana mass sensitivity

## Assumptions:

- 1) 200kg of Xe enriched to 80% in  $^{136}$
- 2)  $\sigma(E)/E = 1.4\%$  obtained in EXO R&D, Conti et al Phys Rev B 68 (2003) 054201
- 3) Low but finite radioactive background:  
20 events/year in the  $\pm 2\sigma$  interval centered around the 2.481MeV endpoint
- 4) Negligible background from  $2\nu\beta\beta$  ( $T_{1/2} > 1 \cdot 10^{22}$ yr R.Bernabei et al. measurement)

Case	Mass (ton)	Eff. (%)	Run Time (yr)	$\sigma_E/E$ @ 2.5MeV (%)	Radioactive Background (events)	$T_{1/2}^{0\nu}$ (yr, 90%CL)	Majorana mass (eV)	
							QRPA	NSM
EXO-200	0.2	70	2	1.6*	40	$6.4 \cdot 10^{25}$	0.133†	0.186*

## *What if Klapdor's observation is correct ?*

Central value  $T_{1/2}(\text{Ge}) = 1.2^{+3}_{-0.5} \cdot 10^{25}$ , ( $\pm 3\sigma$ )  
 (Phys. Lett. B 586 (2004) 198-212)  
 consistently use Rodin's matrix elements for both Ge and Xe)

In 200kg EXO, 2yr:

- Worst case (QRPA, upper limit) 15 events on top of 40 events bkgd  $\rightarrow 2\sigma$
- Best case (NSM, lower limit) 162 events on top of 40 bkgd  $\rightarrow 11\sigma$

# EXO-200kg Majorana mass sensitivity

## Assumptions:

- 1) 200kg of Xe enriched to 80% in 136
- 2)  $\sigma(E)/E = 1.4\%$  obtained in EXO R&D, Conti et al Phys Rev B 68 (2003) 054201
- 3) Low but finite radioactive background:  
20 events/year in the  $\pm 2\sigma$  interval centered around the 2.481MeV endpoint
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Case	Mass (ton)	Eff. (%)	Run Time (yr)	$\sigma_E/E$ @ 2.5MeV (%)	Radioactive Background (events)	$T_{1/2}^{0\nu}$ (yr, 90%CL)	Majorana mass (eV)	
							QRPA	NSM
EXO-200	0.2	70	2	1.6*	40	$6.4 \cdot 10^{25}$	0.13†	0.19*

## *What if Klapdor's observation is correct ?*

Central value  $T_{1/2} (\text{Ge}) = 2.23^{+0.44}_{-0.31} \cdot 10^{25}, (\pm 3\sigma)$

In EXO-200, 2yr:

• Worst case (QRPA, upper limit) 46 events on top of 40 events bkgd  $\rightarrow 5.0 \sigma$

• Best case (NSM, lower limit) 170 events on top of 40 bkgd  $\rightarrow 11.7 \sigma$

# EXO neutrino effective mass sensitivity

## Assumptions:

- 1) 80% enrichment in  $^{136}\text{Xe}$
- 2) Intrinsic low background + Ba tagging eliminate all radioactive background
- 3) Energy res only used to separate the  $0\nu$  from  $2\nu$  modes:  
Select  $0\nu$  events in a  $\pm 2\sigma$  interval centered around the 2.481 MeV endpoint
- 4) Use for  $2\nu\beta\beta$   $T_{1/2} > 1 \cdot 10^{22}\text{yr}$  (Bernabei et al. measurement)

Case	Mass (ton)	Eff. (%)	Run Time (yr)	$\sigma_E/E$ @ 2.5 MeV (%)	$2\nu\beta\beta$ Background (events)	$T_{1/2}^{0\nu}$ (yr, 90%CL)	Majorana mass (meV)	
							QRPA <sup>‡</sup>	NSM <sup>#</sup>
Conservative	1	70	5	1.6*	0.5 (use 1)	$2 \cdot 10^{27}$	24	33
Aggressive	10	70	10	1 <sup>†</sup>	0.7 (use 1)	$4.1 \cdot 10^{28}$	5.3	7.3

\*  $\sigma(E)/E = 1.4\%$  obtained in EXO R&D, Conti et al Phys Rev B 68 (2003) 054201

<sup>†</sup>  $\sigma(E)/E = 1.0\%$  considered as an aggressive but realistic guess with large light collection area

<sup>‡</sup> Rodin, et. al., Nucl. Phys. A 793 (2007) 213-215

<sup>#</sup> Courier, et. al., arXiv:0709.2137v1

## Status of $2\nu$ mode in $^{136}\text{Xe}$

$2\nu\beta\beta$  decay has never been observed in  $^{136}\text{Xe}$ .  
Some of the lower limits on its half life are close to (and in one case below) the theoretical expectation.

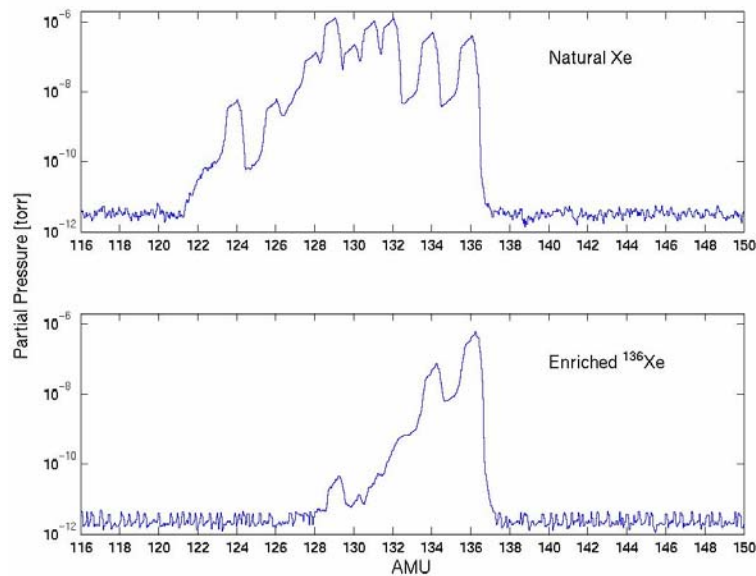
	$T_{1/2}$ (yr)	evts/year in the 200kg prototype (no efficiency applied)
<b>Experimental limit</b>		
Leuscher et al	$>3.6 \cdot 10^{20}$	$<1.3 \text{ M}$
Gavriljuk et al	$>8.1 \cdot 10^{20}$	$<0.6 \text{ M}$
Bernabei et al	$>1.0 \cdot 10^{22}$	$<48 \text{ k}$
<b>Theoretical prediction</b>		
QRPA (Staudt et al) [ $T_{1/2}^{\text{max}}$ ]	$=2.1 \cdot 10^{22}$	$=23 \text{ k}$
QRPA (Vogel et al)	$=8.4 \cdot 10^{20}$	$=0.58 \text{ M}$
NSM (Caurier et al)	$(=2.1 \cdot 10^{21})$	$(=0.23 \text{ M})$

*EXO-200 should definitely resolve this issue*

# EXO-200

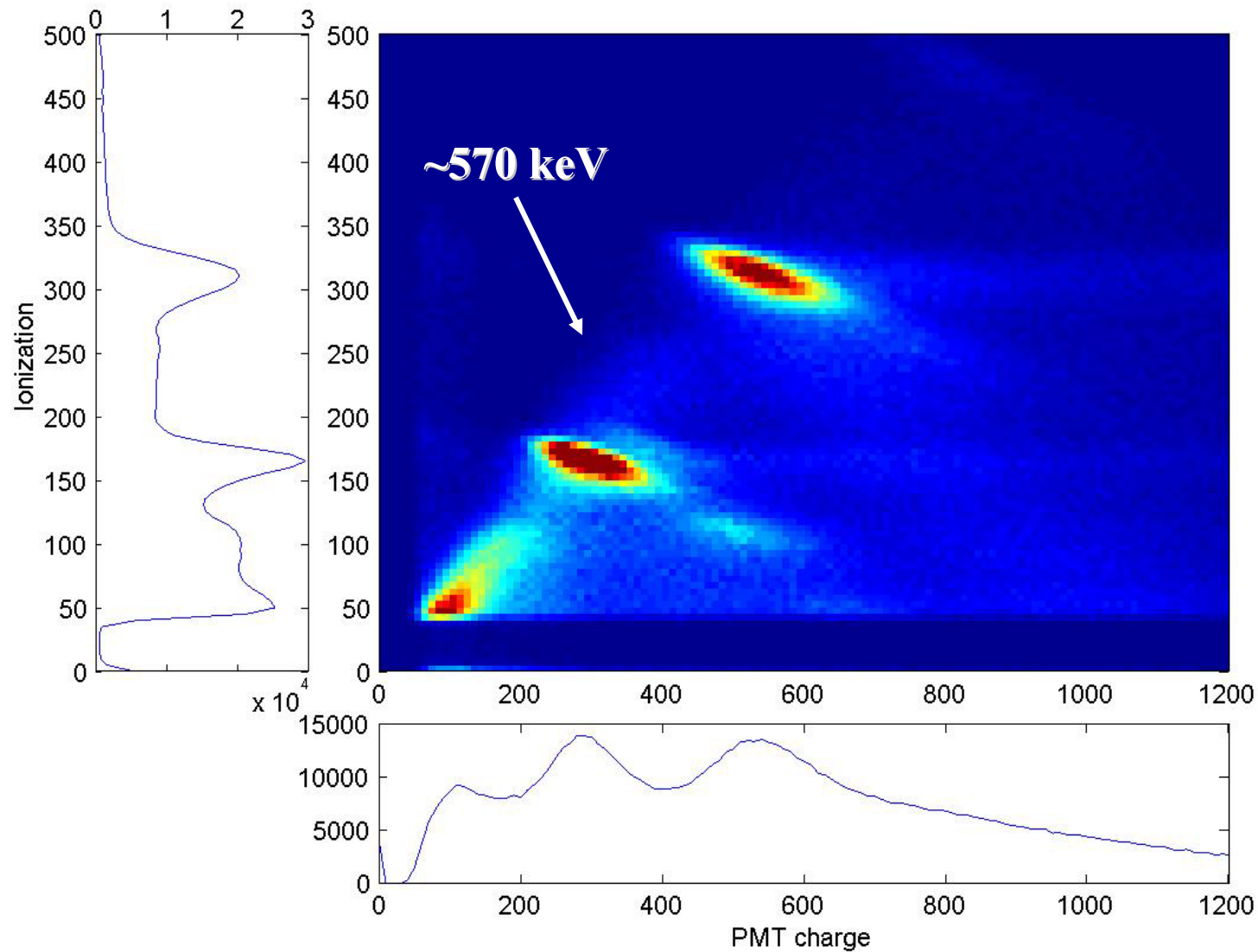


# 200 kg $^{136}\text{Xe}$ test production completed in spring '03 (80% enrichment)

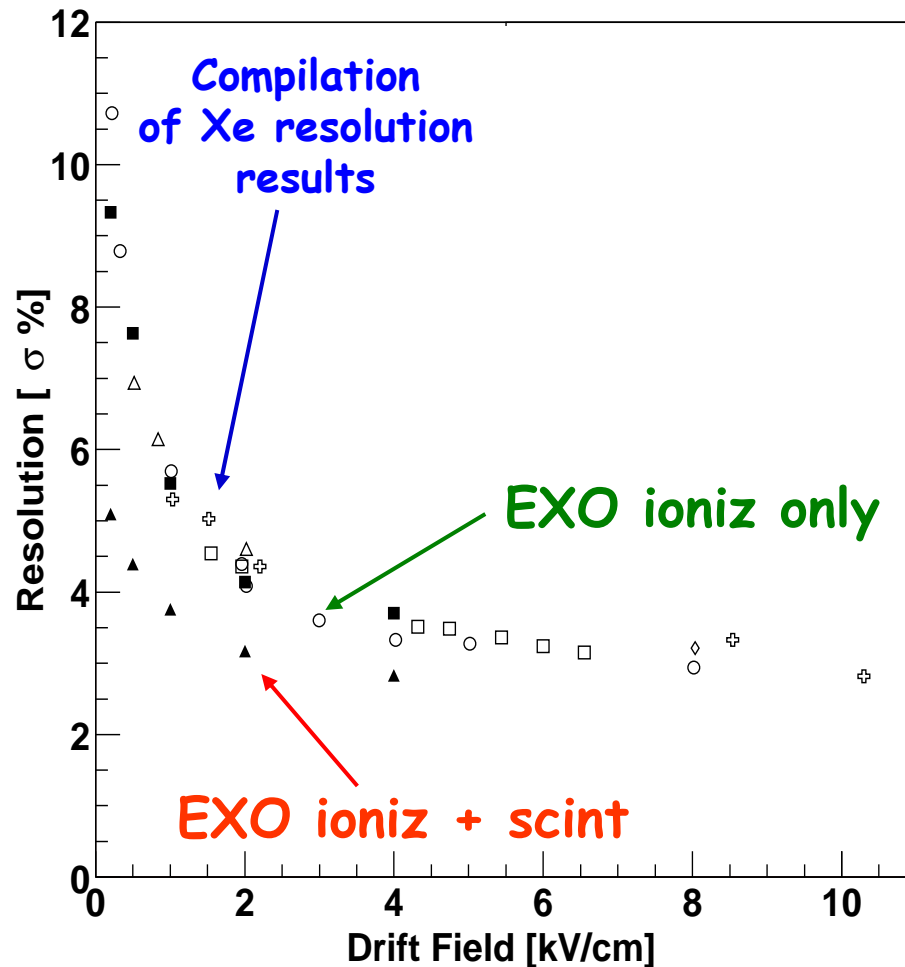


- Largest highly enriched stockpile not related to nuclear industry
- Largest sample of separated  $\beta\beta$  isotope (by ~factor of 10)

**EXO R&D showed the way to improved energy resolution in LXe: Use (anti)correlations between ionization and scintillation signals**



# Anti-correlated ionization and scintillation improves the energy resolution in LXe



Ionization alone:

$$\sigma(E)/E = 3.8\% \text{ @ } 570 \text{ keV}$$

or 1.8% @  $Q_{\beta\beta}$

Ionization & Scintillation:

$$\sigma(E)/E = 3.0\% \text{ @ } 570 \text{ keV}$$

or 1.4% @  $Q_{\beta\beta}$

(a factor of 2 better than the Gotthard TPC)

*E. Conti et al.*

*Phys. Rev. B: 68 (2003) 054201*

and by now other groups have used this

[e.g. *E. Aprile et al. PRB 76 (2007) 014115*]

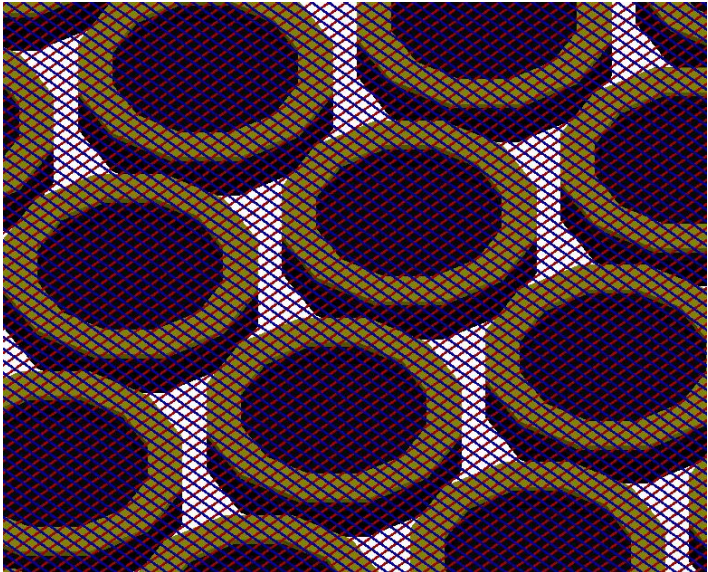
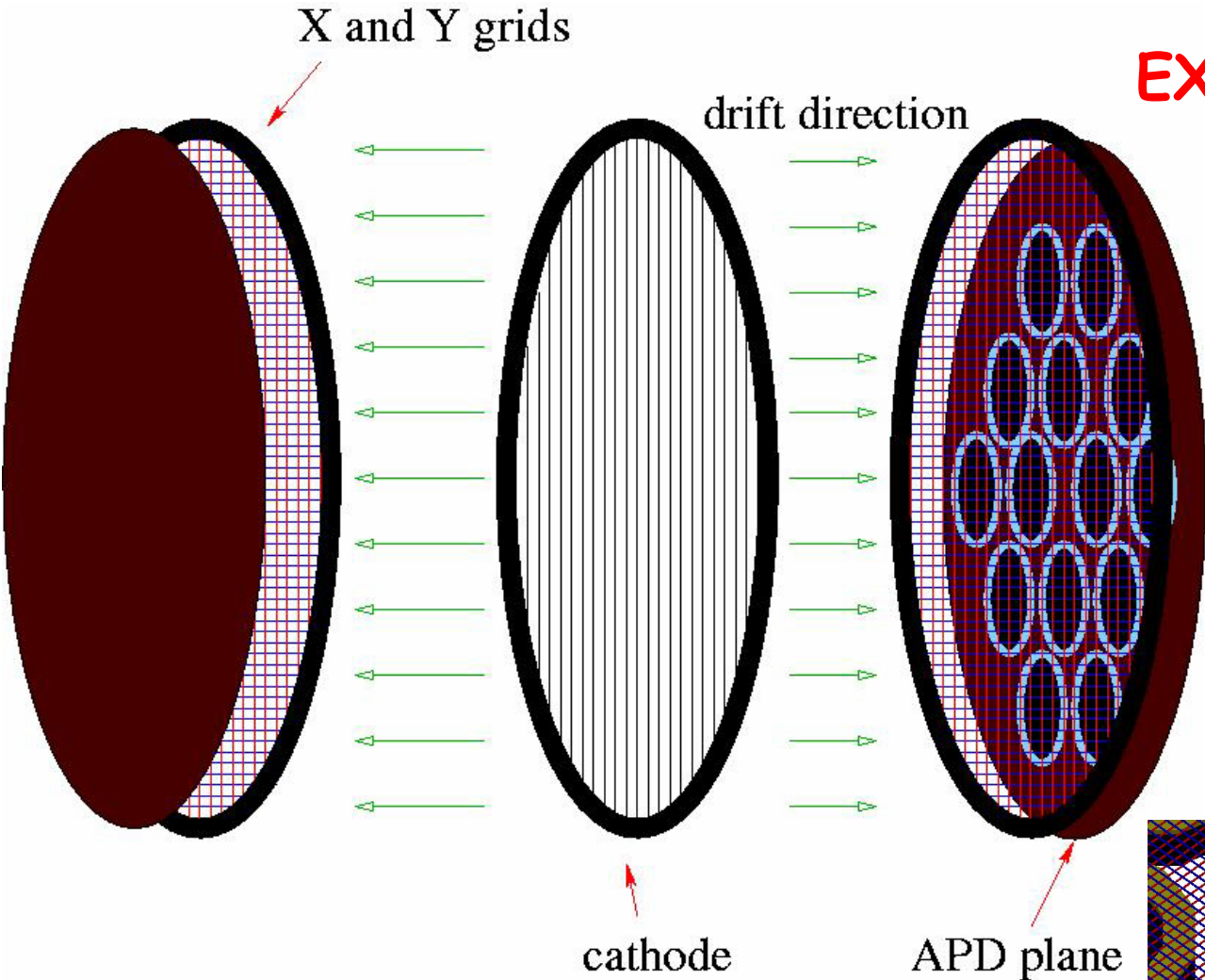
EXO-200 will collect 3-4 times as much scintillation... further improvement possible



# EXO-200 TPC basics

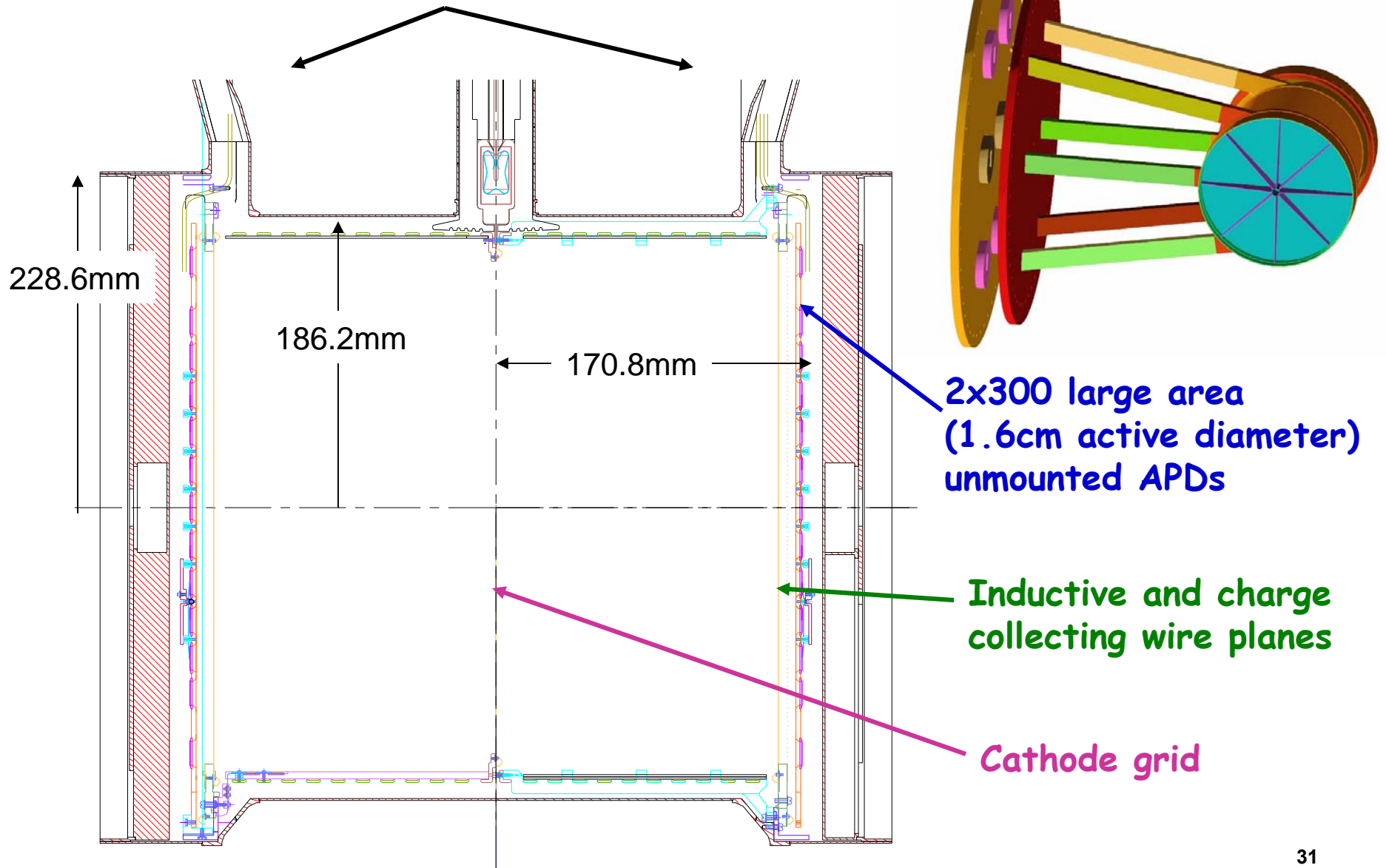
- The detector measures both the ionization electrons and the scintillation light to get best energy resolution.
  - In addition, the position of the decay is measured to get spatial distributions and (for later) the position of the Ba ion.
  - Info on event topology also important for background separation.
- The detector is a cylinder of ~40cm ID by ~40cm inner length.
- The cylinder is split by a cathode plane at the center so there are two symmetric drift regions. The cathode runs at negative HV.
  - Max HV is ~ 70kV (~3.5 kV/cm drift). Energy resolution improves with drift field, but there are arguments that separation of 1 vs 2 primary electrons might be better at lower fields.
    - field optimization is an important mission of EXO-200
- Readout "style":
  - Crossed wires, 100 $\mu$ m wires, 3mm pitch, ganged in groups of 3
  - 48ch x, 48ch y, total 96 ch per 1/2 detector
  - (Pad readout rejected because of high channel count)

# EXO-200 TPC basics



# EXO chamber "hugs" the fiducial volume very closely!

## Mechanical supports and cable/Xe conduits







Low activity copper traveled from Germany to...

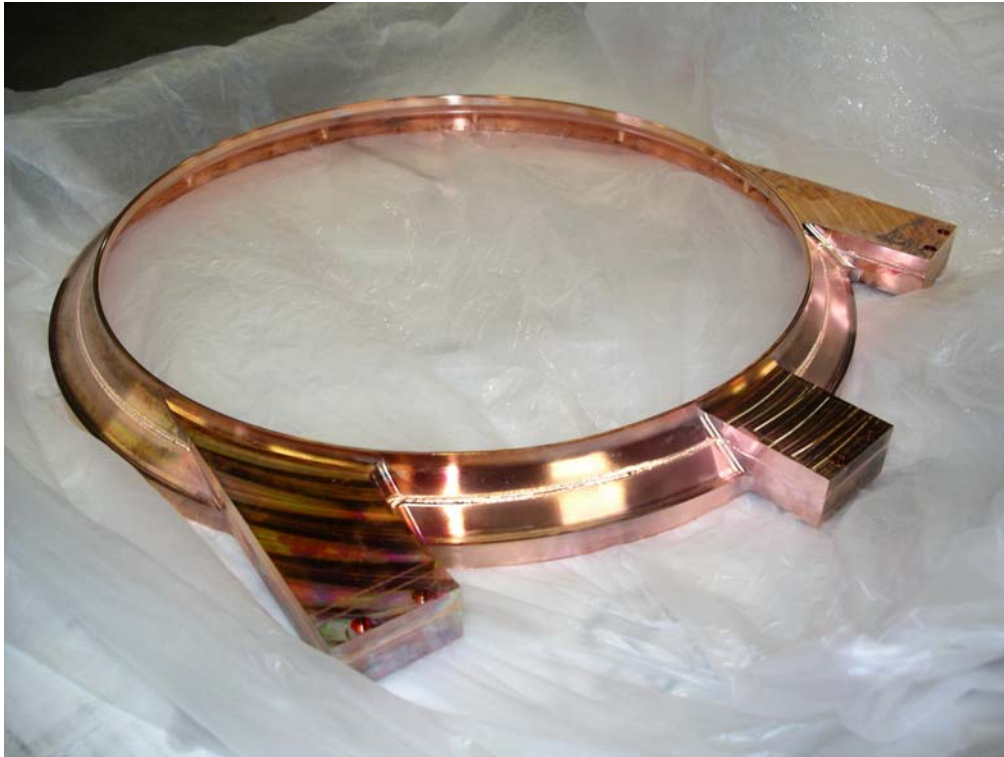
...California in a shielded container



Berkeley Mar 19, 2009

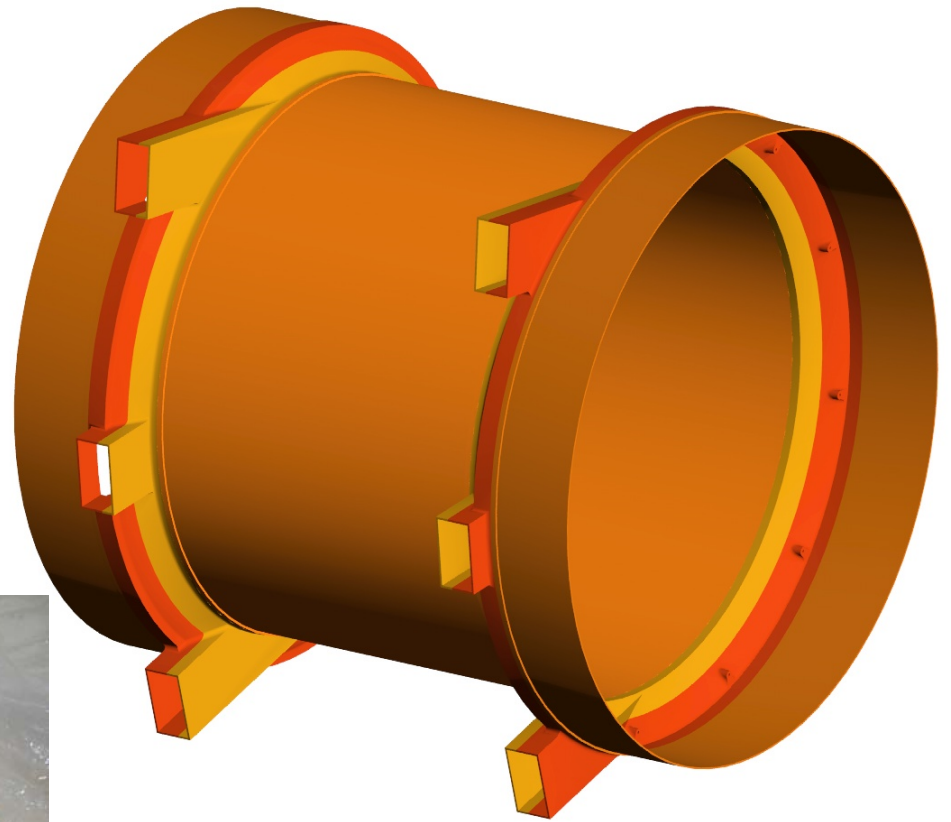
# Ultra-low activity Cu vessel

- Very light (~1.5mm thin, ~15kg) to minimize materials



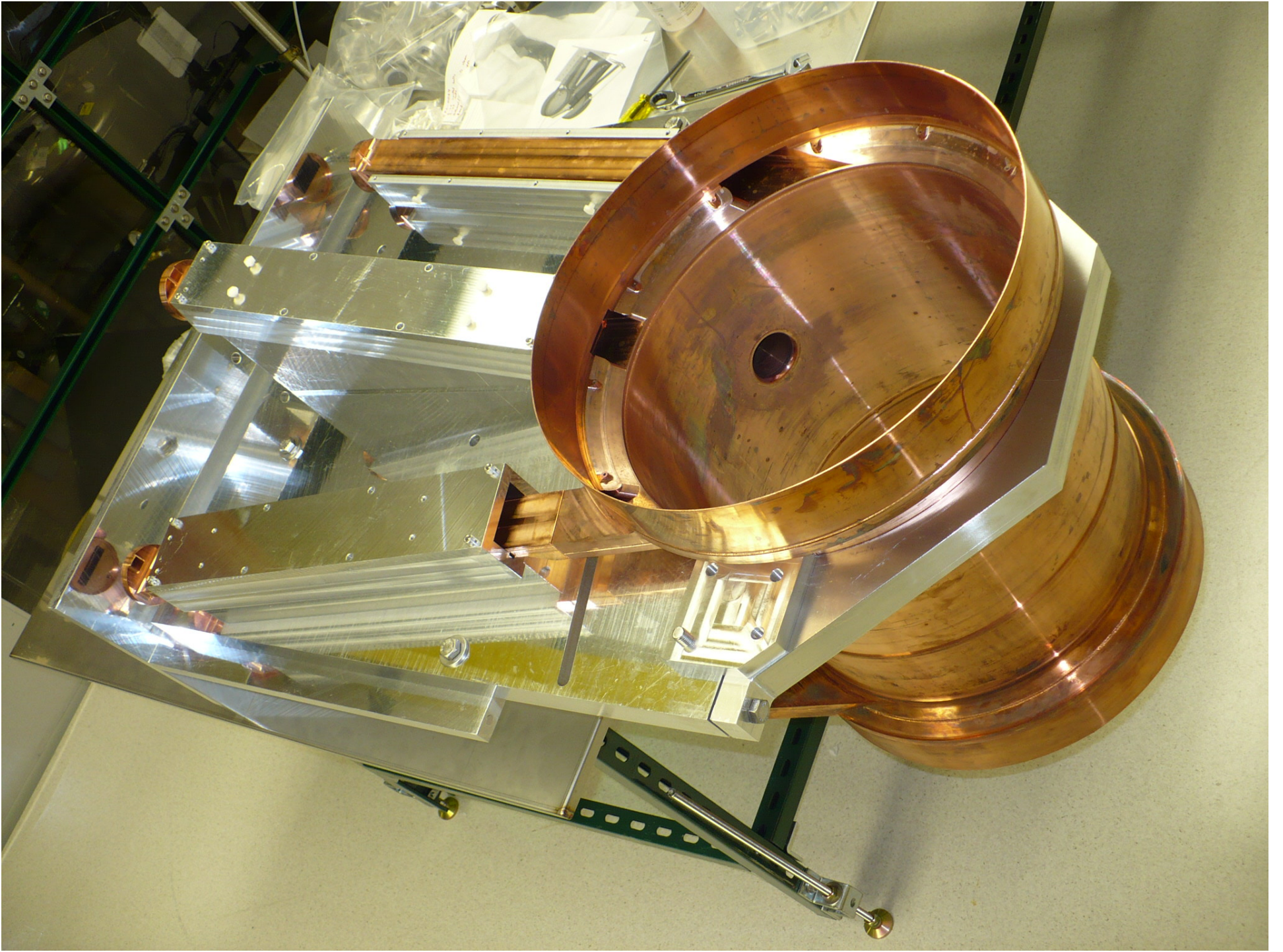
Berkeley Mar 19, 2009

Giorgio Gratta, LLNL



- Different parts e-beam welded together (San Leandro)
- Field TIG weld(s) to seal the vessel after assembly (TIG technology tested for radioactivity)
- All machining done by in a shop under a shallow cosmic ray shielding



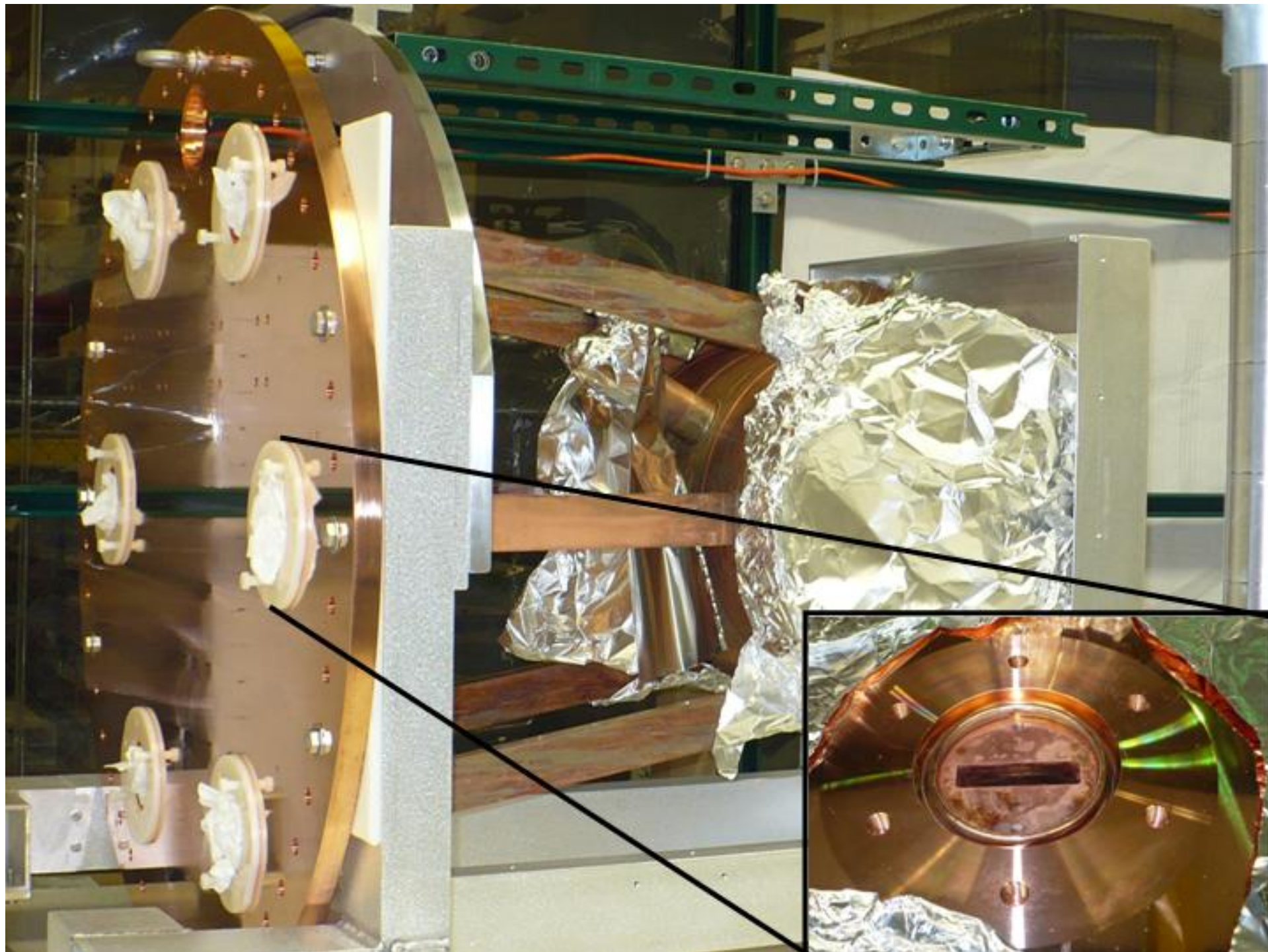






Berkeley Mar 19,





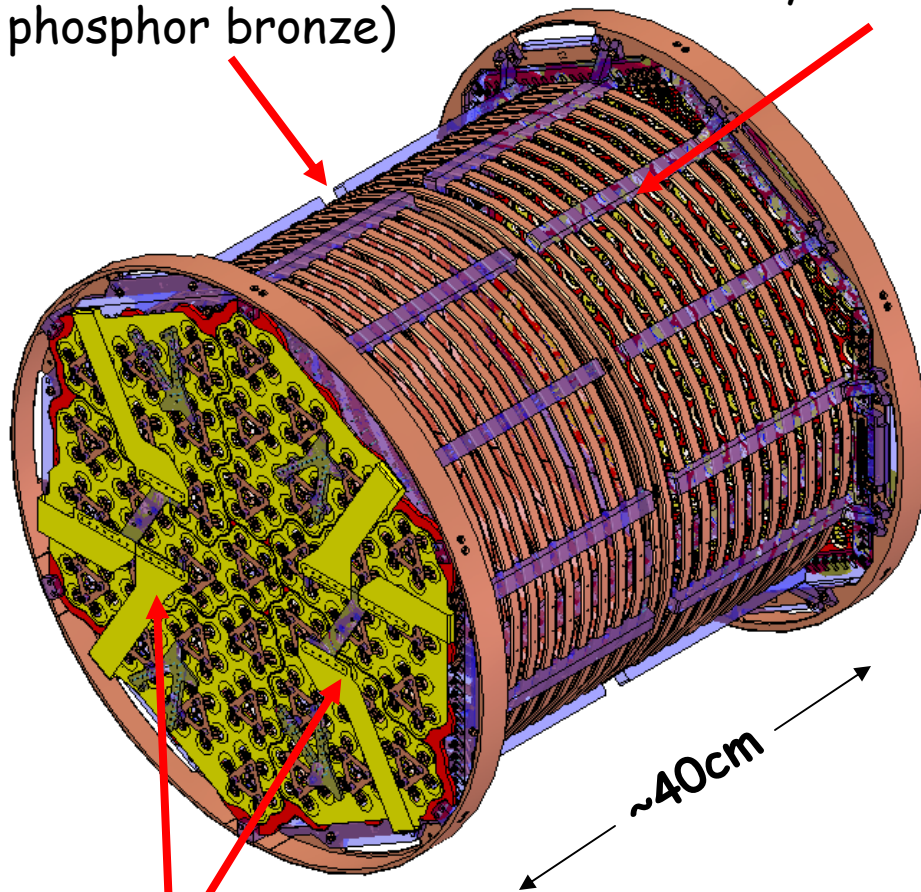


# EXO-200 LXe TPC field cage & readout planes

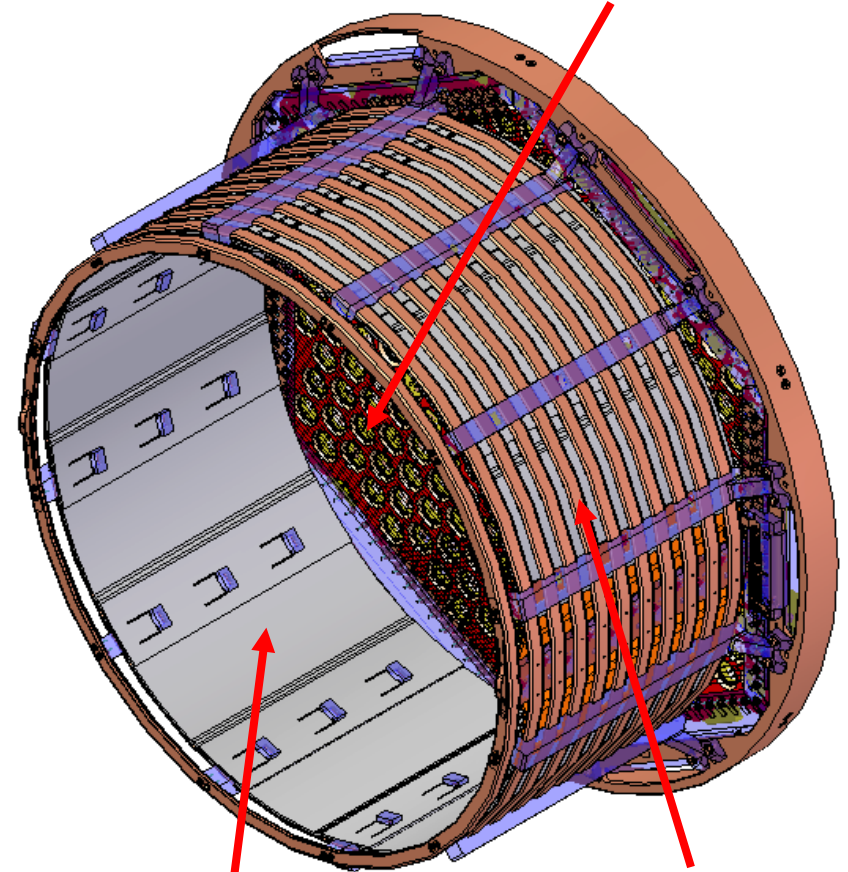
Central HV plane  
(photo-etched  
phosphor bronze)

acrylic supports

APD plane (copper) and  
grid plane (photo-etched  
phosphor bronze)



flex cables on back of APD plane

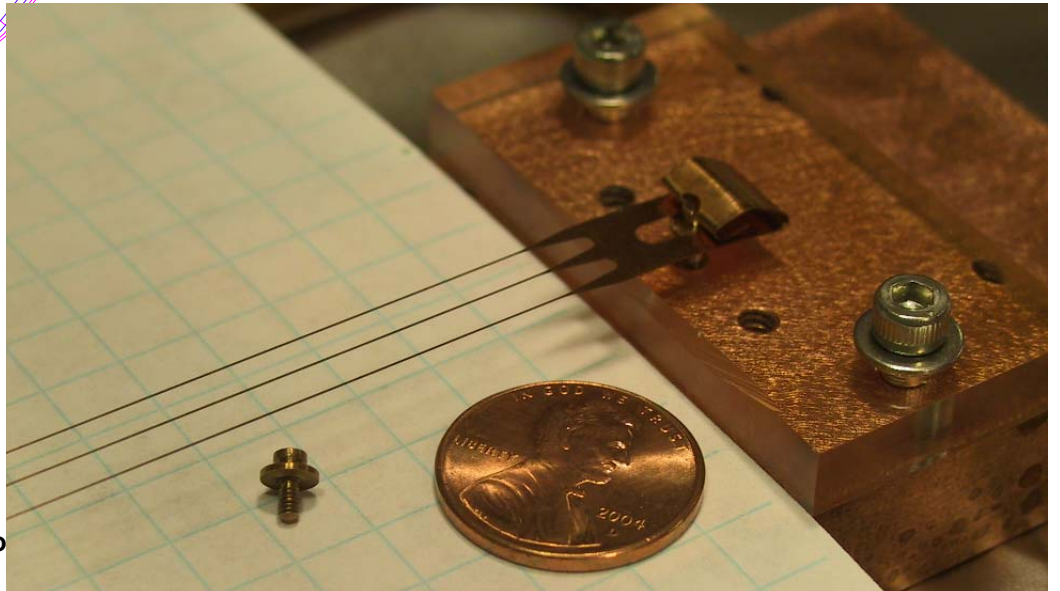
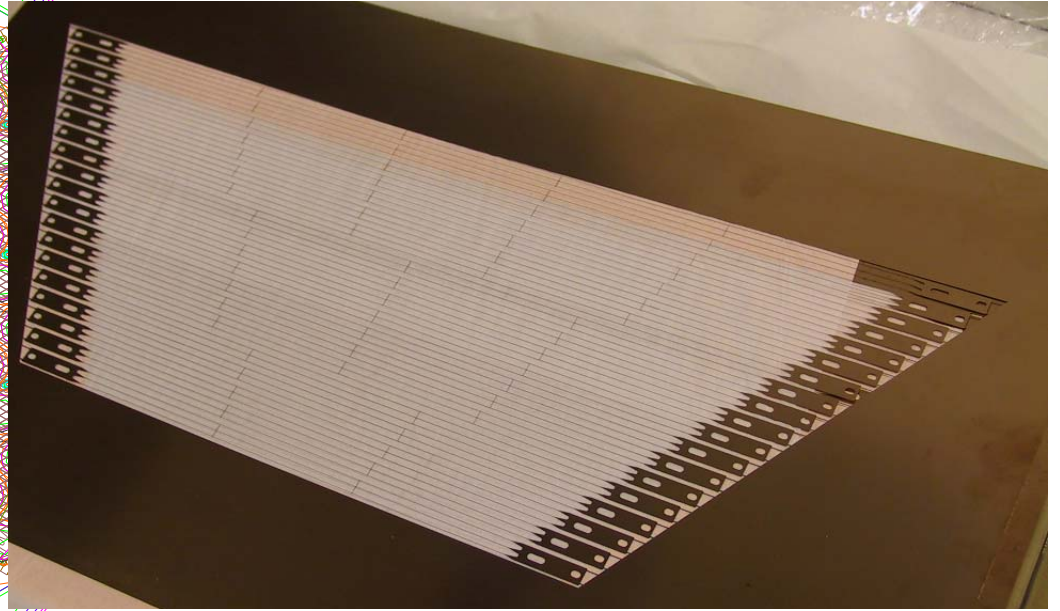
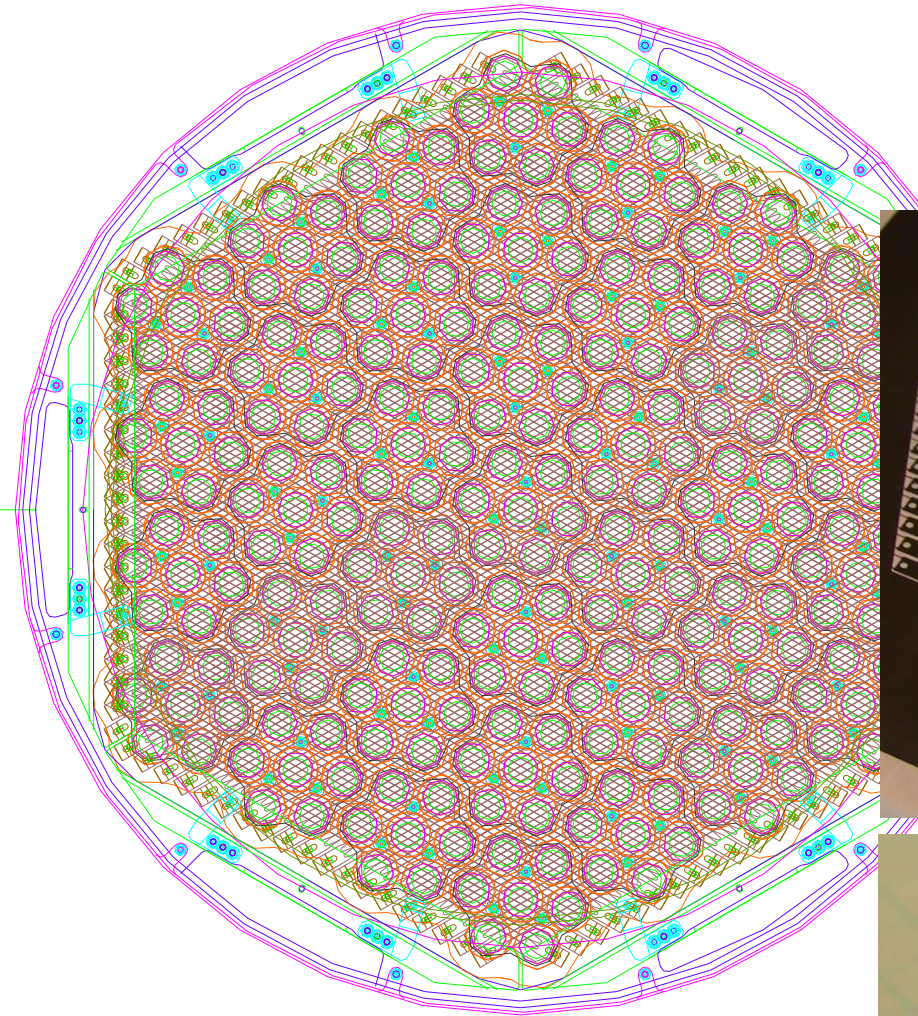


teflon light  
reflectors

field shaping  
rings (copper)



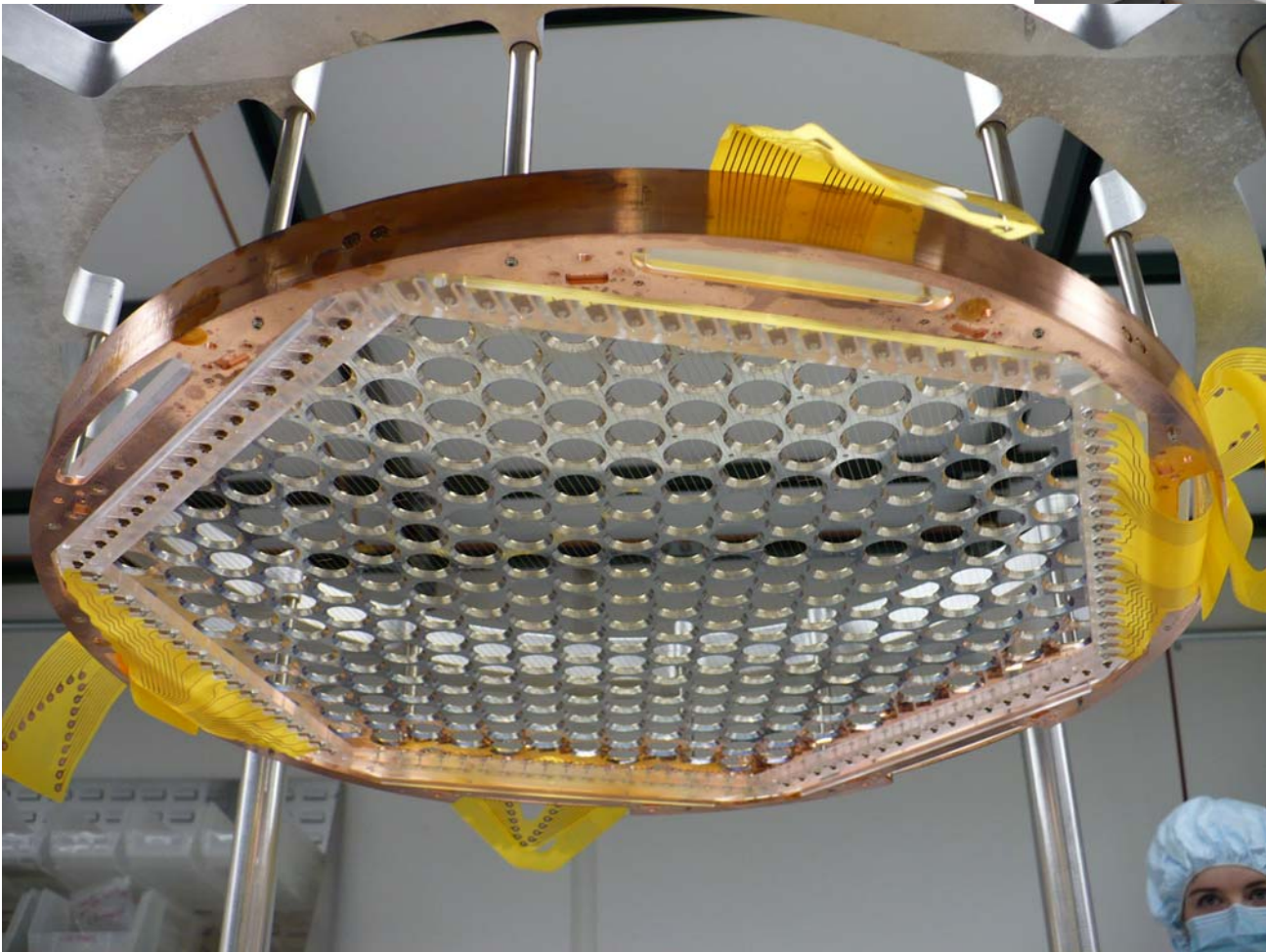
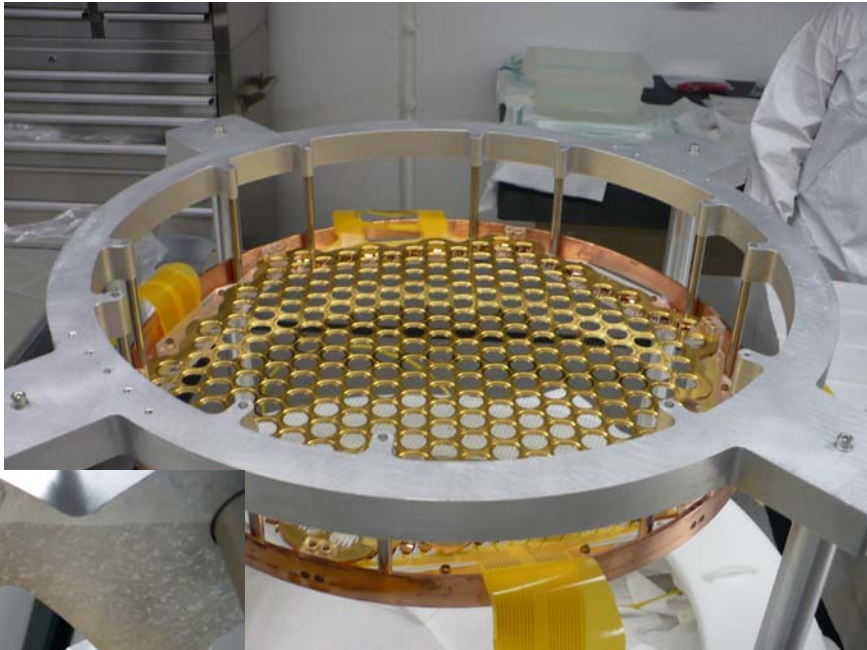
# One readout pancake



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Gio





## 650 bare LAAPD from Advanced Photonix

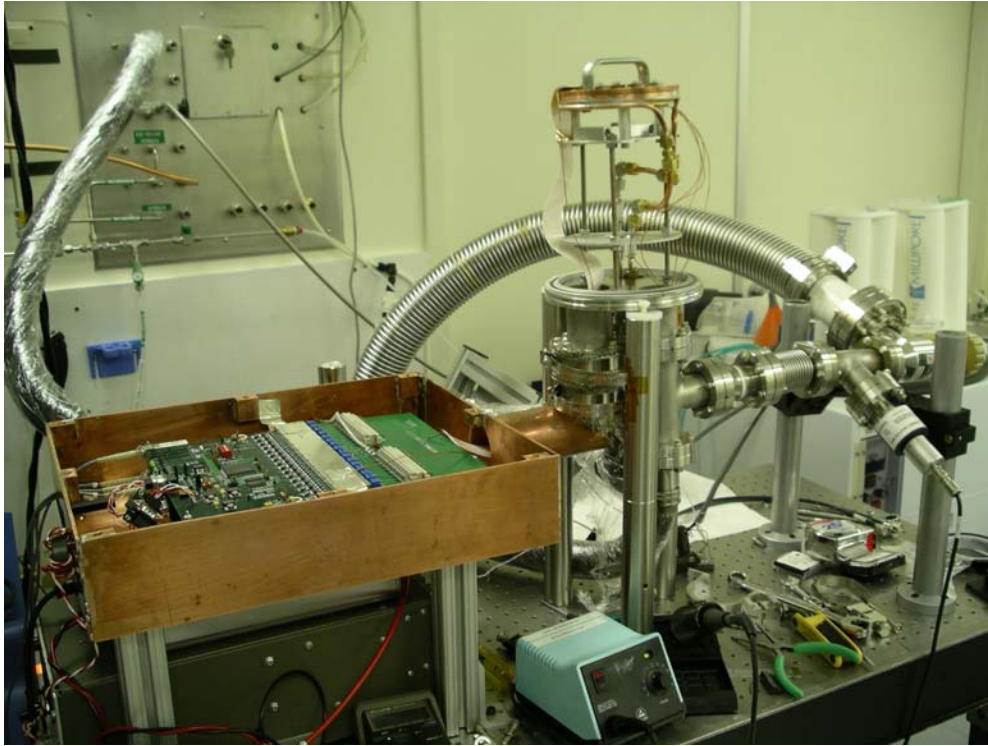


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Giorgio Gratta, EXO

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**~500 "Bare" LAAPD**

**APDs are ideal for our application:**

- very clean & light-weight,**
- very sensitive to VUV**

**QE > 1 at 175nm**

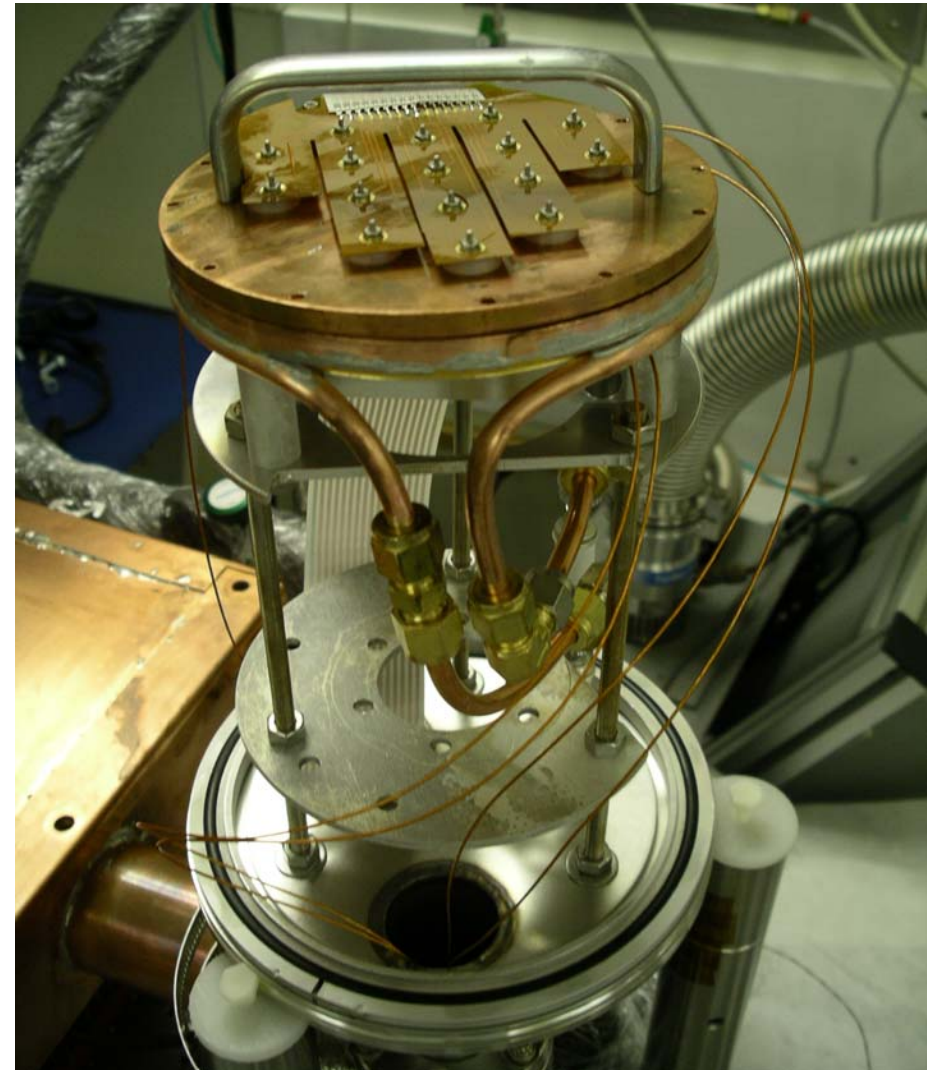
**Gain set at 100-150**

**V~1500V**

**$\Delta V < \pm 0.5V$**

**$\Delta T < \pm 1K$  APD is the driver  
for temperature stability**

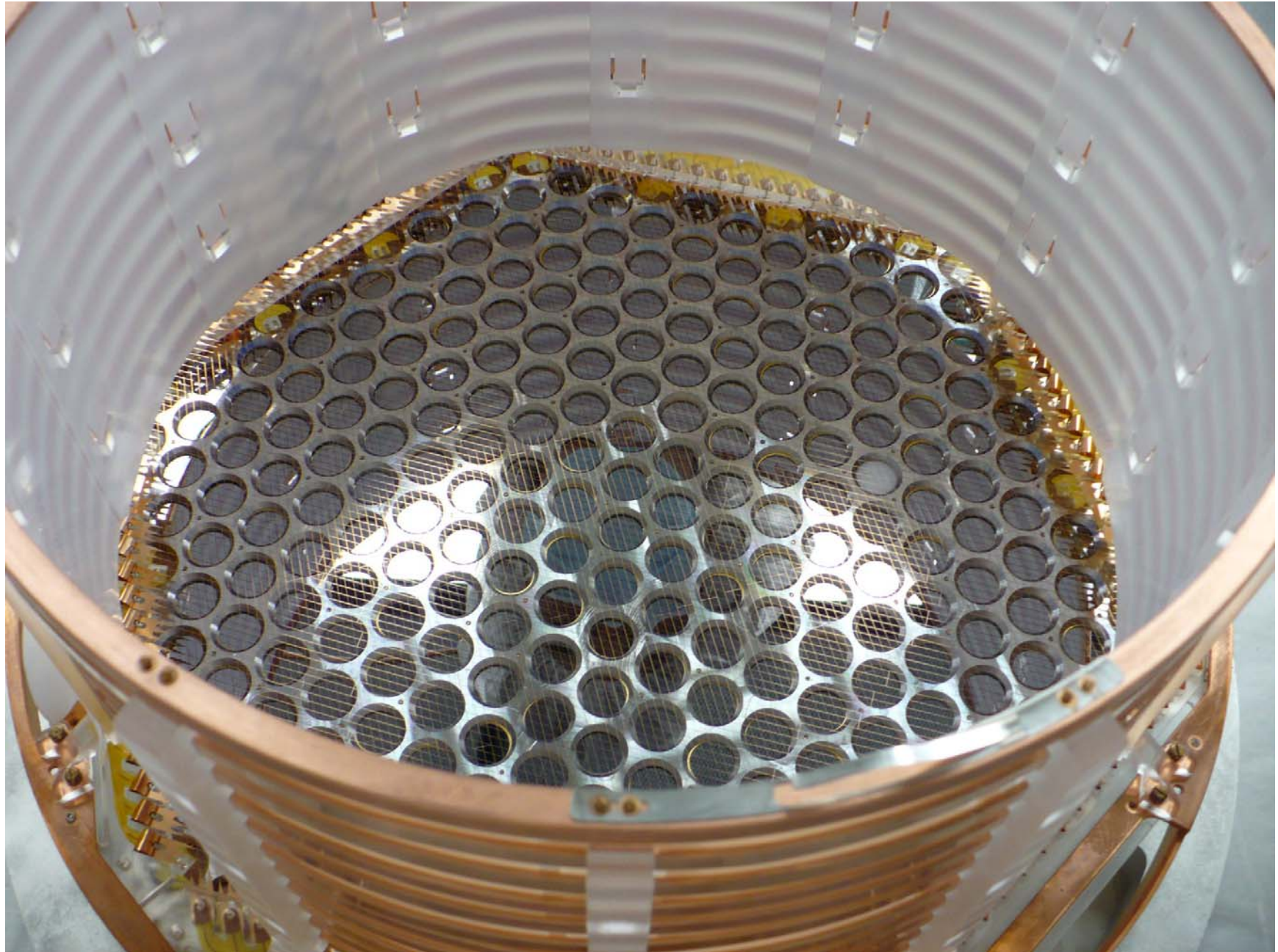
**Leakage current OK cold**



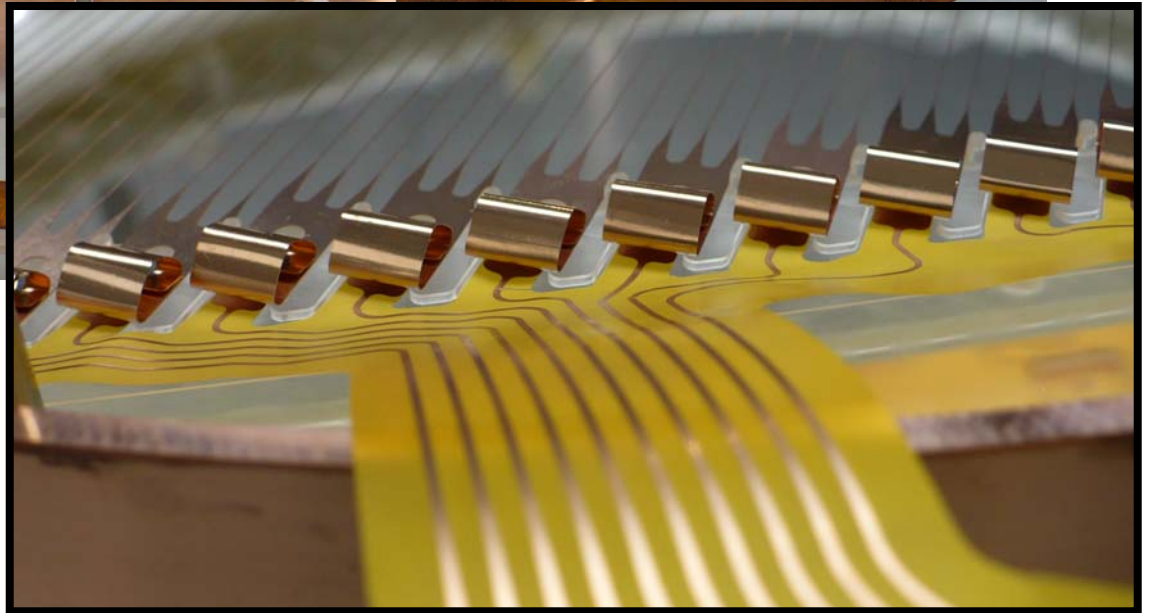
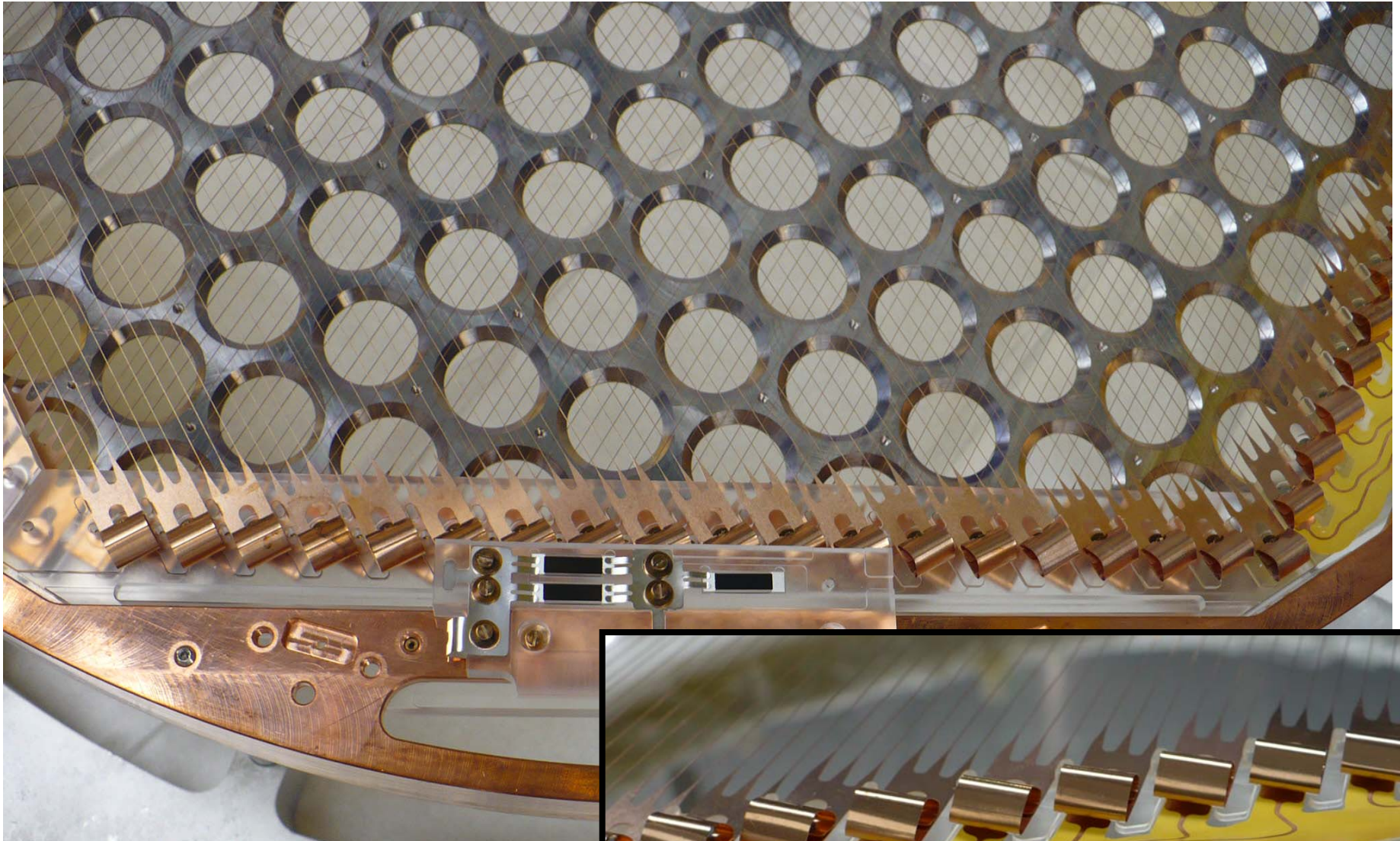




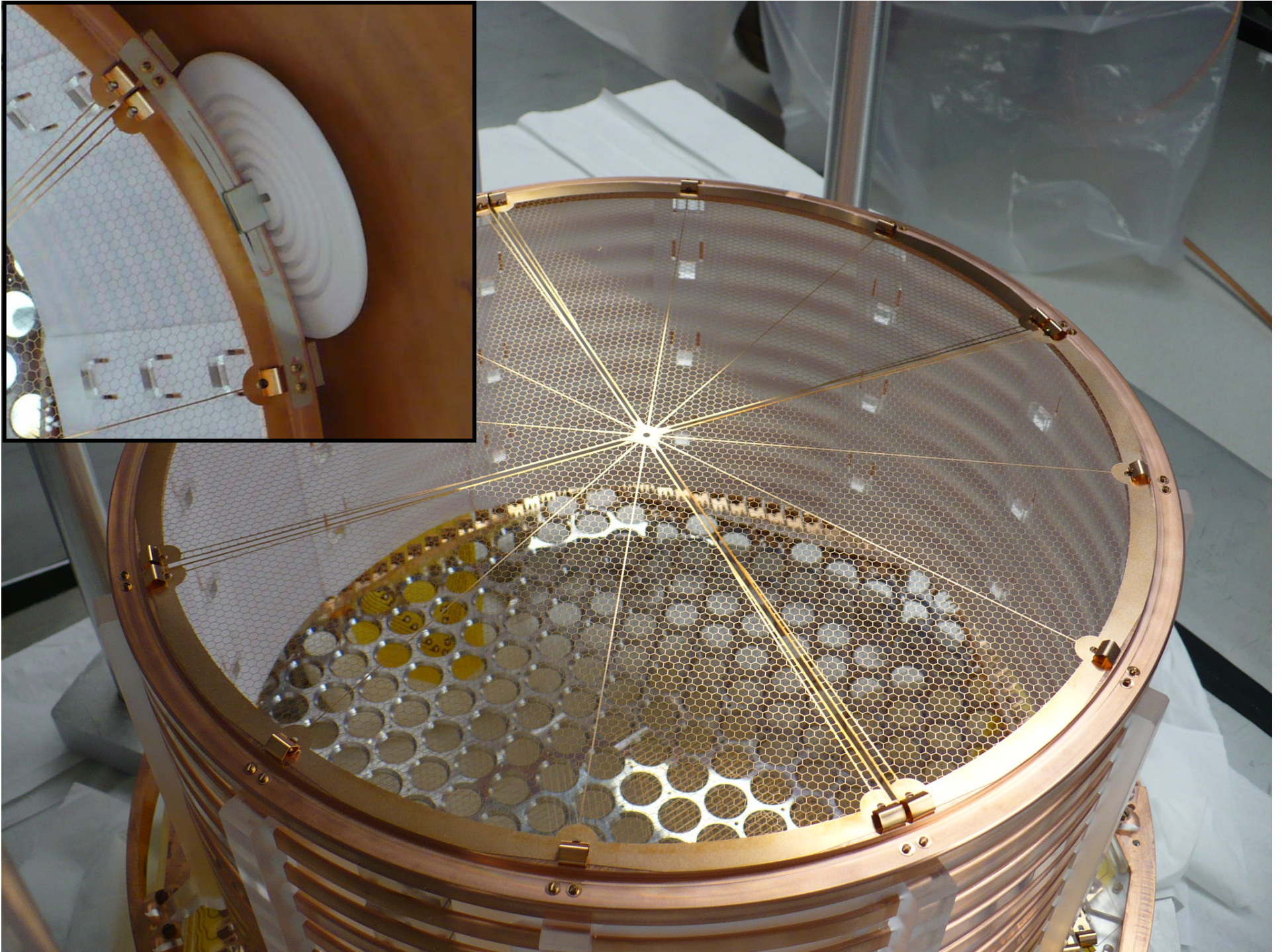




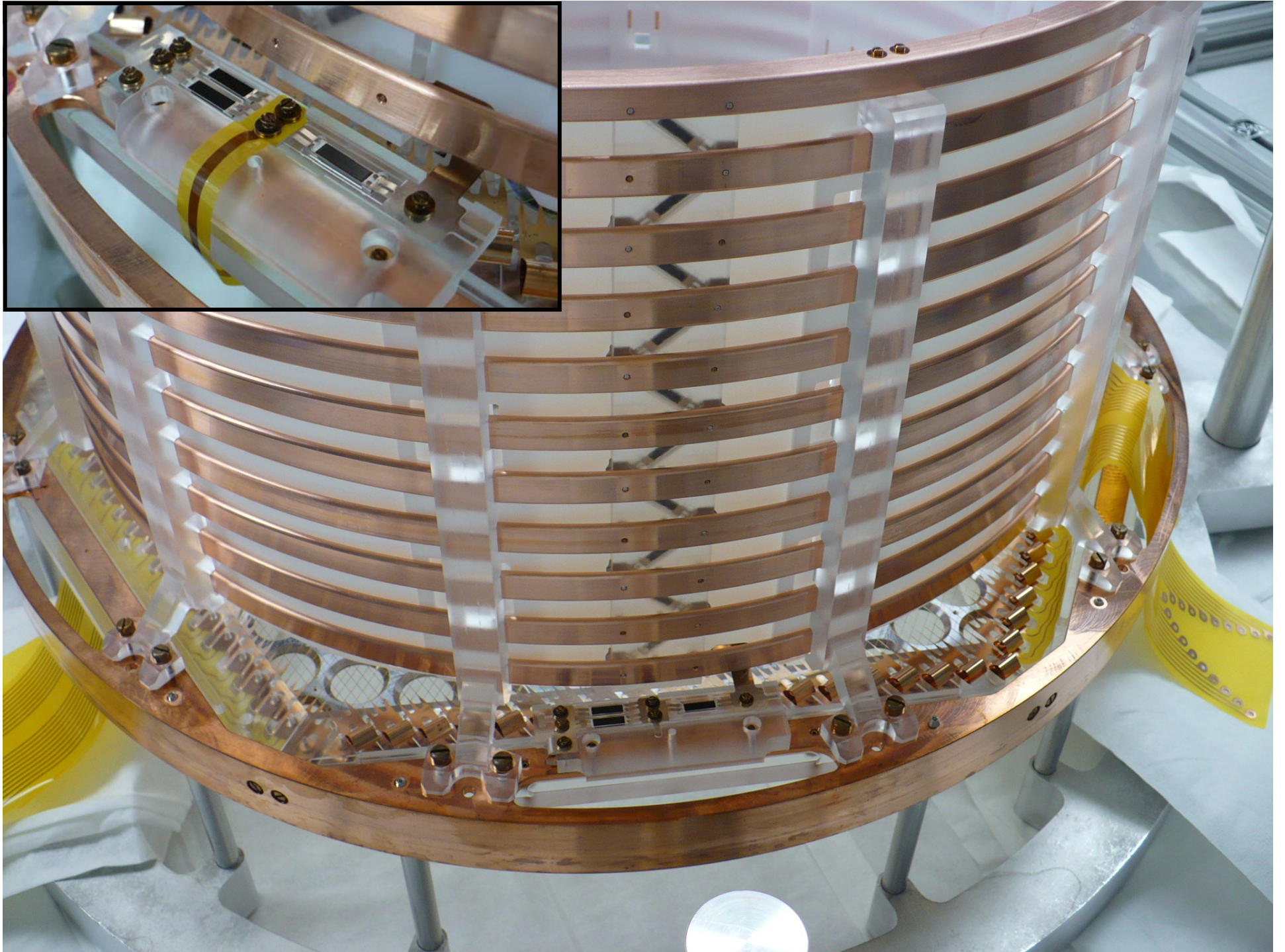
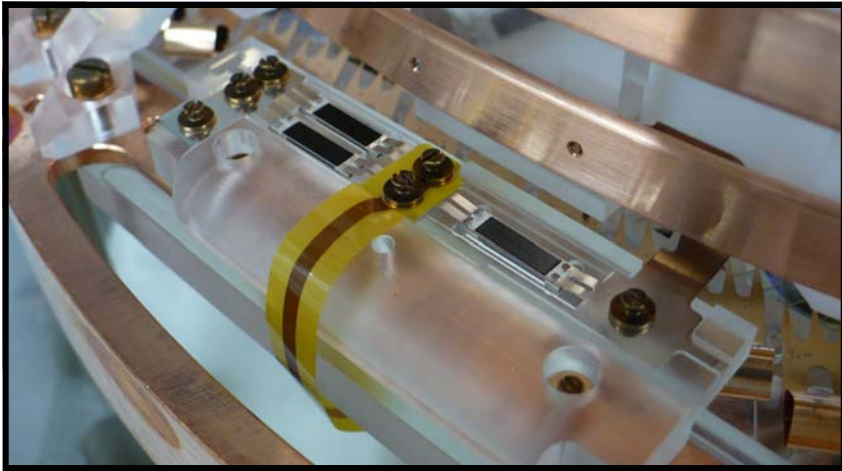




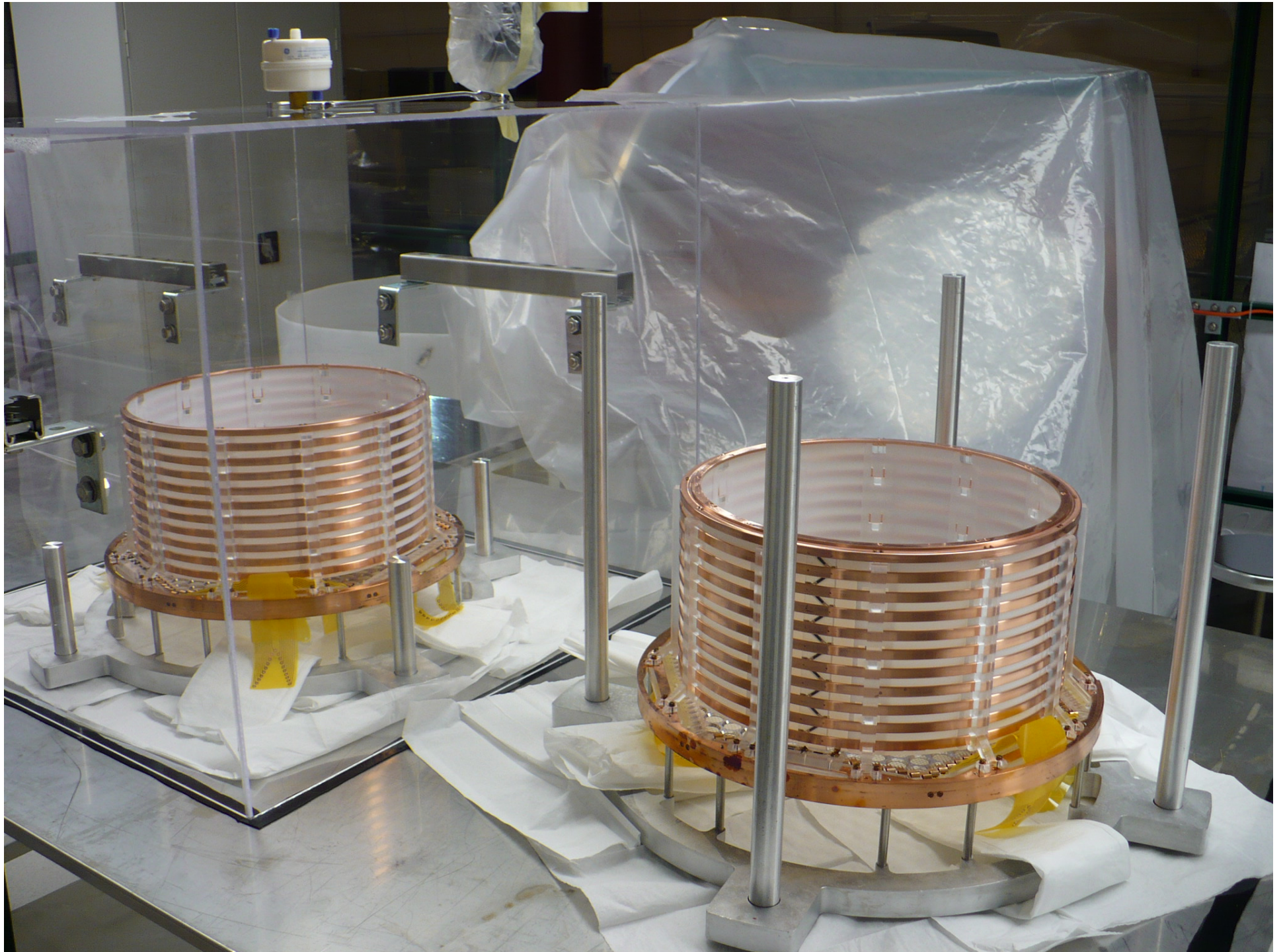












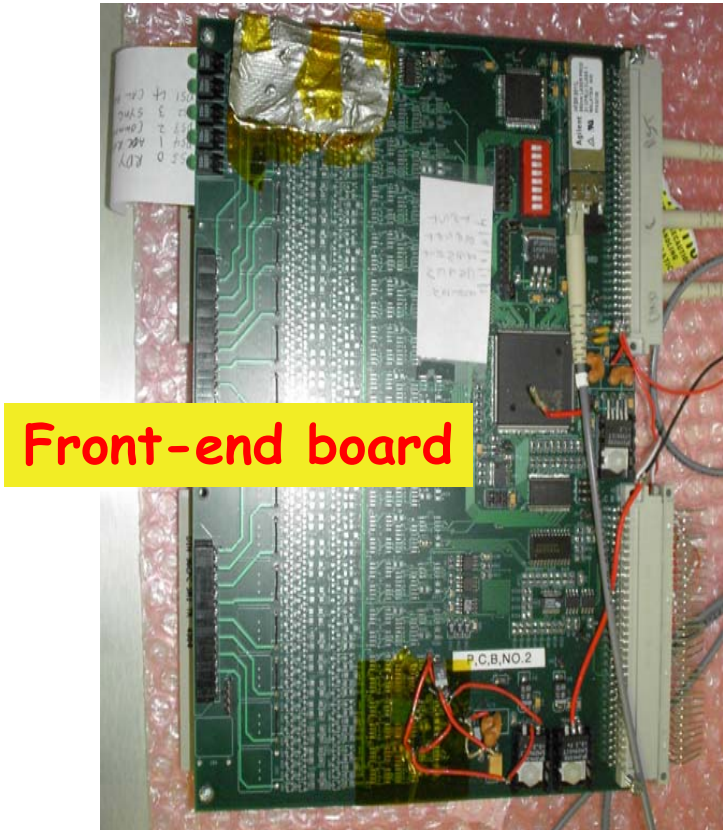




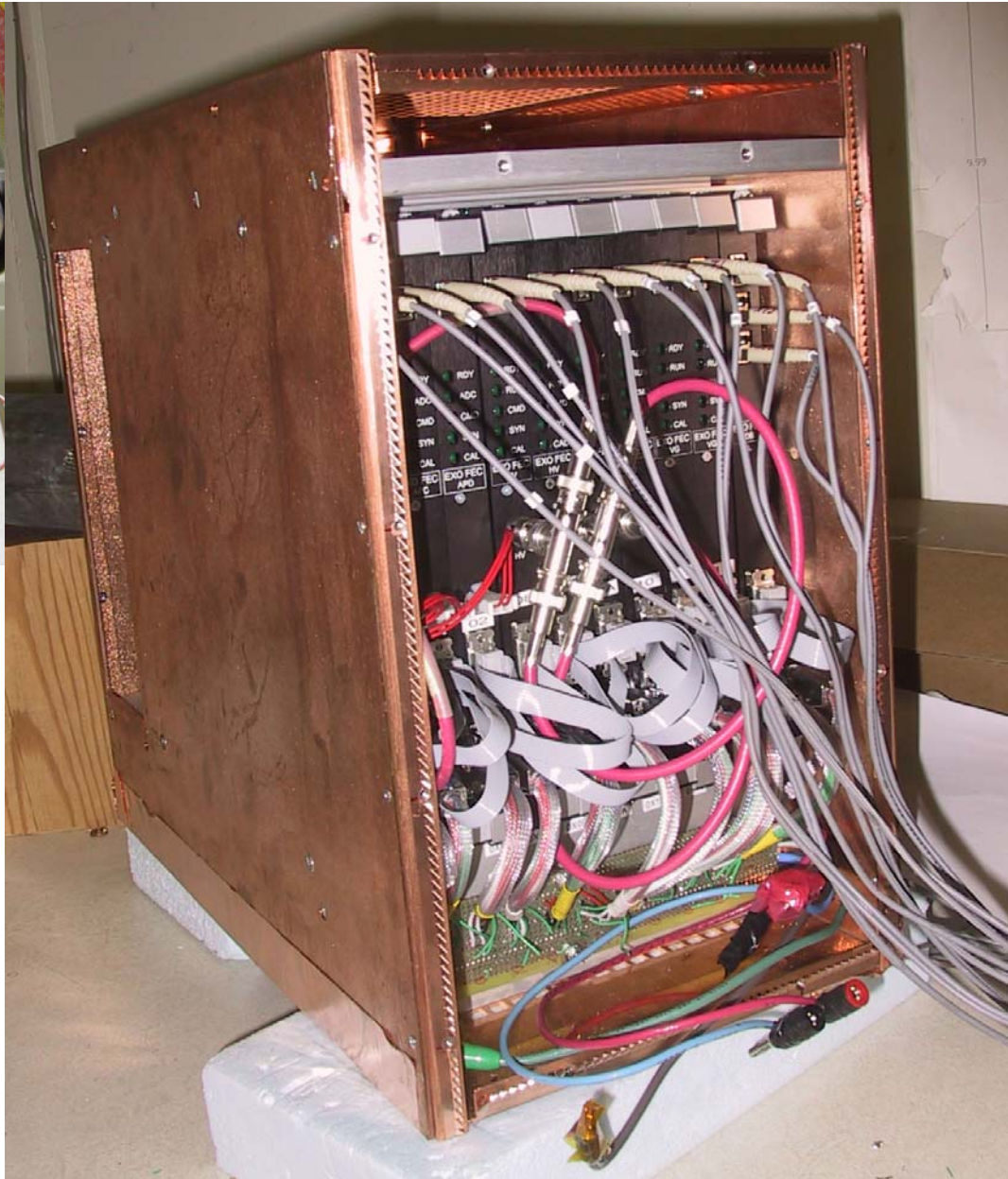




**Trigger module**

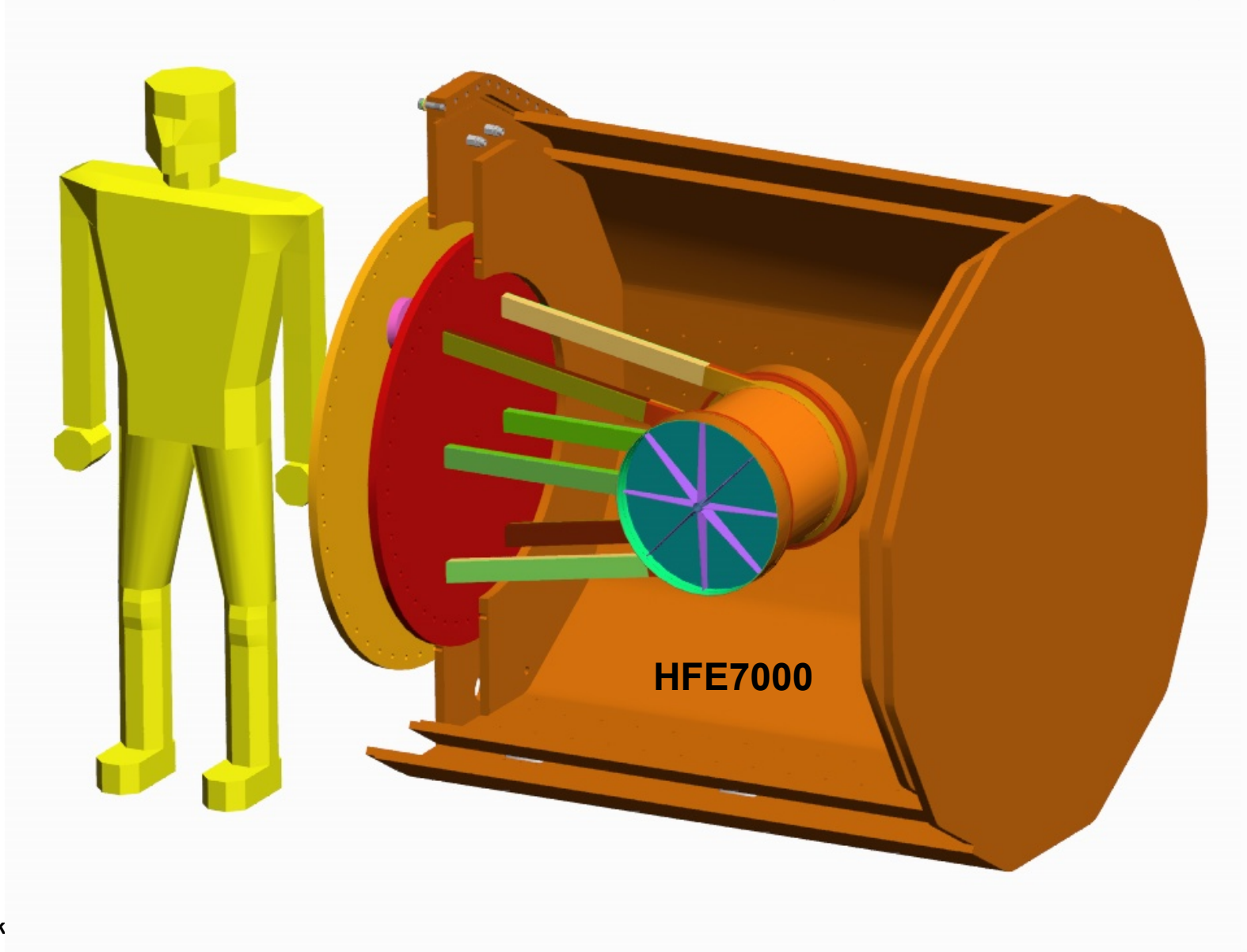


**Front-end board**



**Electronics completed, in calibration**

The TPC vessel sits in a low background, copper cryostat filled with ultra-clean HFE7000 refrigerant/shielding fluid

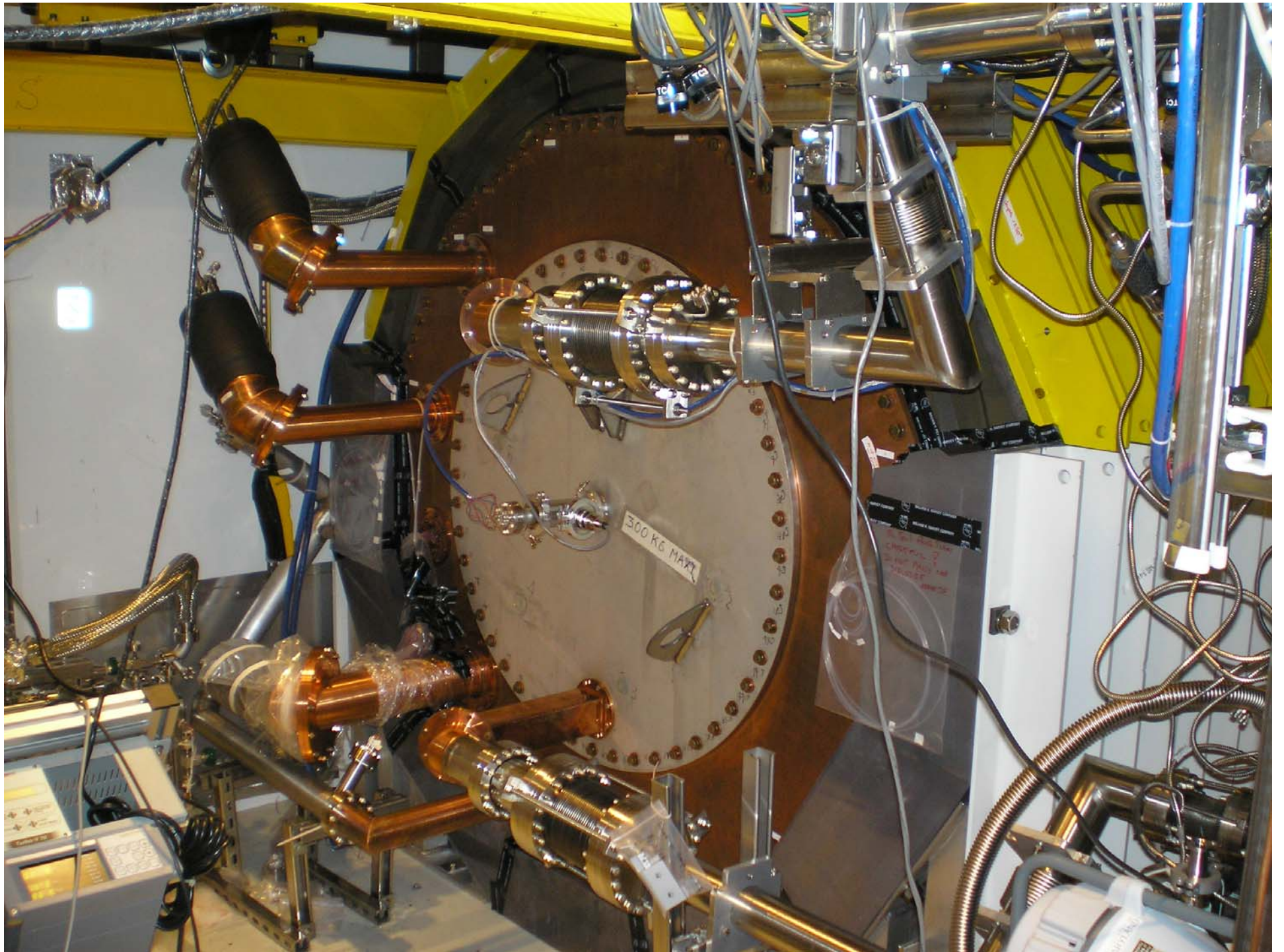




# Cryostat was fabricated at SDMS (Grenoble)



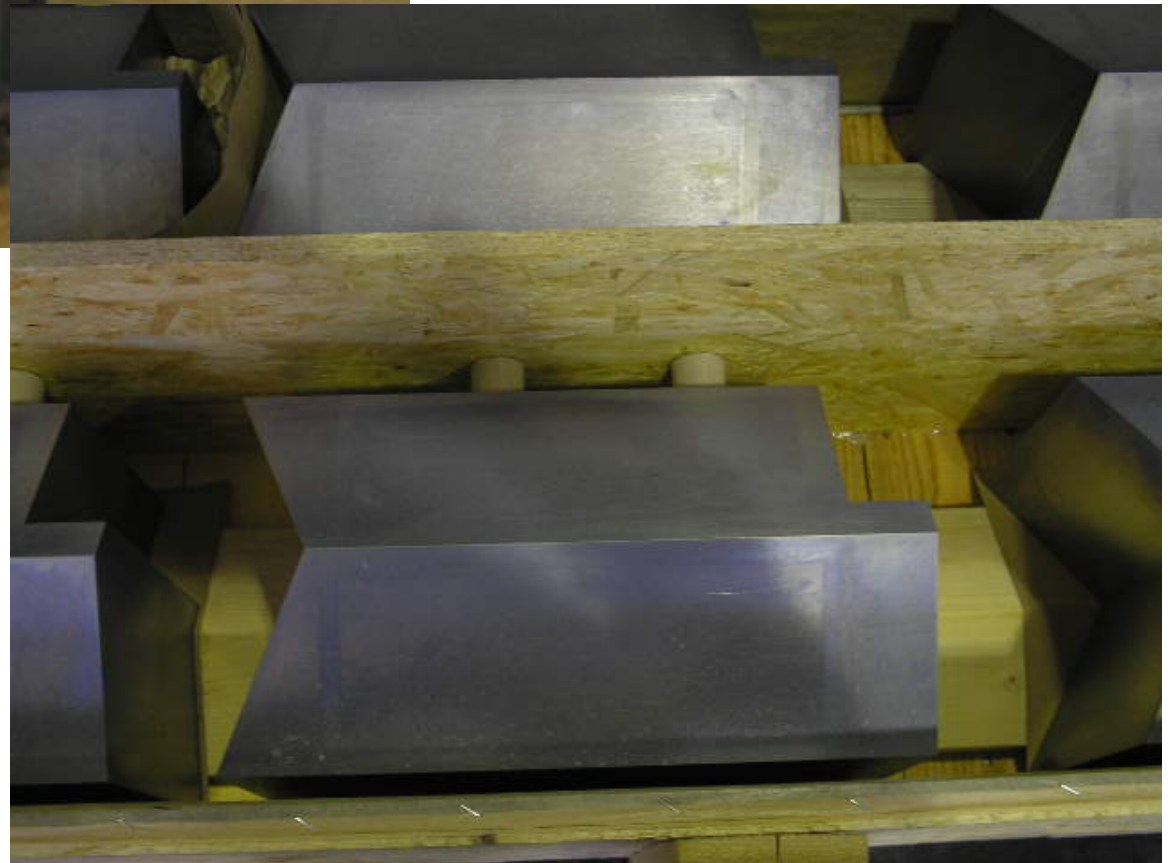








Each brick is epoxy painted at the factory in Germany to reduce Pb dust





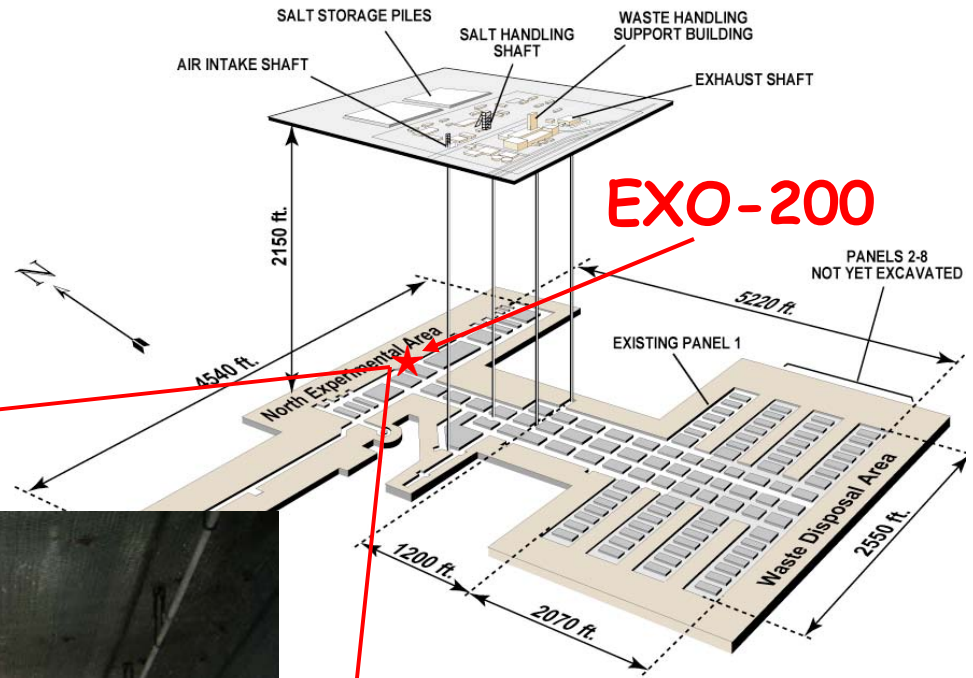
and shipped to the WIPP underground site



Jul 5, 07, the first shipment (Mods 5 & 6)



# EXO-200 at WIPP





Aug 22, 07, 6PM MDT  
Module 1 goes underground









Part of the EXO-200 infrastructure at WIPP, NM

## Massive effort on material radioactive qualification using:

- NAA
- Low background  $\gamma$ -spectroscopy
- $\alpha$ -counting
- Radon counting
- High sensitivity GD-MS and ICP-MS

*At present the database of characterized materials includes >300 entries*

## MC simulation of backgrounds

The impact of every screw within the Pb shielding is evaluated before acceptance

*NIM article published on the subject with entries for 225 materials [D.Leonard et al., Nucl. Inst. and Meth. A 591 3 (2008) ]*





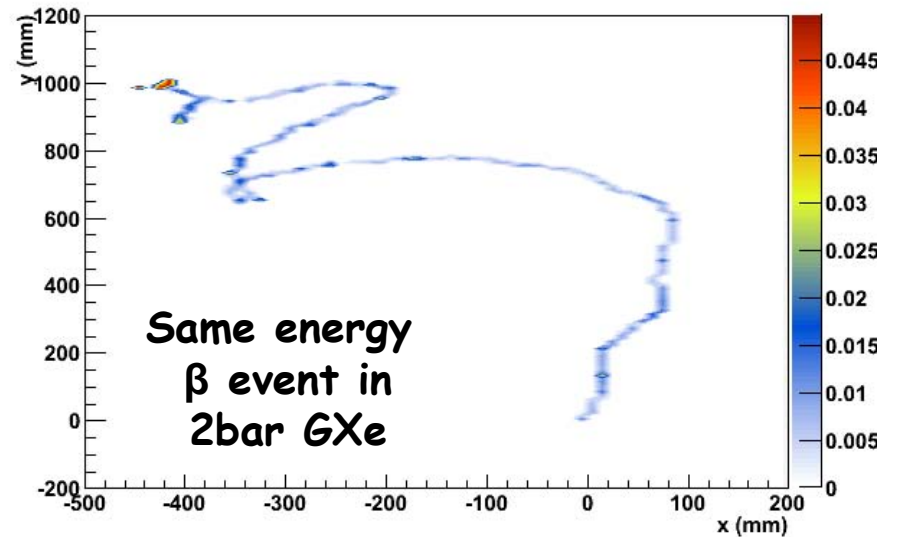
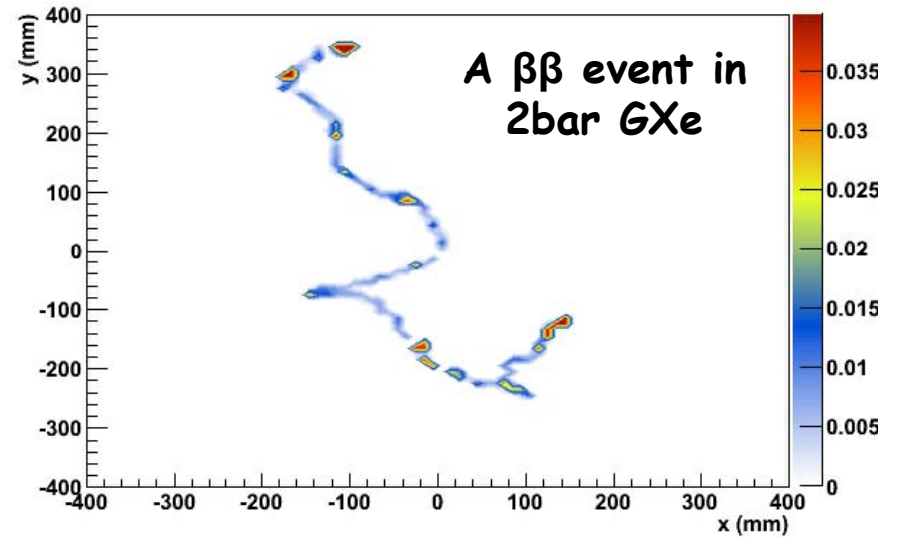
# GXe R&D program



# Gas Xe has the potential of providing event topology information along with very good energy resolution

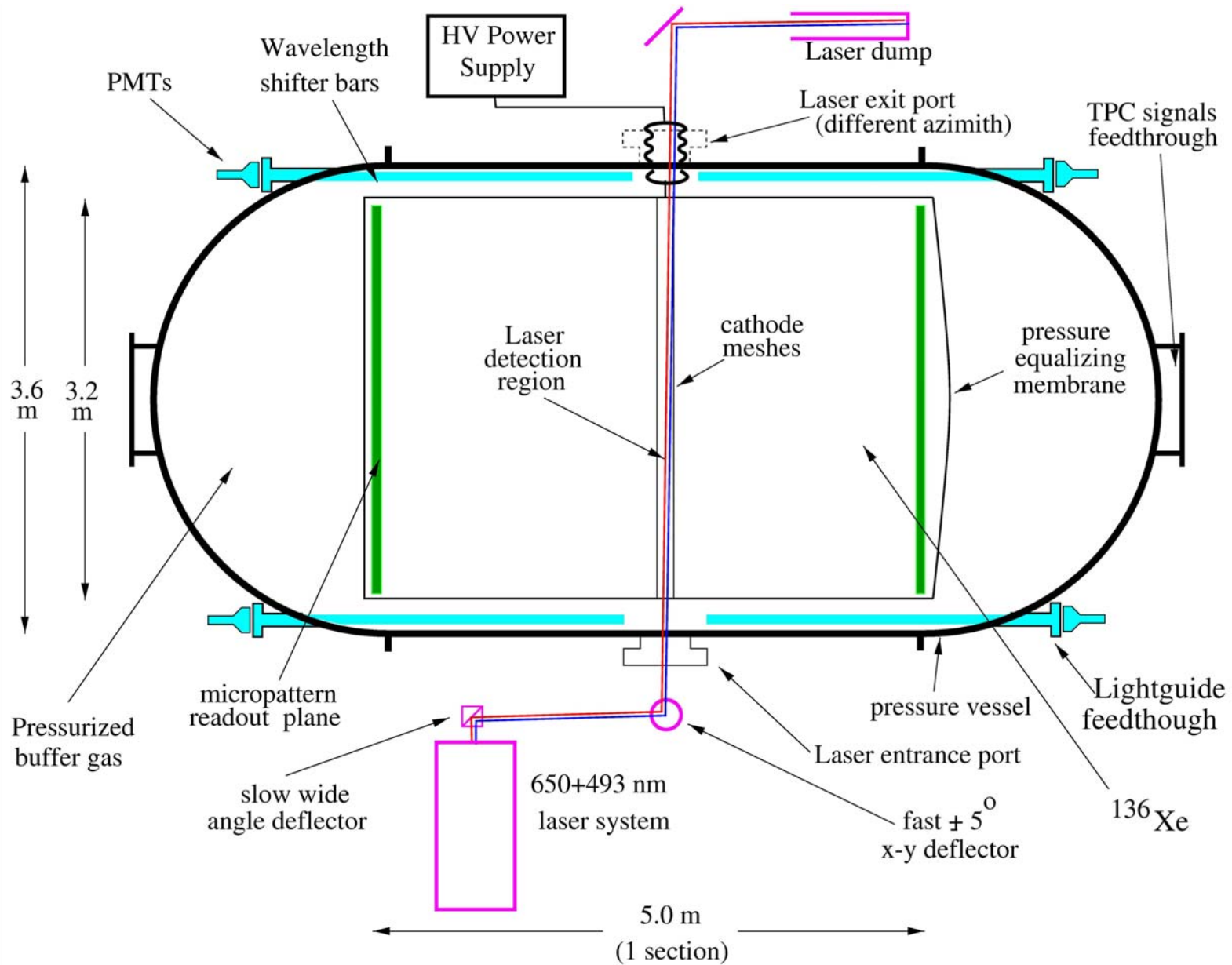
How well this works depends of course on the pressure.

1000 kg, OC, cylinder with 2R=L		
Pressure (atm)	Volume (m <sup>3</sup> )	Diameter or Length (m)
1	179	6.1
2	89	4.8
5	35	3.5
10	17	2.8



# First EXO concept

*M. Danilov et al (EXO coll), Phys Lett 480 (2000) 12*

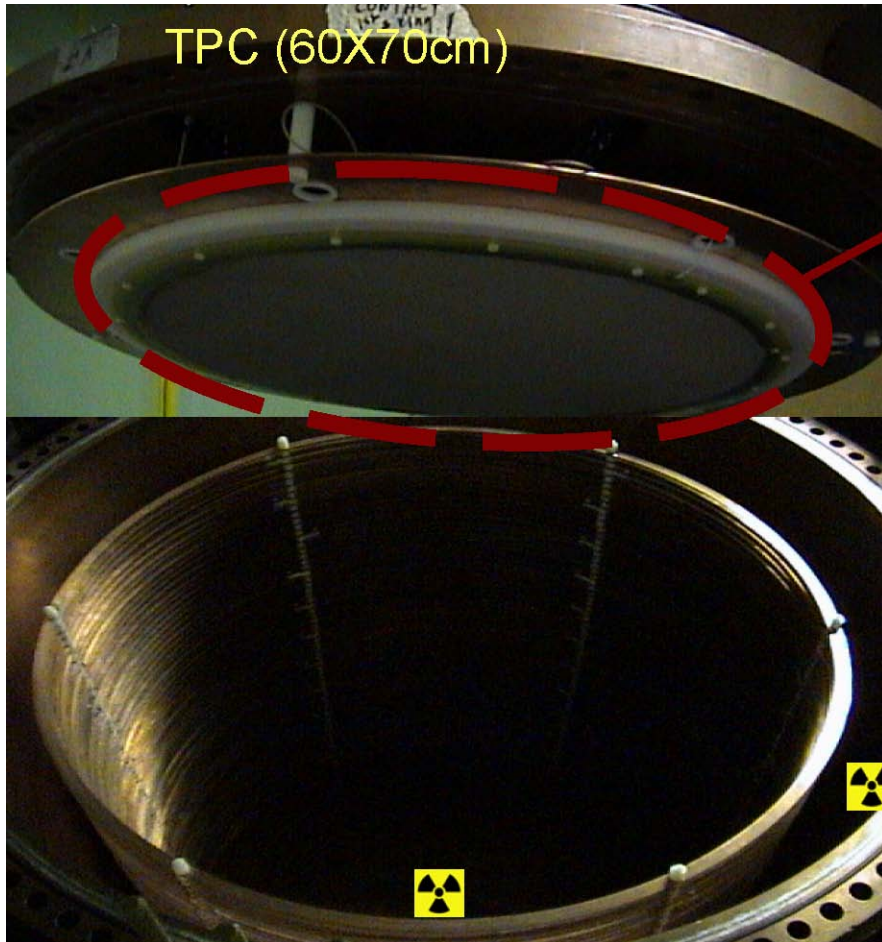




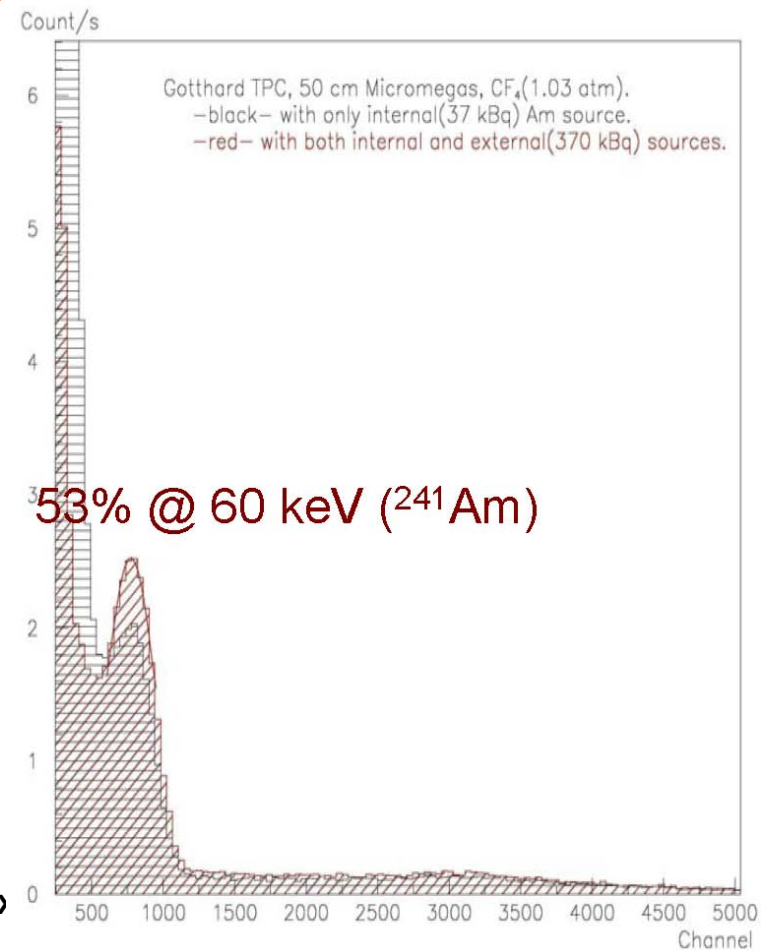
# Old Gotthard TPC has a new life in EXO R&D



50cm diameter  
micromegas



TPC (60X70cm)

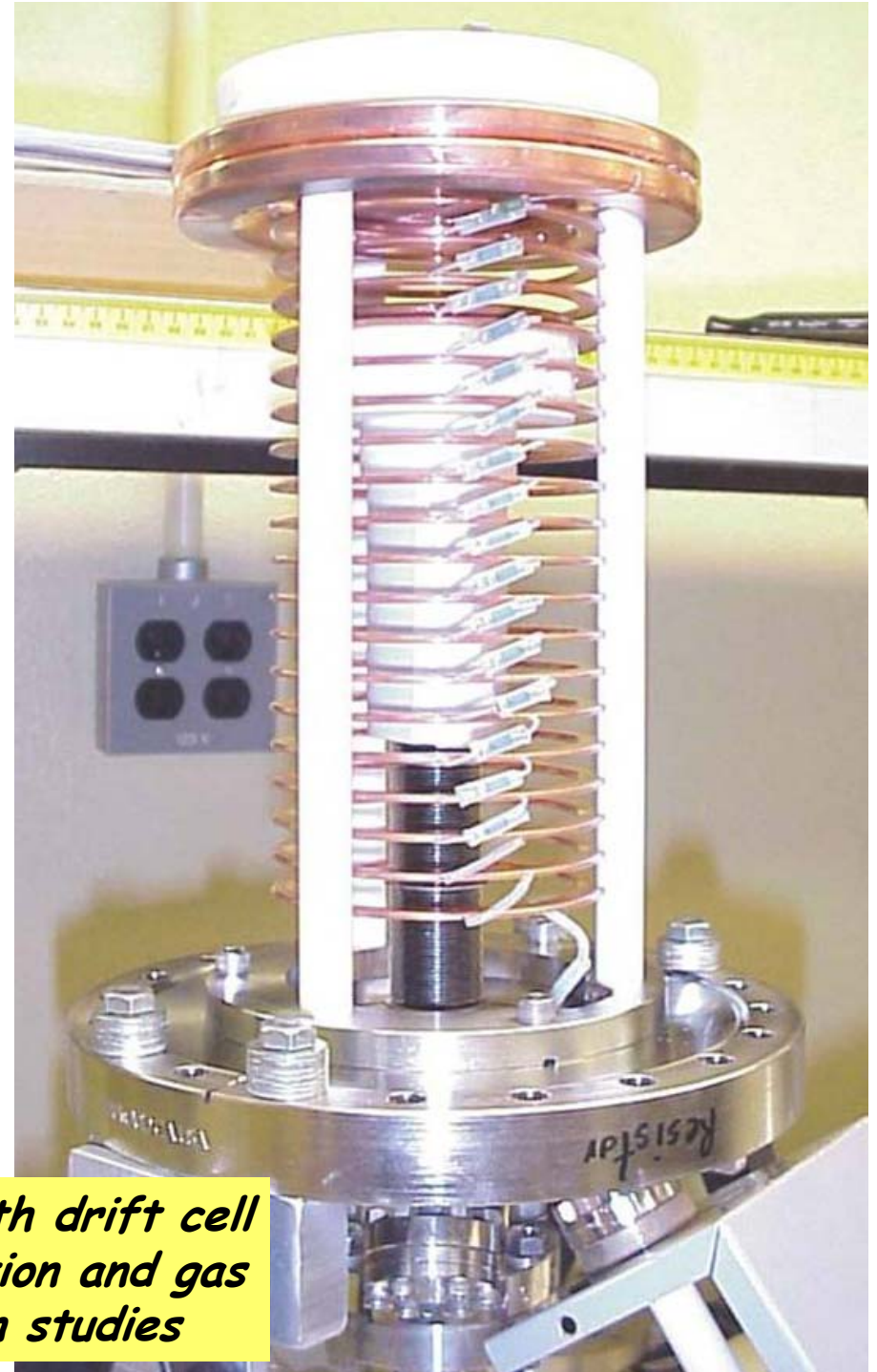


Conventional gain, including GEMs and Micromegas, has a number of problems for EXO:

1. Energy resolution not ideal
2. Additive quenchers likely eat up the Ba or at least add another constraint to an already challenging problem

If feasible, pure Xe option appears better suited

*Variable length drift cell for resolution and gas composition studies*





Next Step: build a high pressure chamber able to contain  $\sim 1\text{MeV}$  electrons with electroluminescent readout

*This chamber will NOT be low background but will use methods that can be made low background*

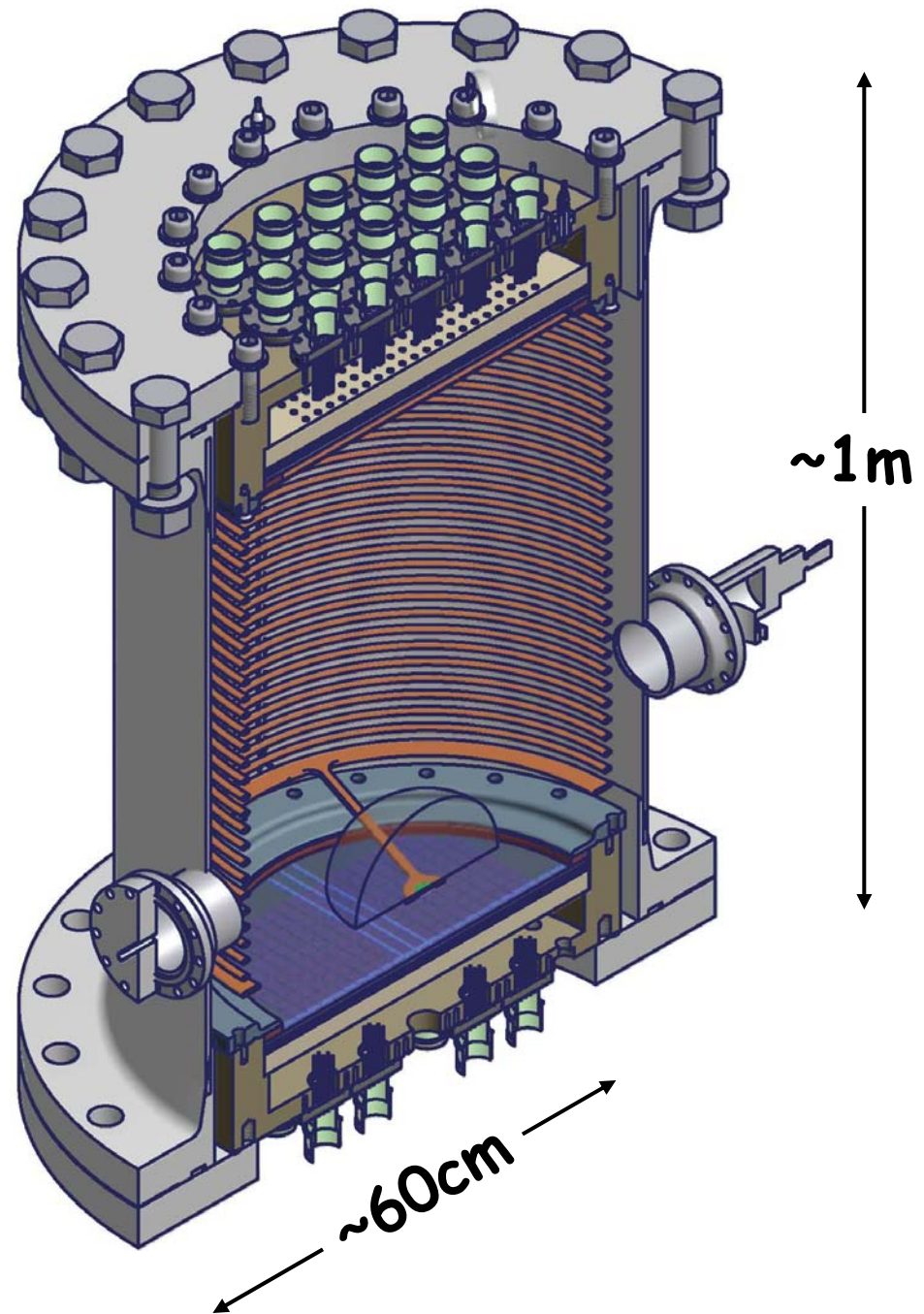
No PMTs. Primary VUV detection option: CsI modular structure will allow the test of gaseous photocathodes (and, possibly wavelength shifting fibers or APDs)

- Goals:
- Measure energy resolution for  $1\text{MeV}$  electrons with segmented readout
  - Evaluate tracking ability and background rejection at different pressures
  - Study E field and readout options
  - Validate the concept for a large detector

Will try to have chamber ready in summer 09 at Carleton

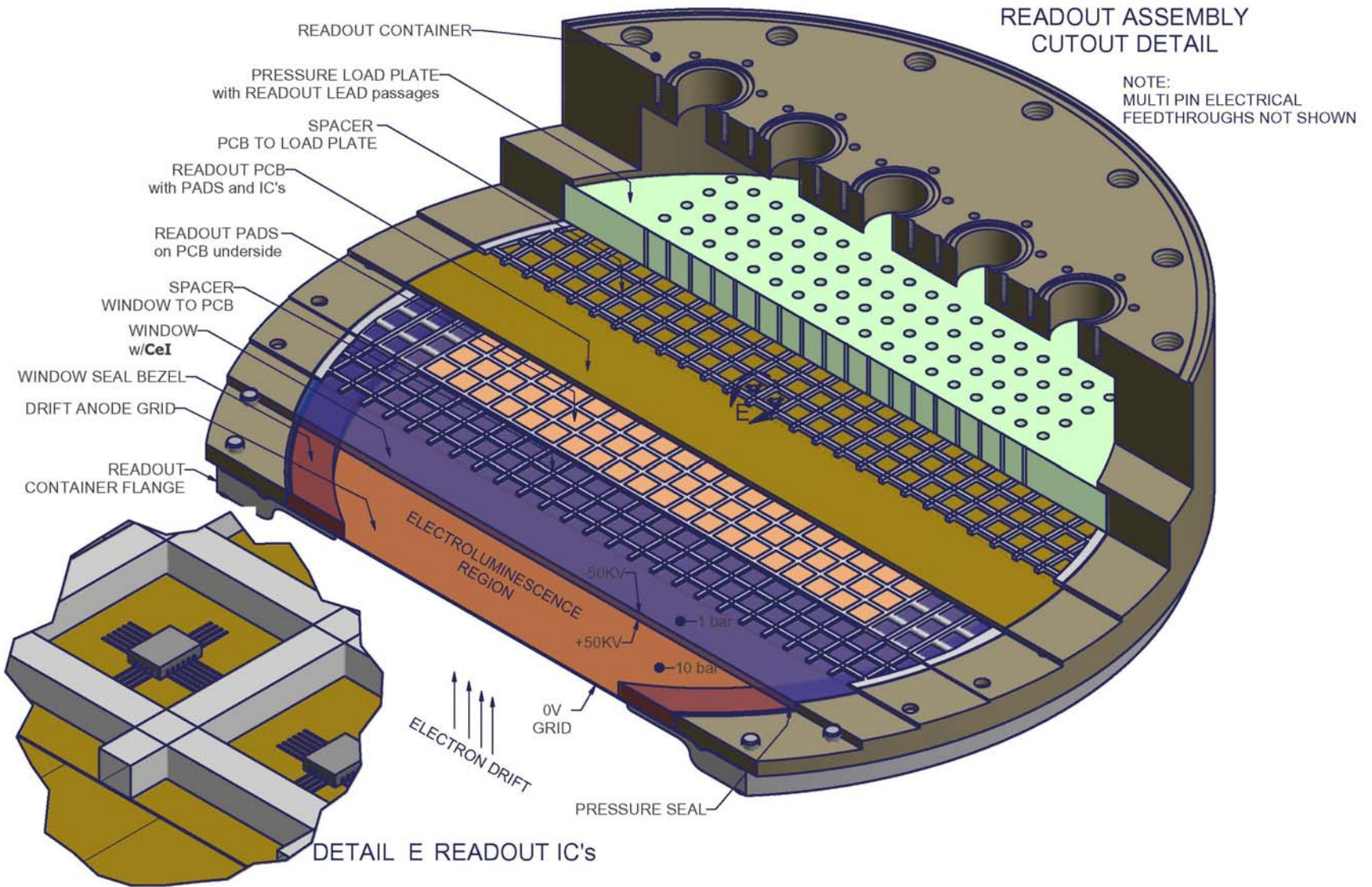
(although competition for personnel to start EXO-200 is harsh)

$V \sim 300 \text{ l}$   
 $P \leq 10 \text{ atm}$



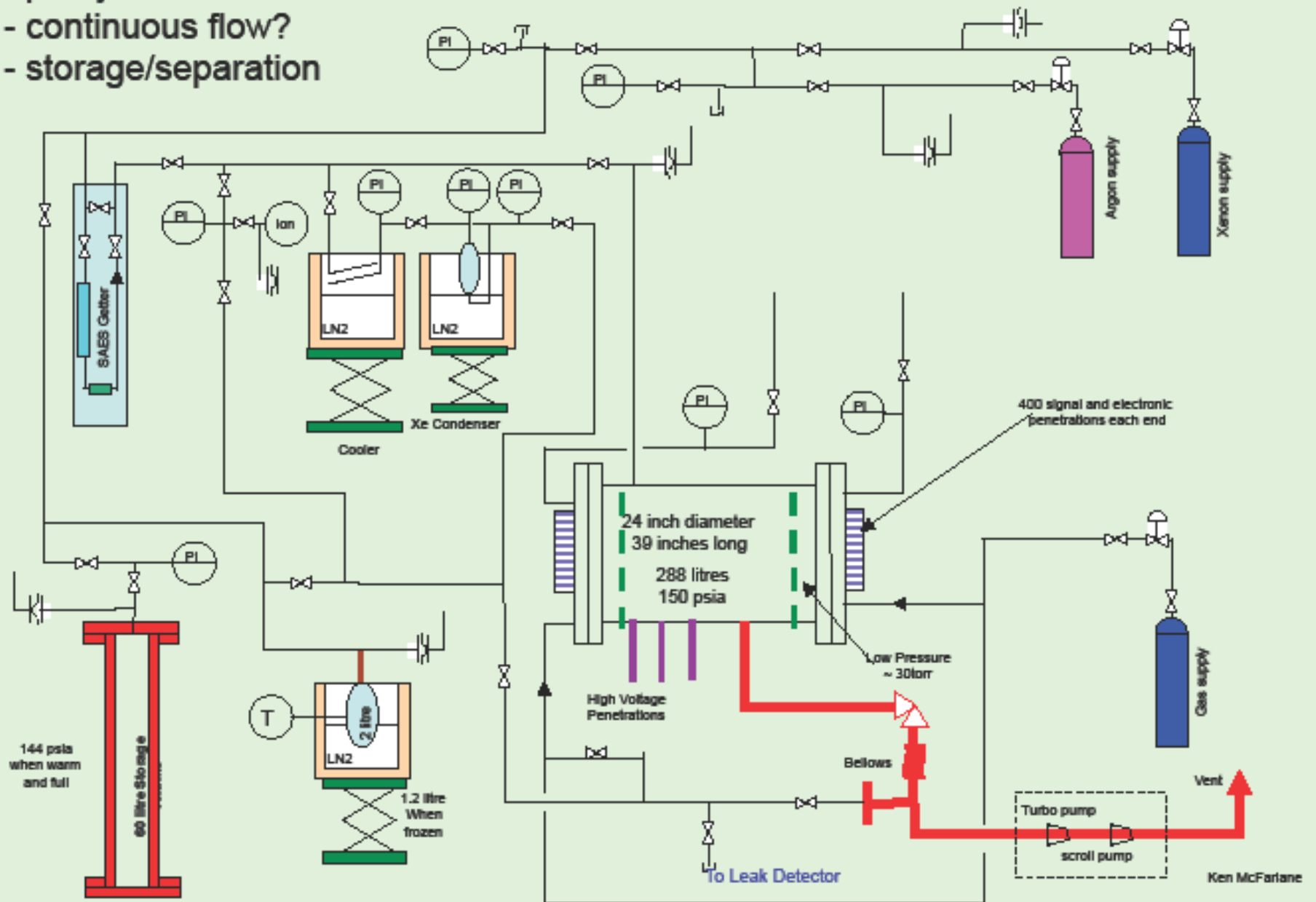


# CsI readout plane



# Process Systems

- purity monitor
- continuous flow?
- storage/separation





## Gaseous photocathode R&D



UV flash  
lamp

Drift  
cell

Distillation  
system

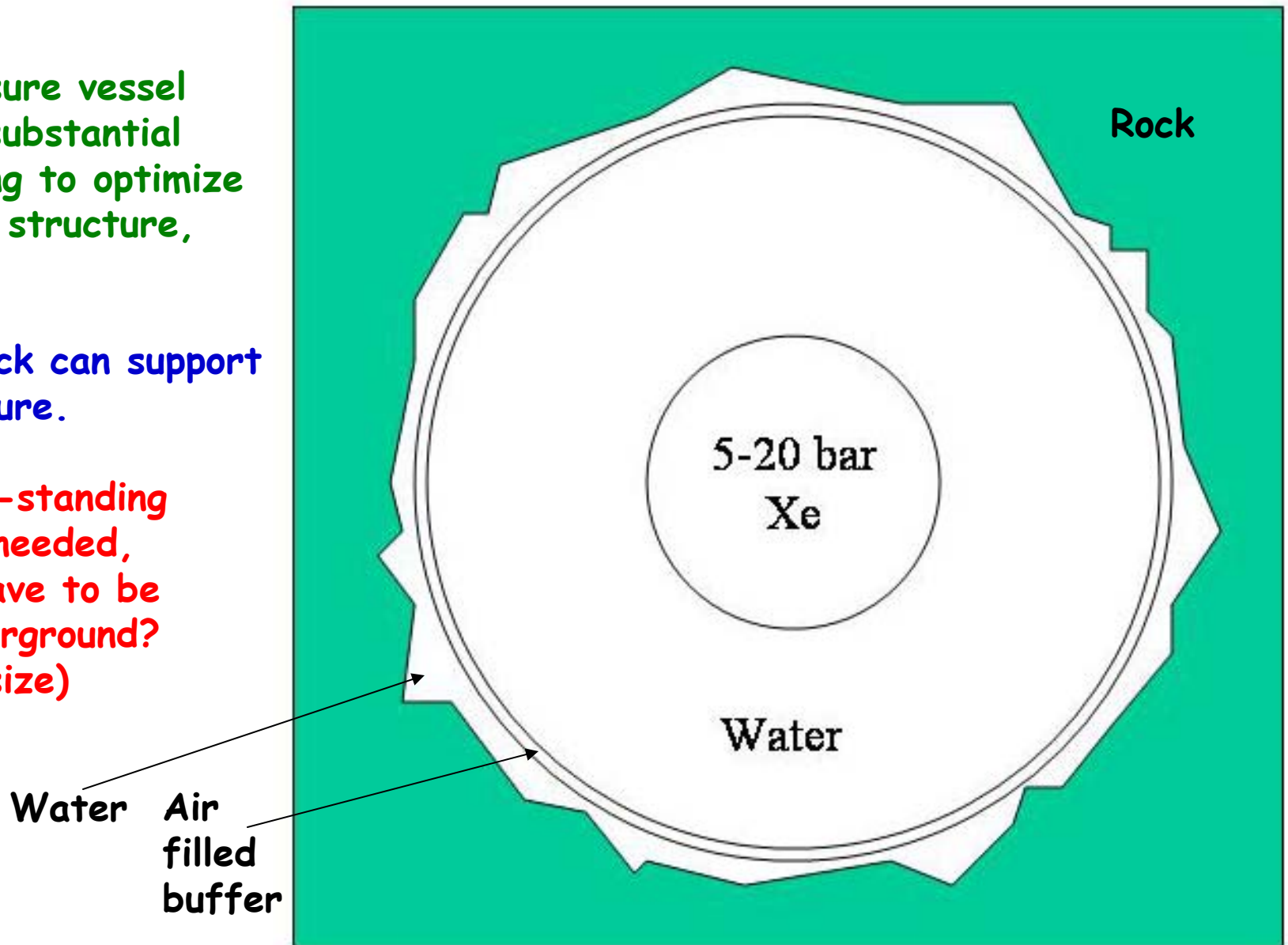
Name	Composition	Ioniz potential (eV)	Vap press at 20C (torr)
TMAE	$C_{10}H_{24}N_4$	5.4	0.35
TEA	$C_{17}H_{44}N_4$	7.5	55
exo-2-chloronorbornane	$C_7H_{11}Cl$	1.65	3.3
(S)-exo-5-norbornen-2-ylchloride	$C_7H_{19}Cl$	1.1	1.6

## Ideas about pressure vessel

The pressure vessel requires substantial engineering to optimize shielding, structure, access, ...

Maybe rock can support the pressure.

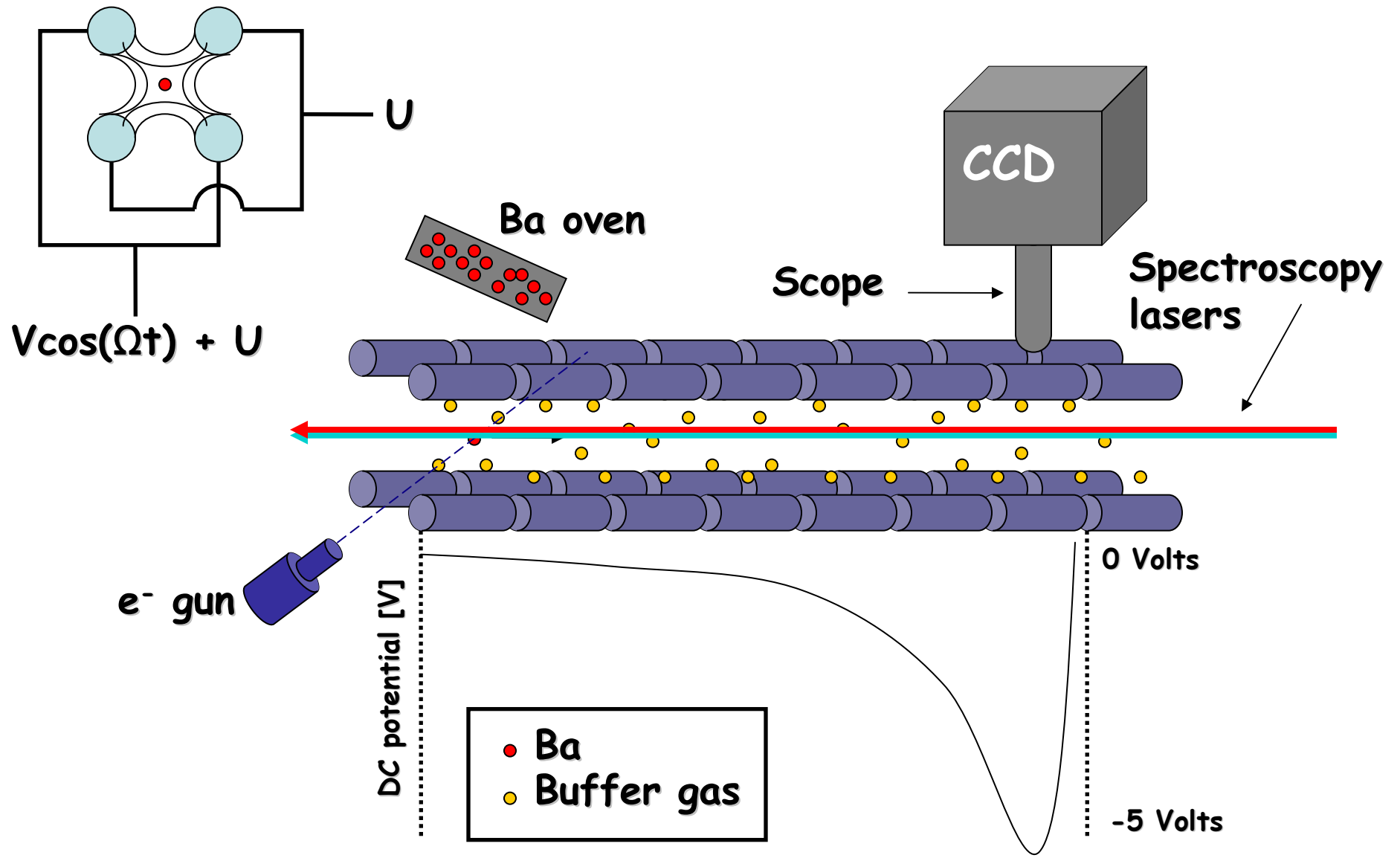
If a free-standing vessel is needed, does it have to be built underground? (weight, size)



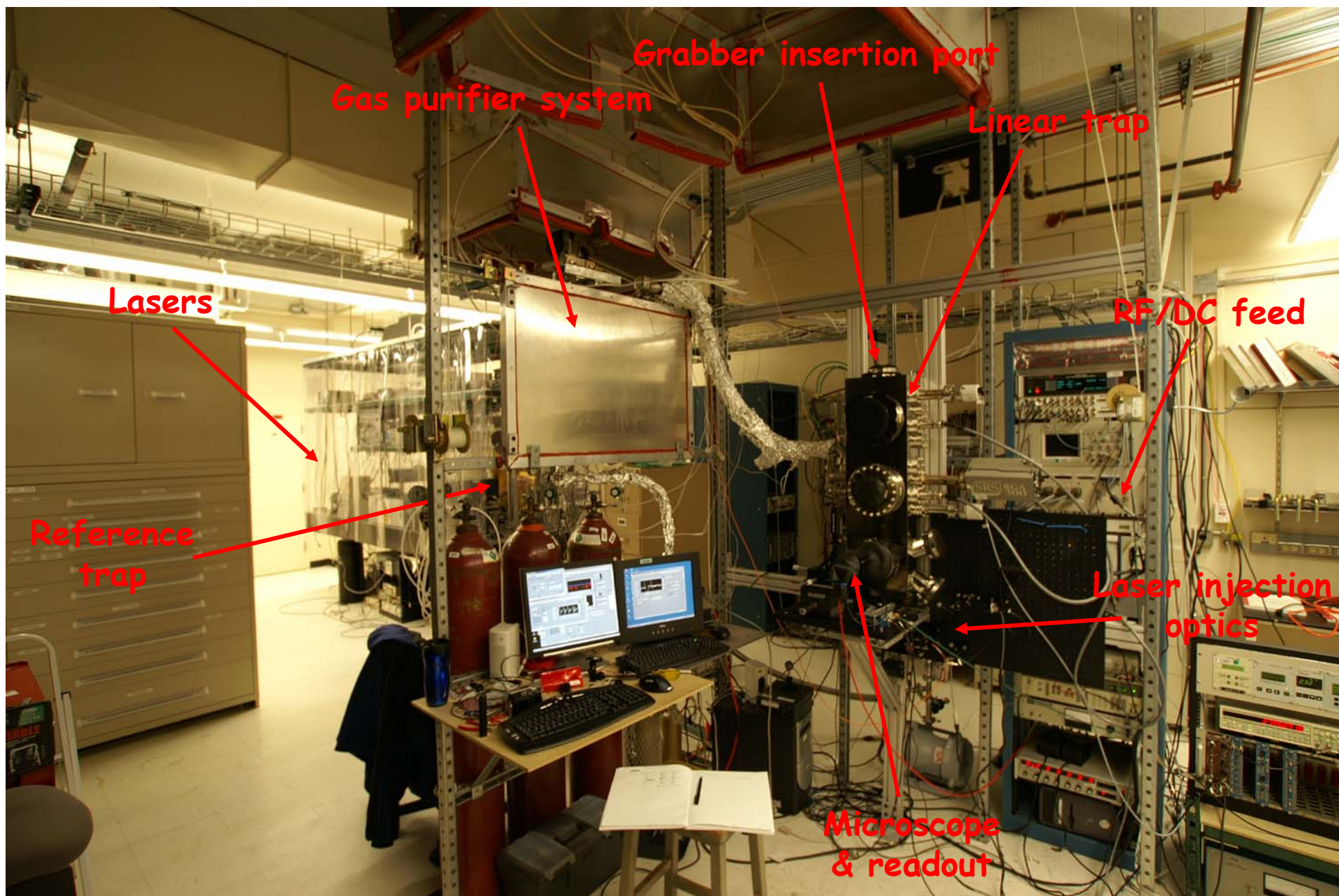


# Ba tagging R&D

# Ba<sup>+</sup> identification in a Linear Ion Trap







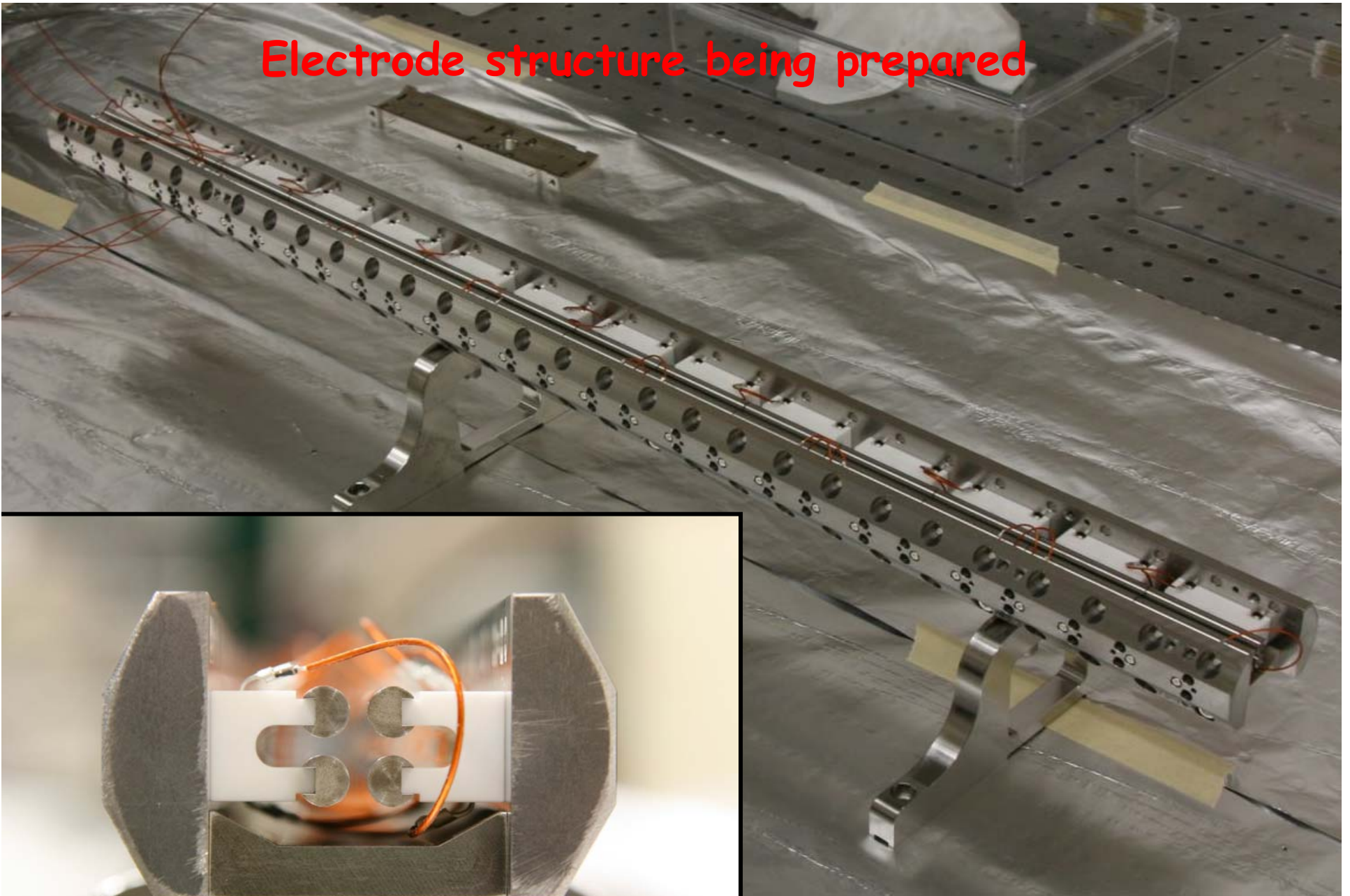
Berkeley Mar 19, 2009

Giorgio Gratta, EXO

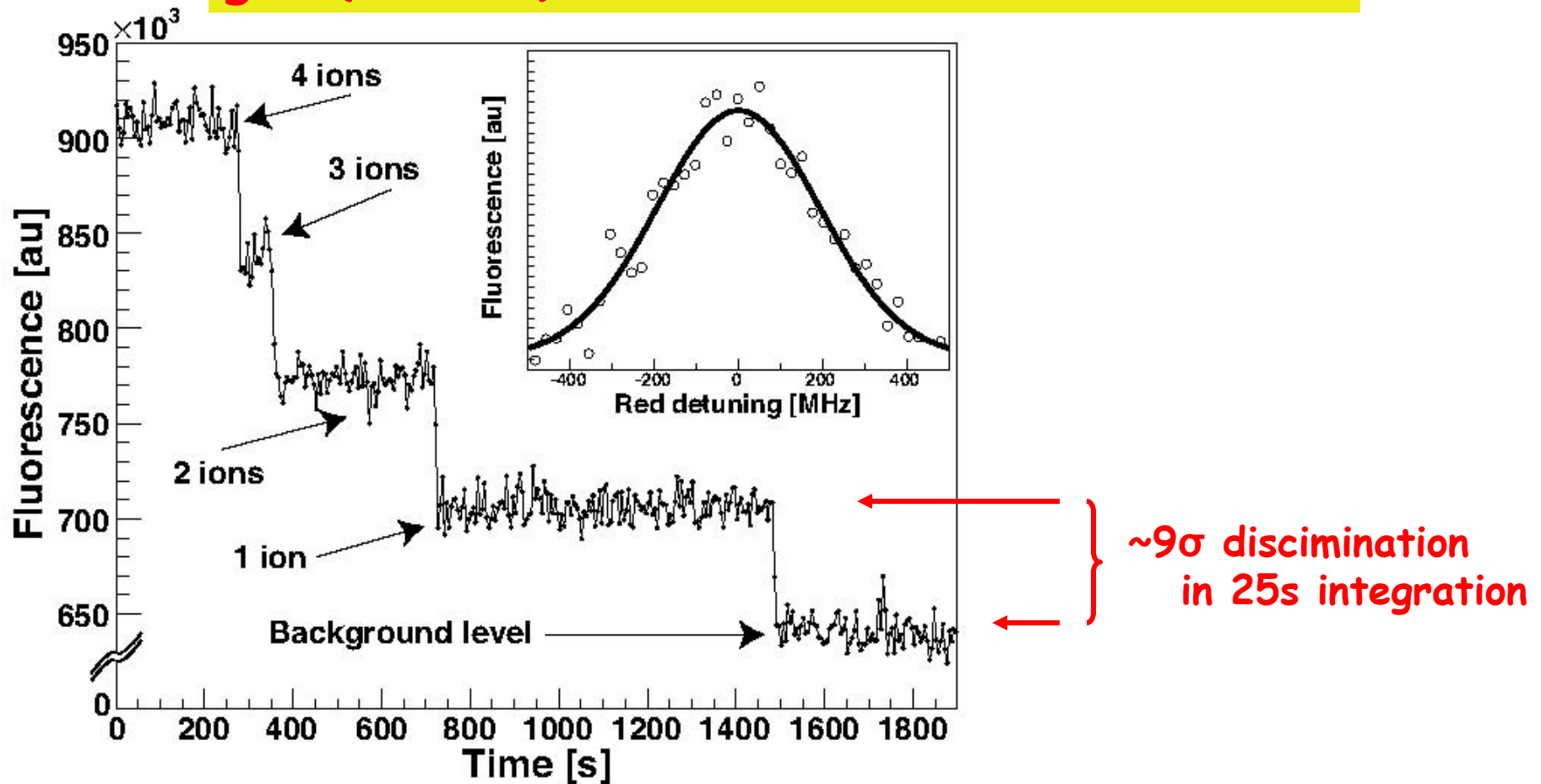
74



Electrode structure being prepared



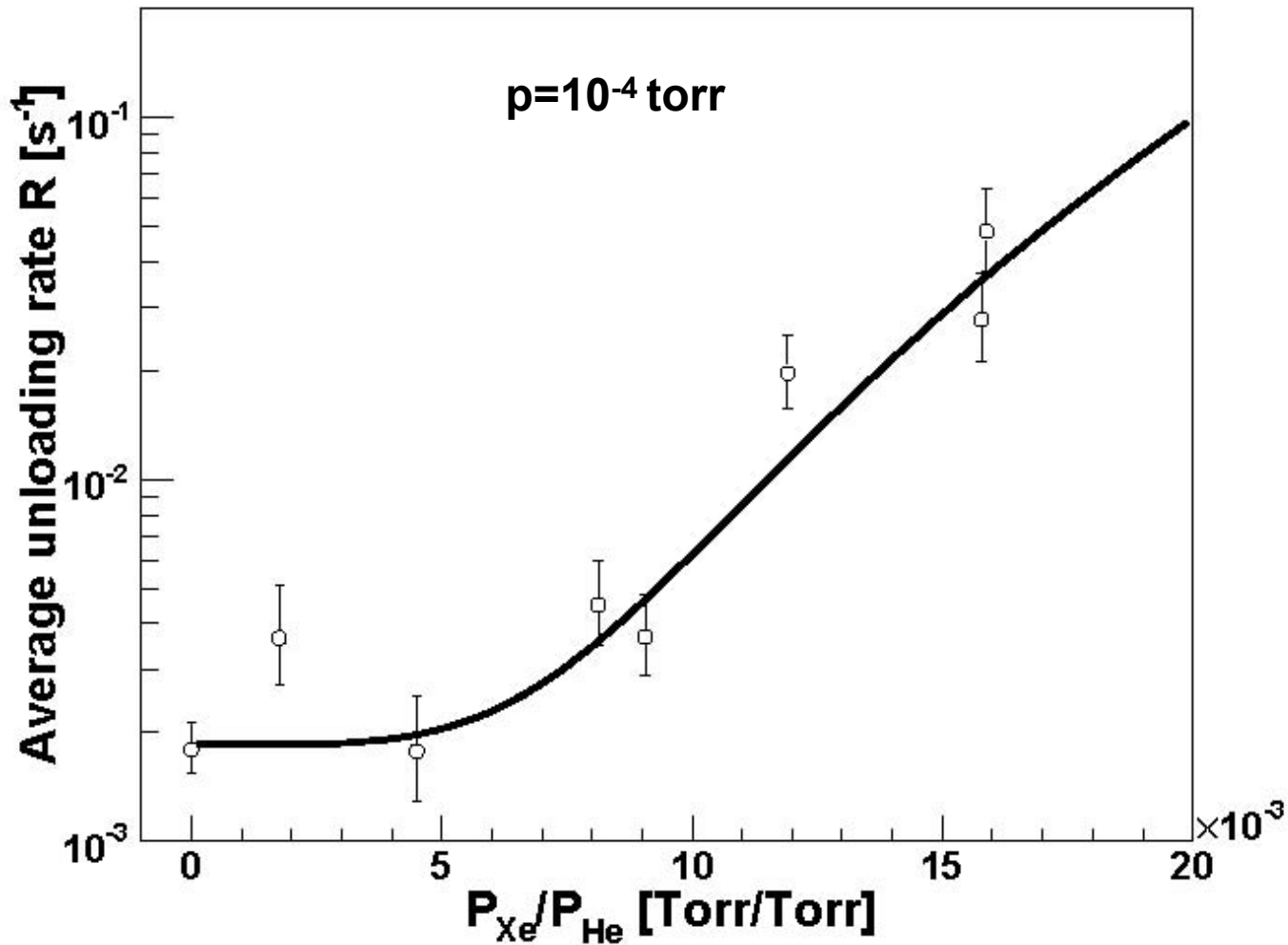
# First single ion detection in high pressure gas (He, Ar)



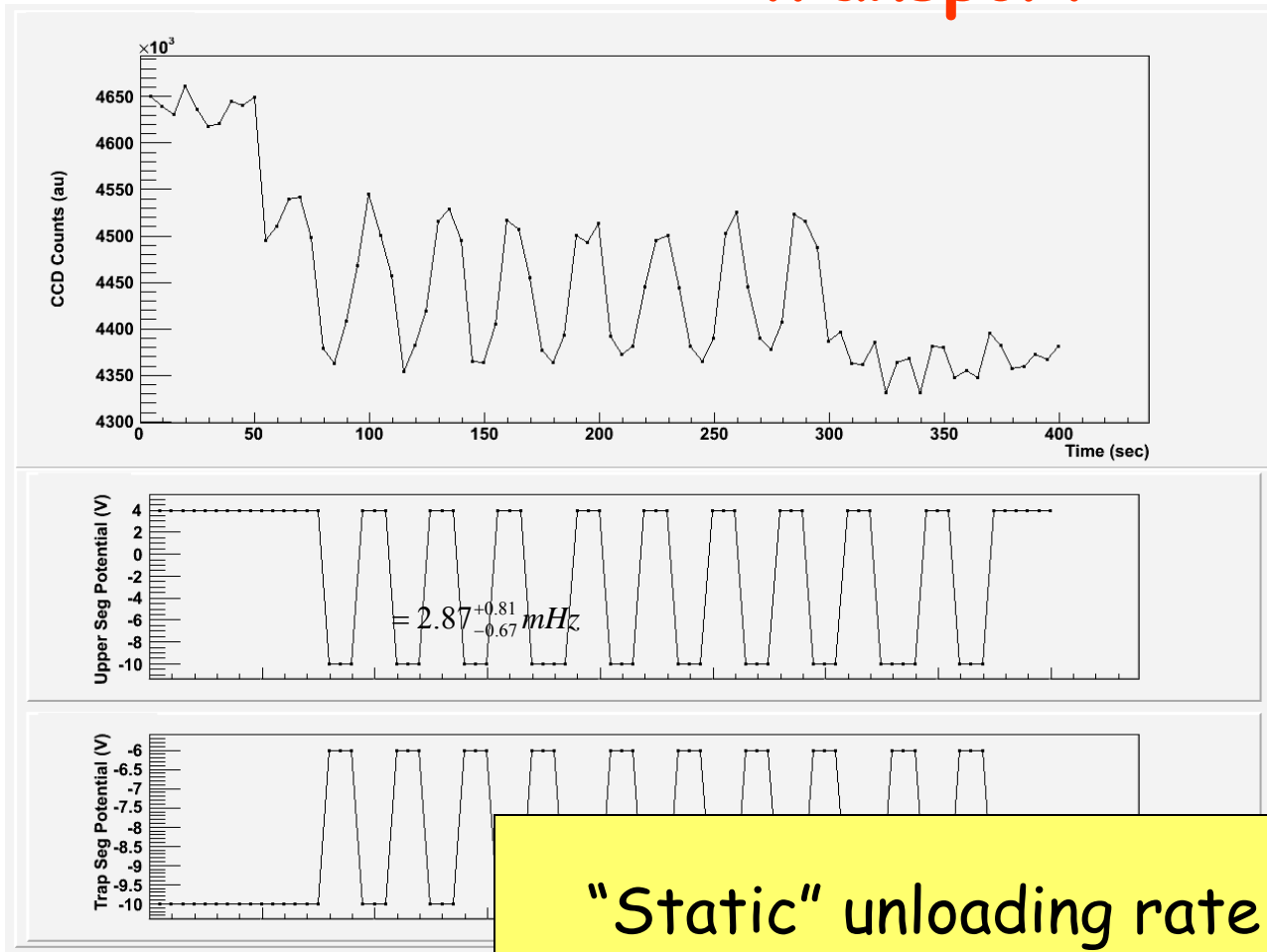
*M.Green et al. arXiv:0702122, Phys Rev A 76 (2007) 023404*  
*B.Flatt et al. arXiv:0704.1646, NIM A 578 (2007) 409*



Single ion spectroscopy & identification possible in some Xe atmosphere provided He is added to the trap



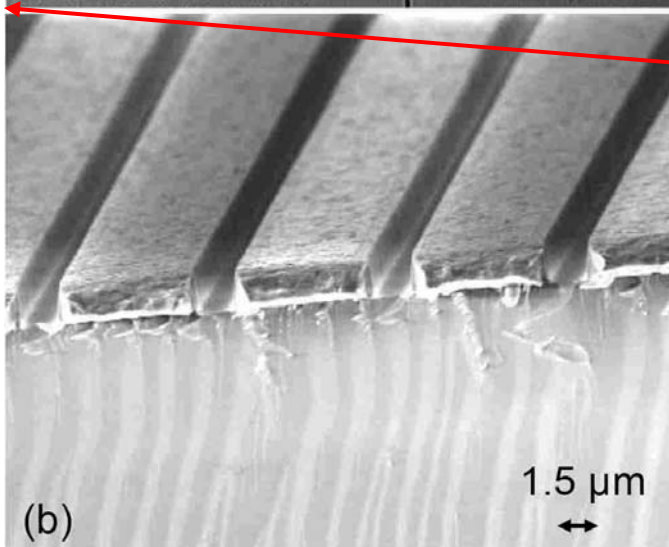
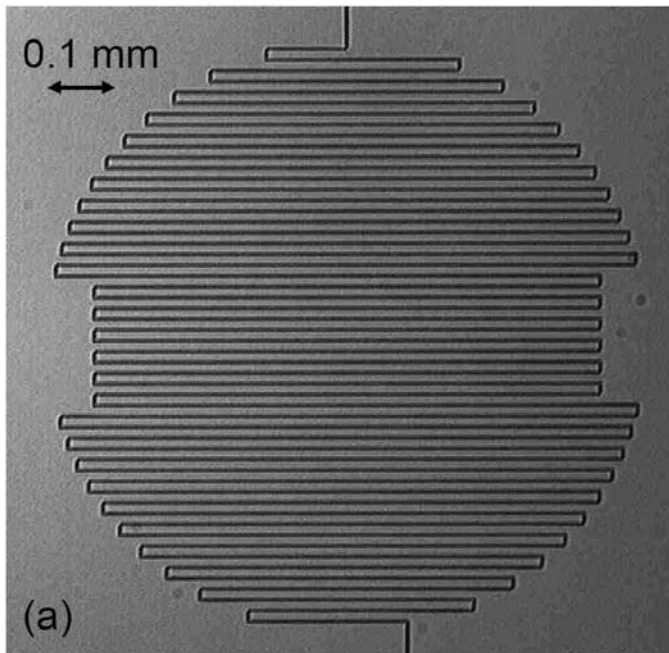
# Trap also allows for very "clean" ion transport



"Static" unloading rate =  $2.87^{+0.81}_{-0.67} \text{ mHz}$

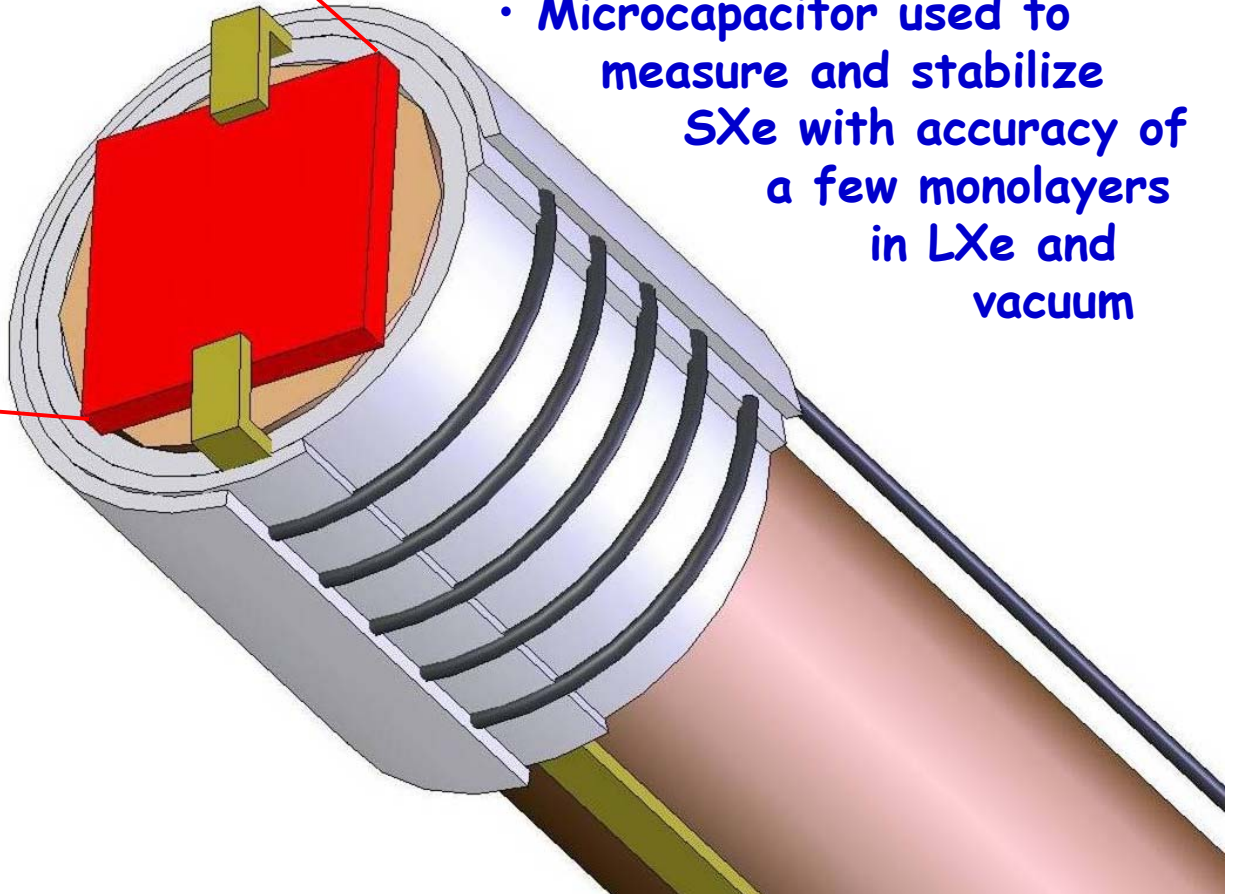
"Cycling" unloading rate =  $7.80^{+1.56}_{-2.34} \text{ mHz}$

Remaining challenge is the efficient transfer of single Ba ions from LXe to the ion trap



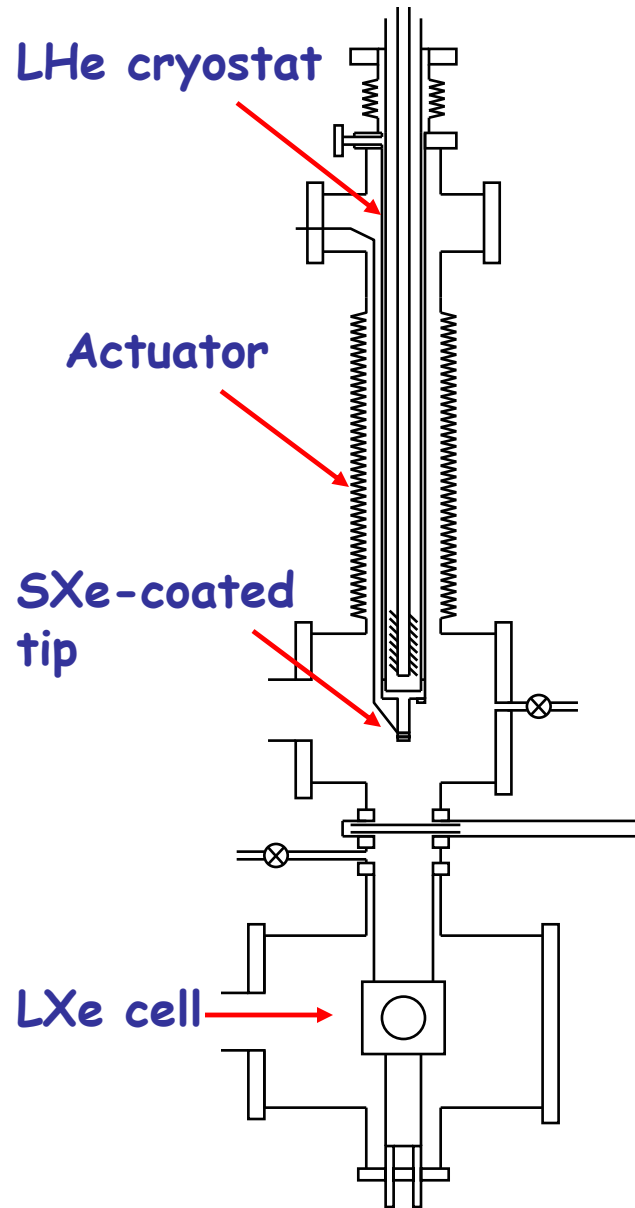
### Cryogenic dipstick

- Capture ion on SXe coating
- LHe cooling ( $\sim 20\text{K}$ ) to maintain stable SXe coating in  $10^{-8}$  torr vacuum
- Microcapacitor used to measure and stabilize SXe with accuracy of a few monolayers in LXe and vacuum

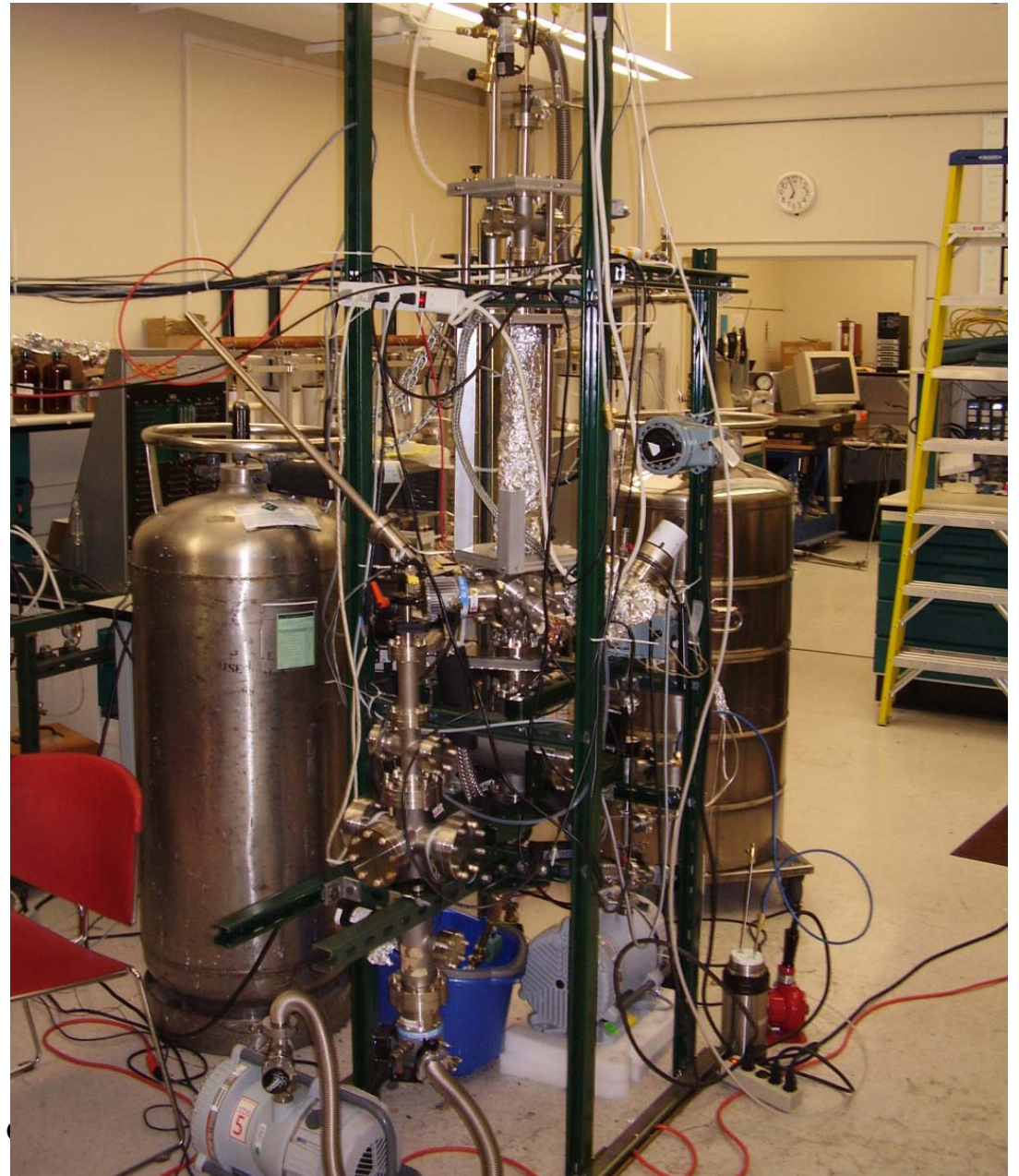


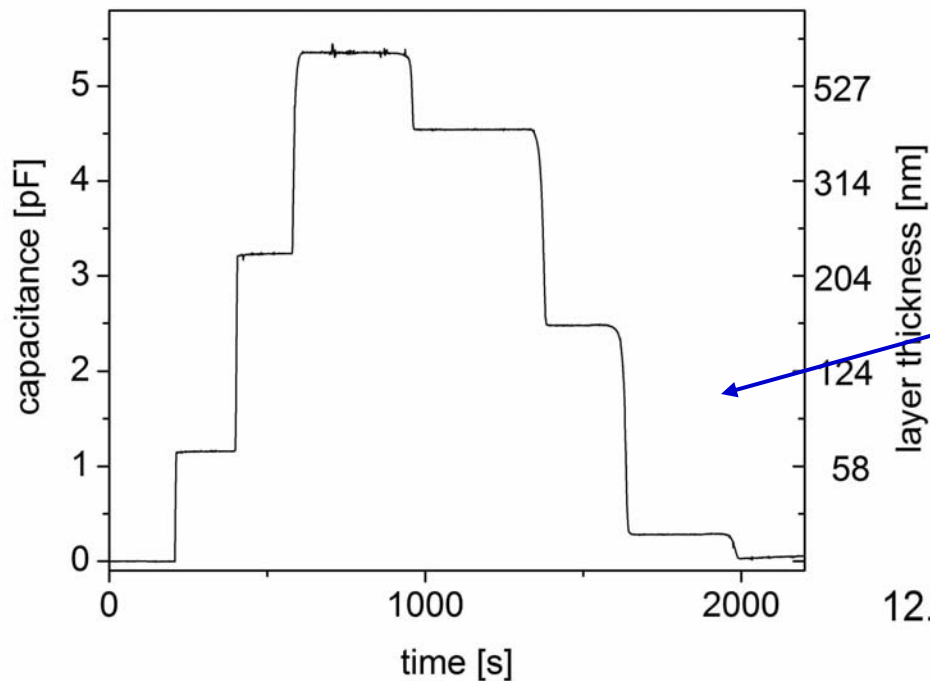


# Complete setup



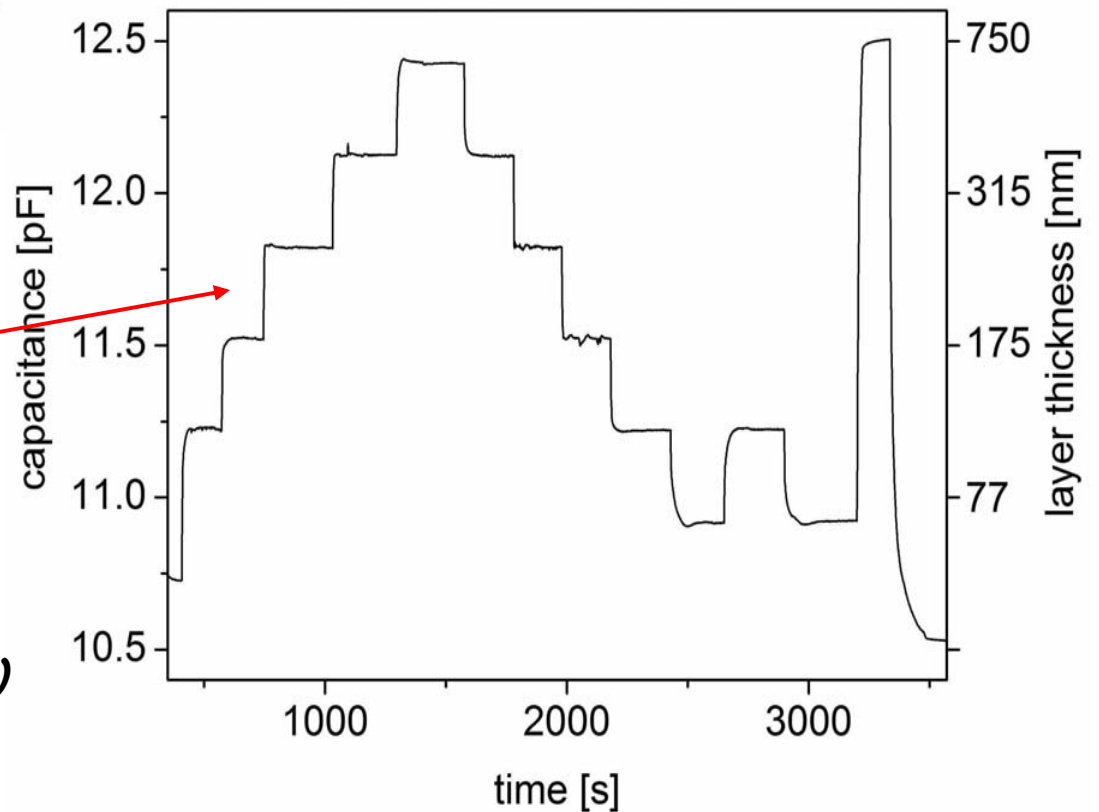
Berkeley Mar 19, 2009





Growing SXe layers from metered amounts of GXe in the vacuum chamber

Growing and maintaining SXe layers in a LXe bath with active feedback



*P. Fierlinger et al,*  
*Rev. Sci. Instr. 79, 045101 (2008)*

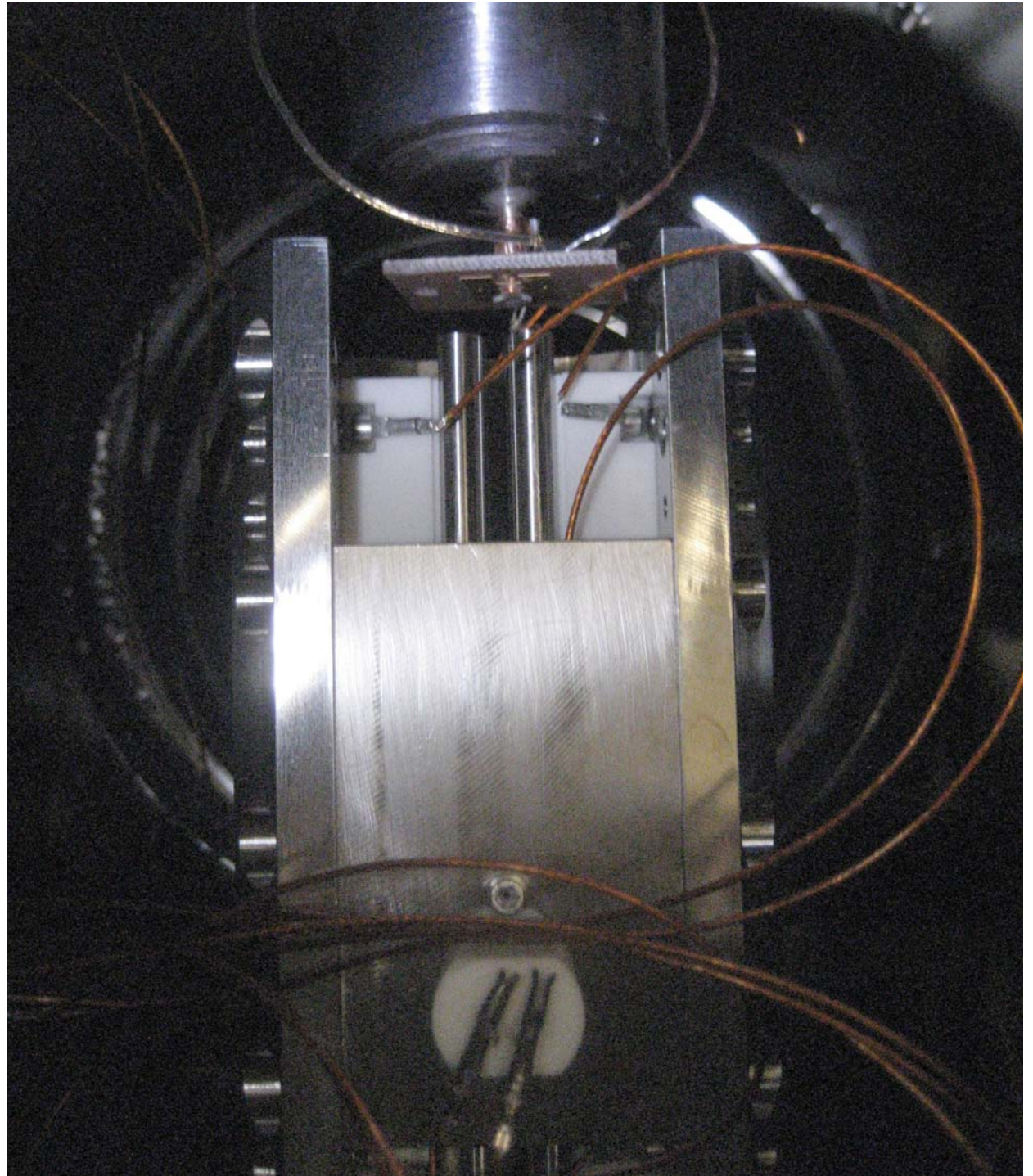


**In progress...**

**Shoot ions from the trap  
onto the cryotip and back  
into the ion trap**

**Measure the product  
of efficiencies**

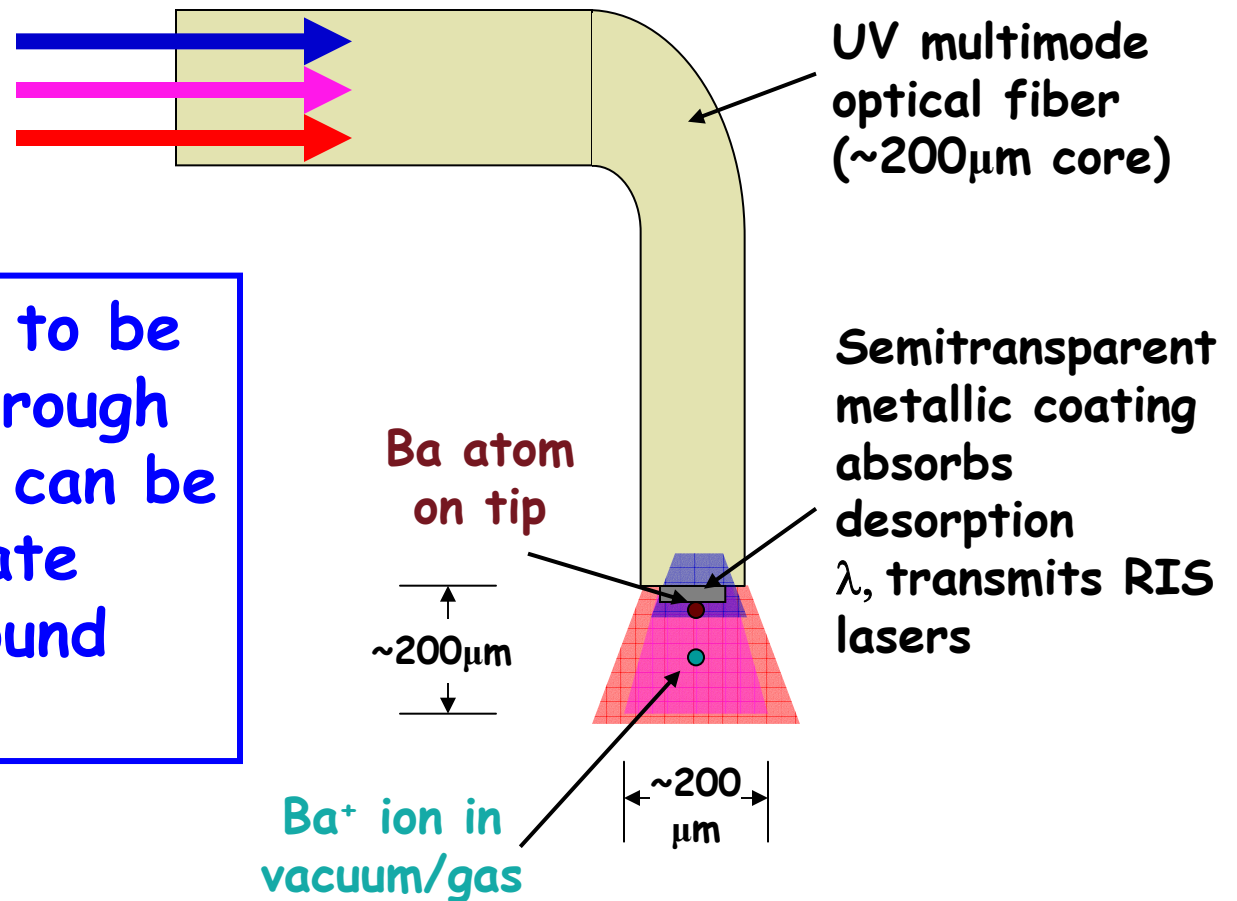
Berkeley Mar 19, 2009





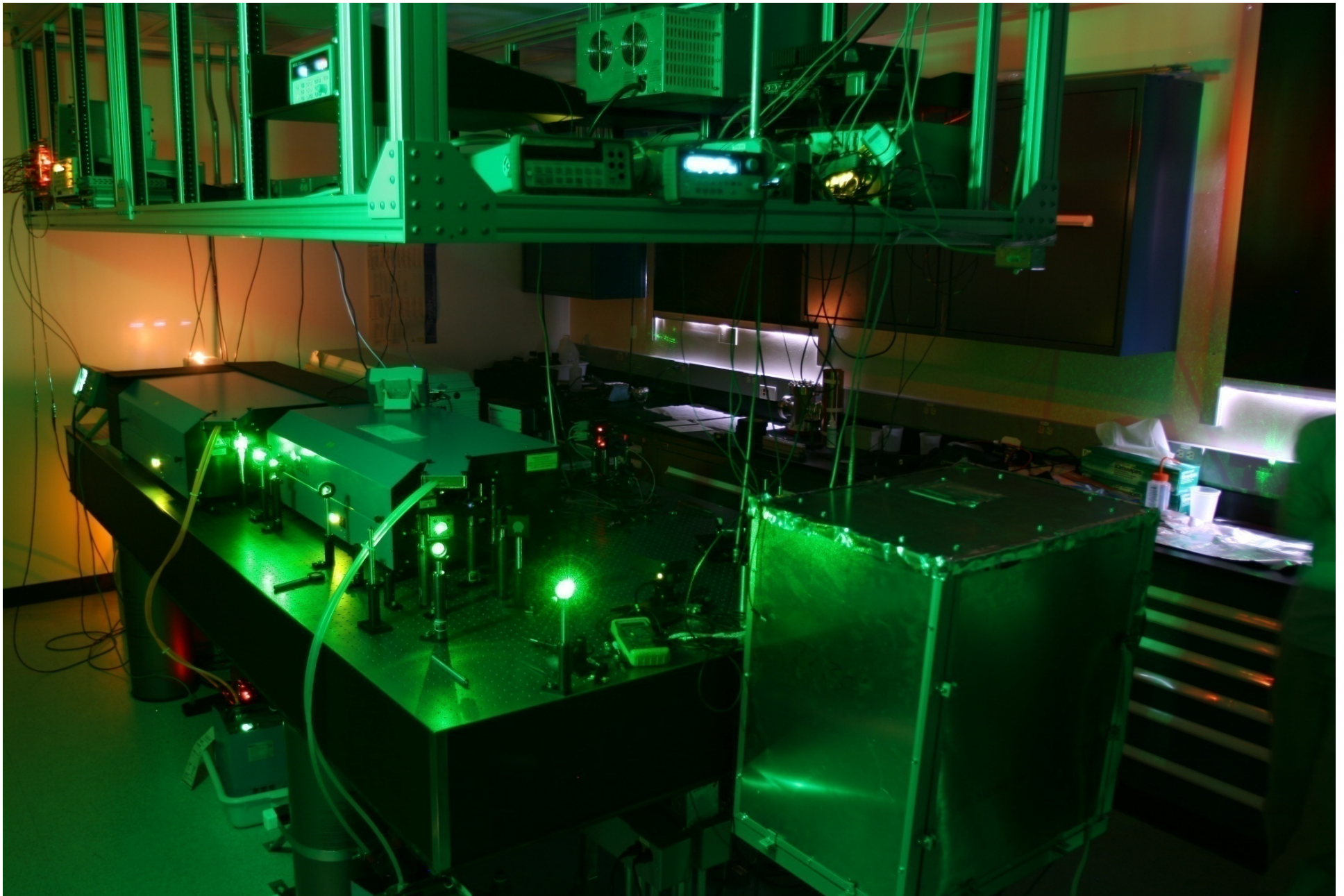
# An alternative way to transport the Ba ion:

This does not have to be necessarily done through a fiber, the lasers can be shot at the substrate where the ion is bound from the "outside"

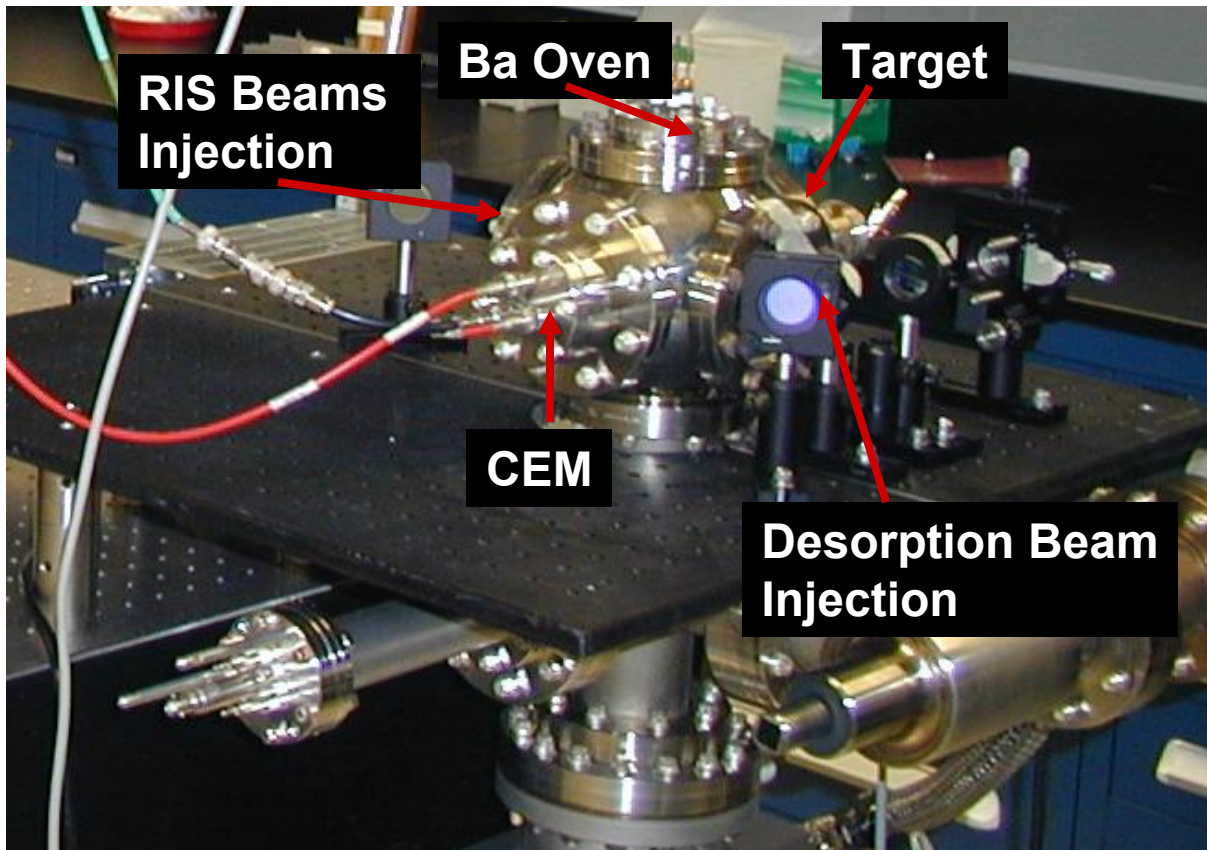


In this case *each step* can be documented to work with high efficiency in the literature !

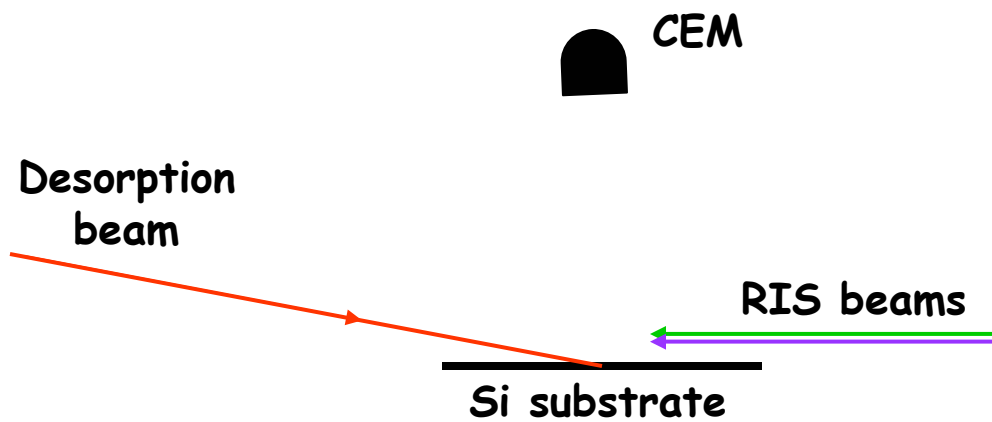
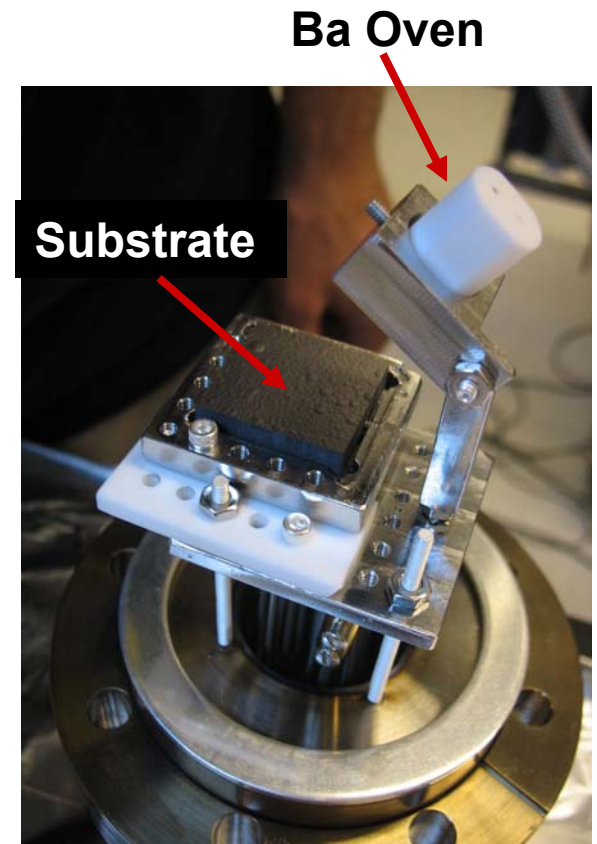
# Resonant ionization spectroscopy lab







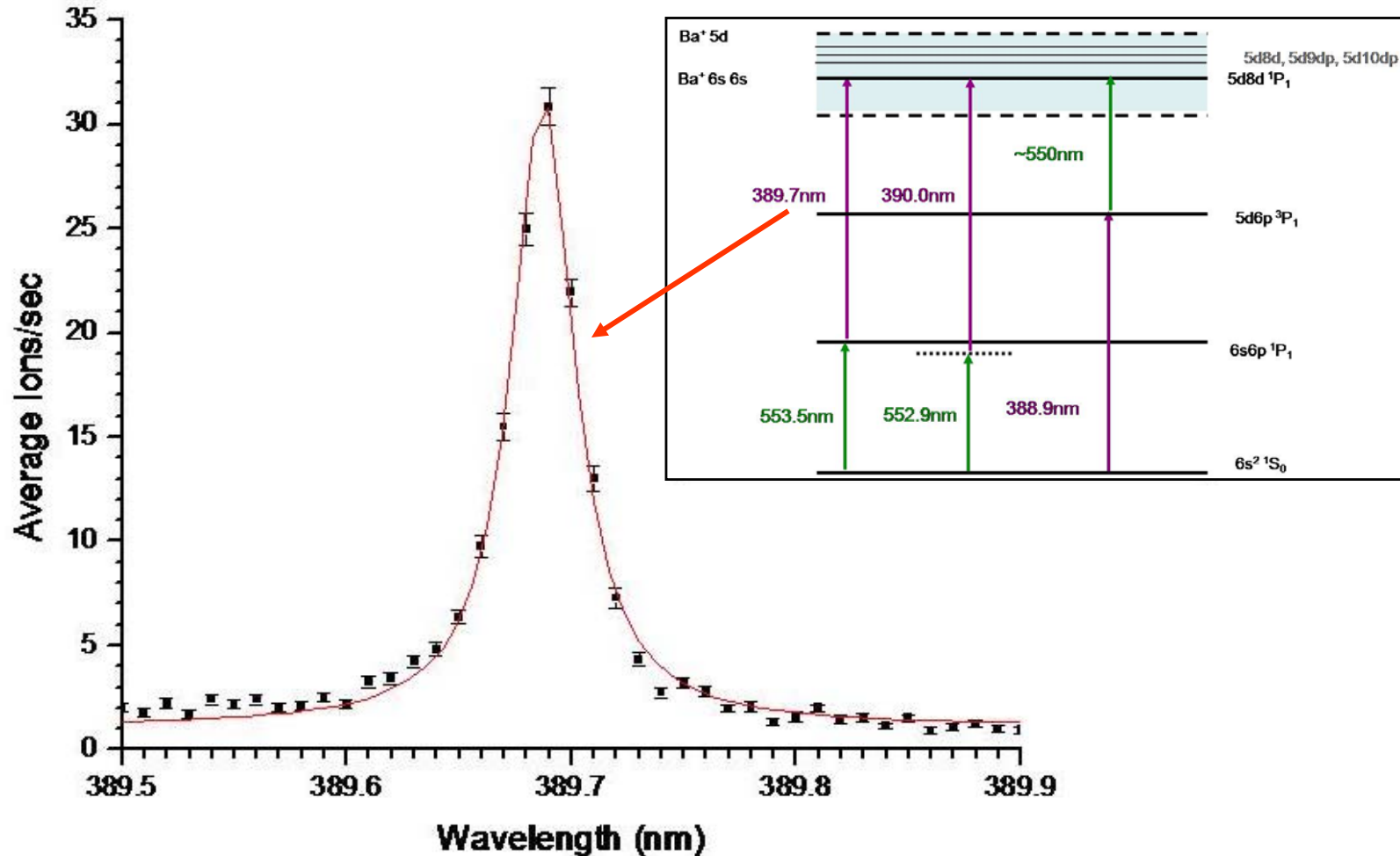
...and cell





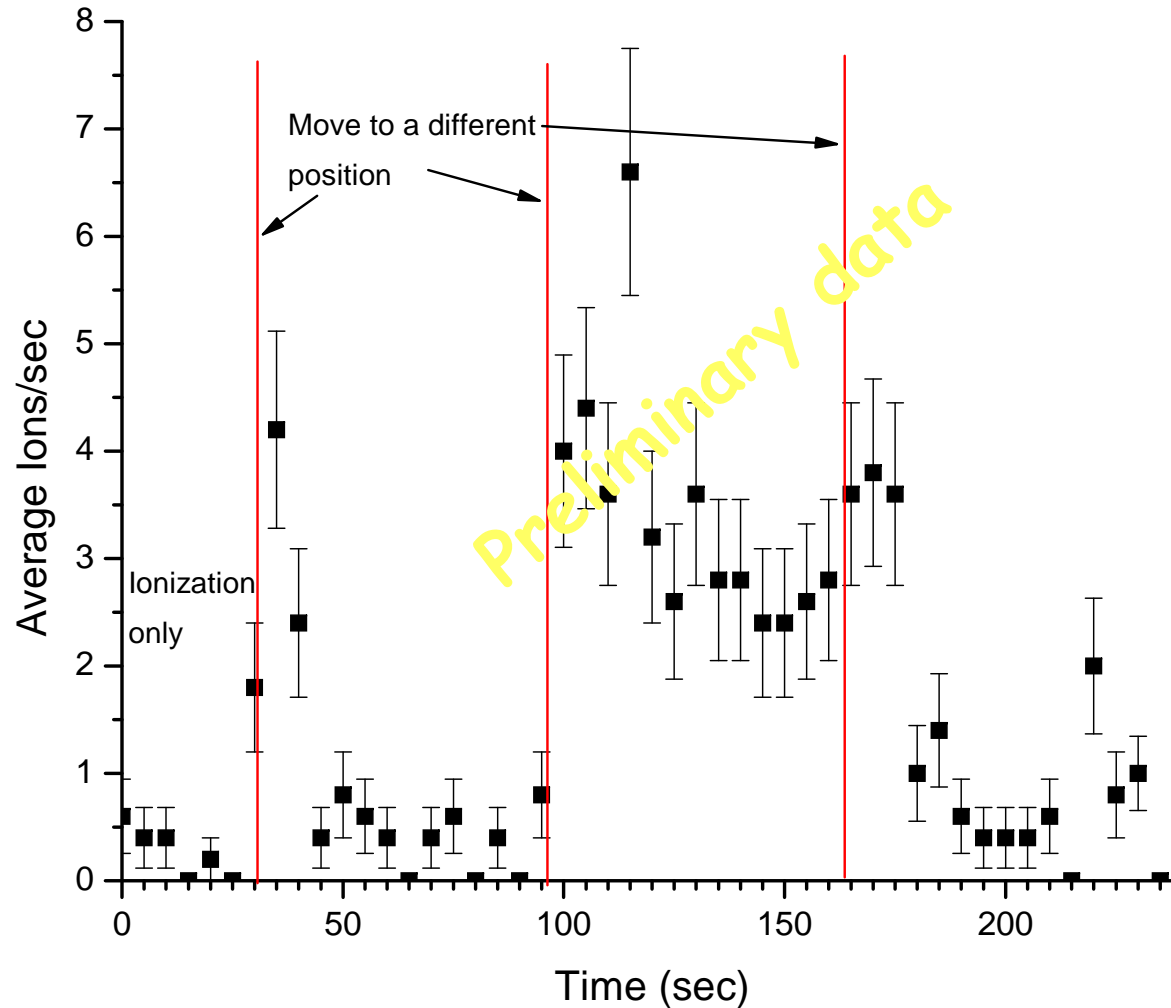
# RIS from Ba vapor phase: things work as they should

Wavelength Scan (Fine) of UV Laser at  $E = 0.24$  mJ/pulse.  
Green Laser at 553.5nm,  $E = 0.072$  mJ/pulse



For very thin Ba coatings a few shorts remove the Ba from a particular spot. **But the signal comes back elsewhere**

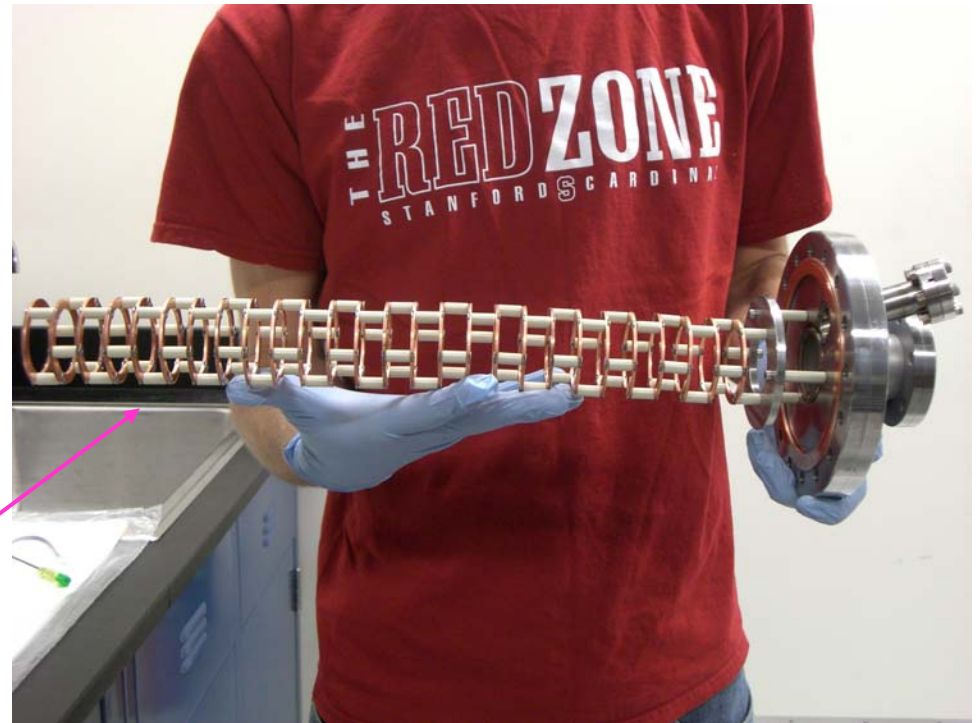
Desorption and Ionization of Barium, Scan Across Si Target  
Desorption E = 2 mJ/pulse, 1.8 usec delay, gate 3.8-5.2 usec



## Can desorb Ba and ionize in the same shot

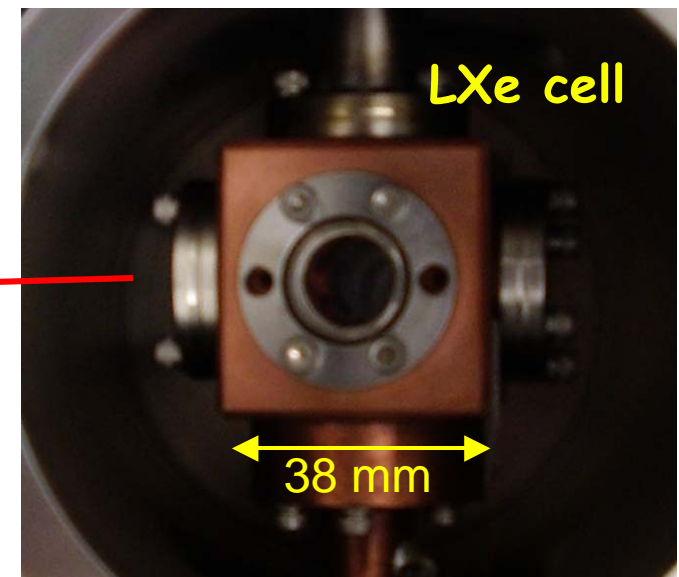
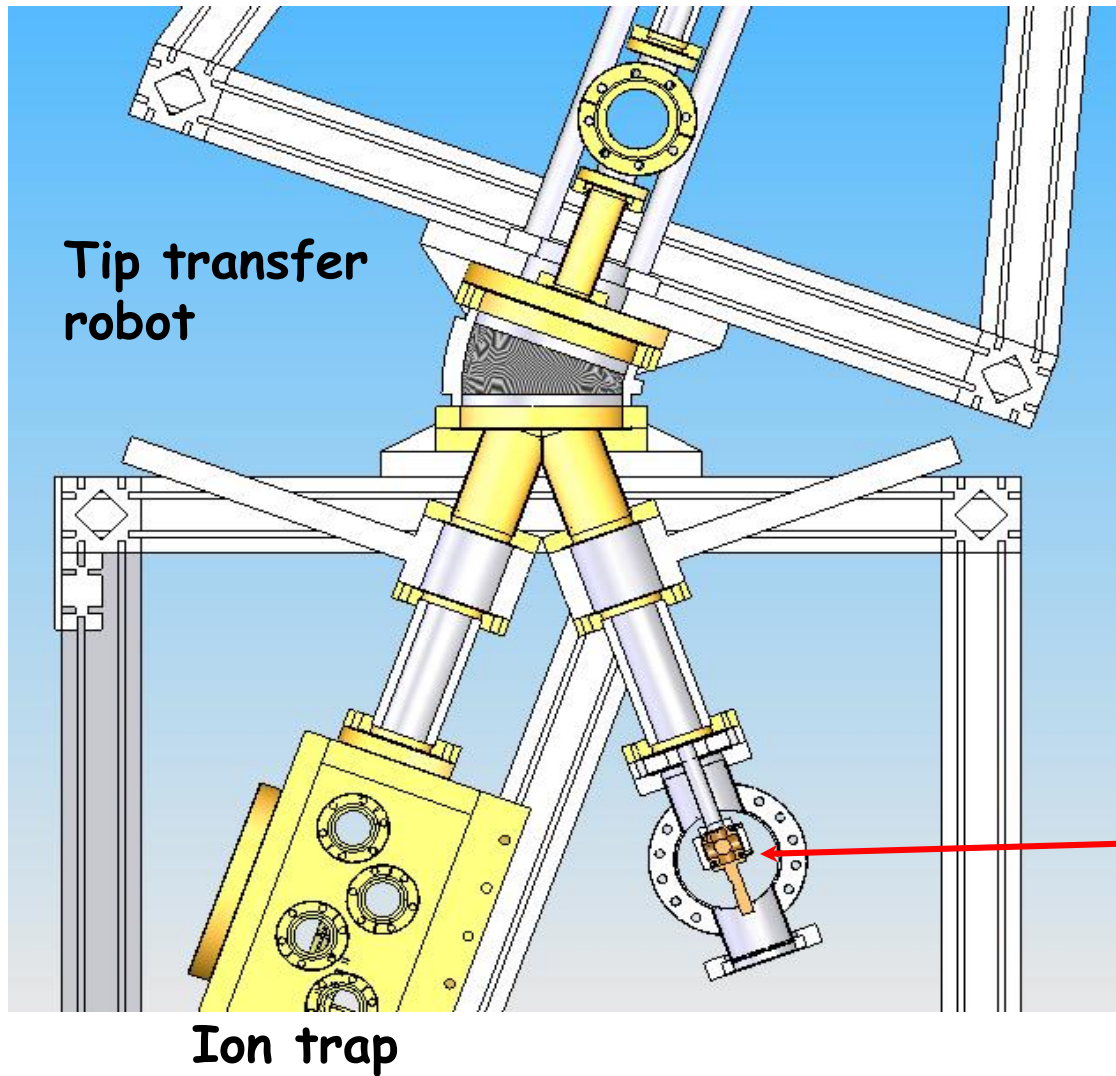
### To do:

- Learn how to deal with the fact that few shots may “burn” all the Ba from a place (this will not be a problem with single ions!)
- Understand better ablation thresholds
- Figure out a nice way to present data
- Install TOF spectrometer
- Build a, possibly tagged, single Ba source
- Work with single ions and measure efficiency



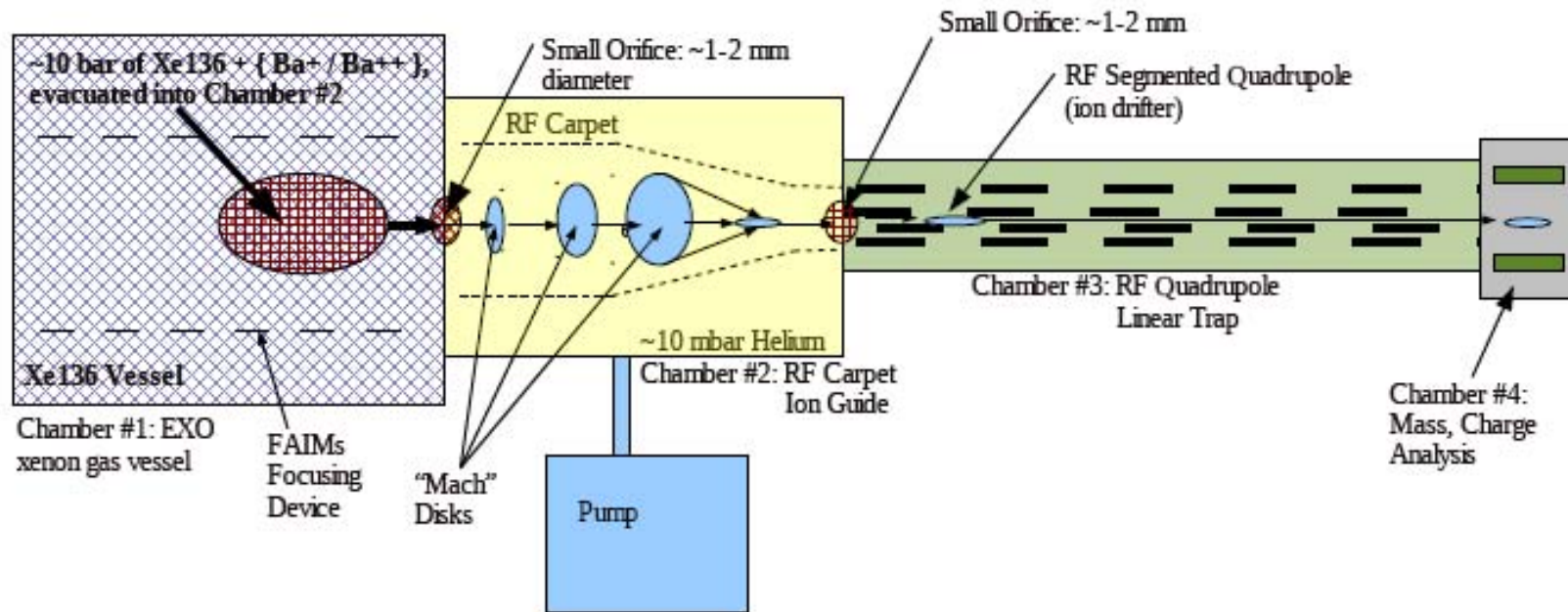


# Following step: transfer Ba-ions from LXe cell to ion trap



# What about Ba tagging from a GXe detector?

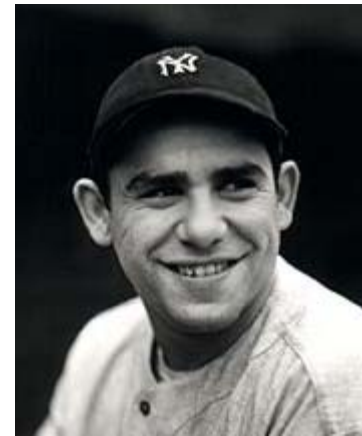
*We have just started to think about this:*



One may be able to borrow techniques from the radioactive ion beams community and build a chain of differentially pumped chambers, separated by nozzles, where the ion is confined by various means before reaching the identification stage in vacuum.

# Plans for the future

**It's tough to make predictions...  
especially about the future**





## *Some possibilities from possible EXO-200 outcomes*

Case 1: EXO-200 sees nothing, really

→ Build the largest possible EXO, with Ba tagging unless the EXO-200 background is spectacular

Go to DUSEL or to SNOlab

*Start in 5 to 8 yrs from now*

Case 2: EXO-200 has a 2 or 3 sigma peak at the right place

→ Build a new ~500kg chamber maybe in the same cryostat and same mine

*Start in <2 after signal is seen*

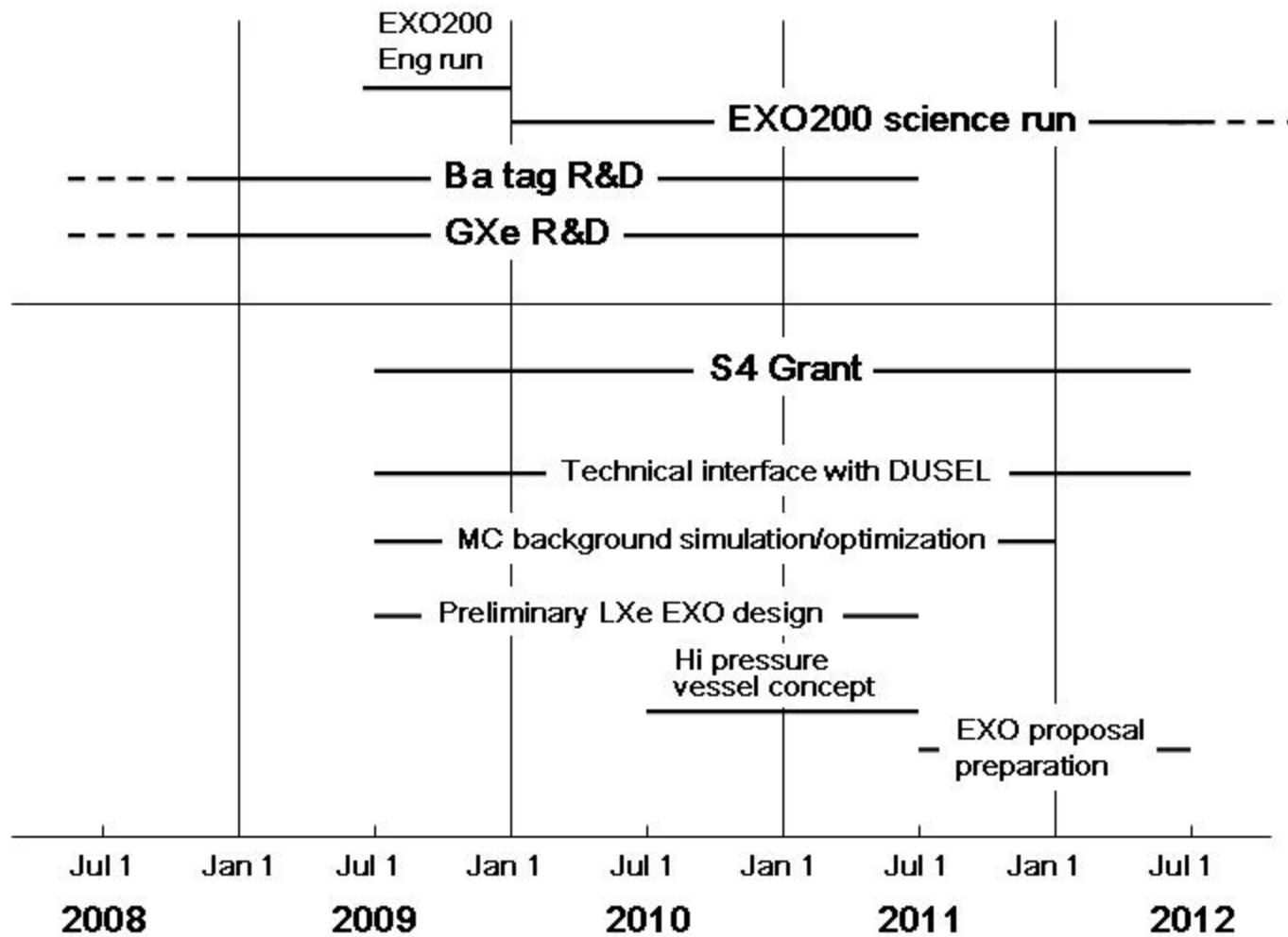
Case 3: EXO-200 has a clear signal

→ (open Champagne and) build a 1 bar GXe TPC for the very same 200kg of Xe to study angular correlations

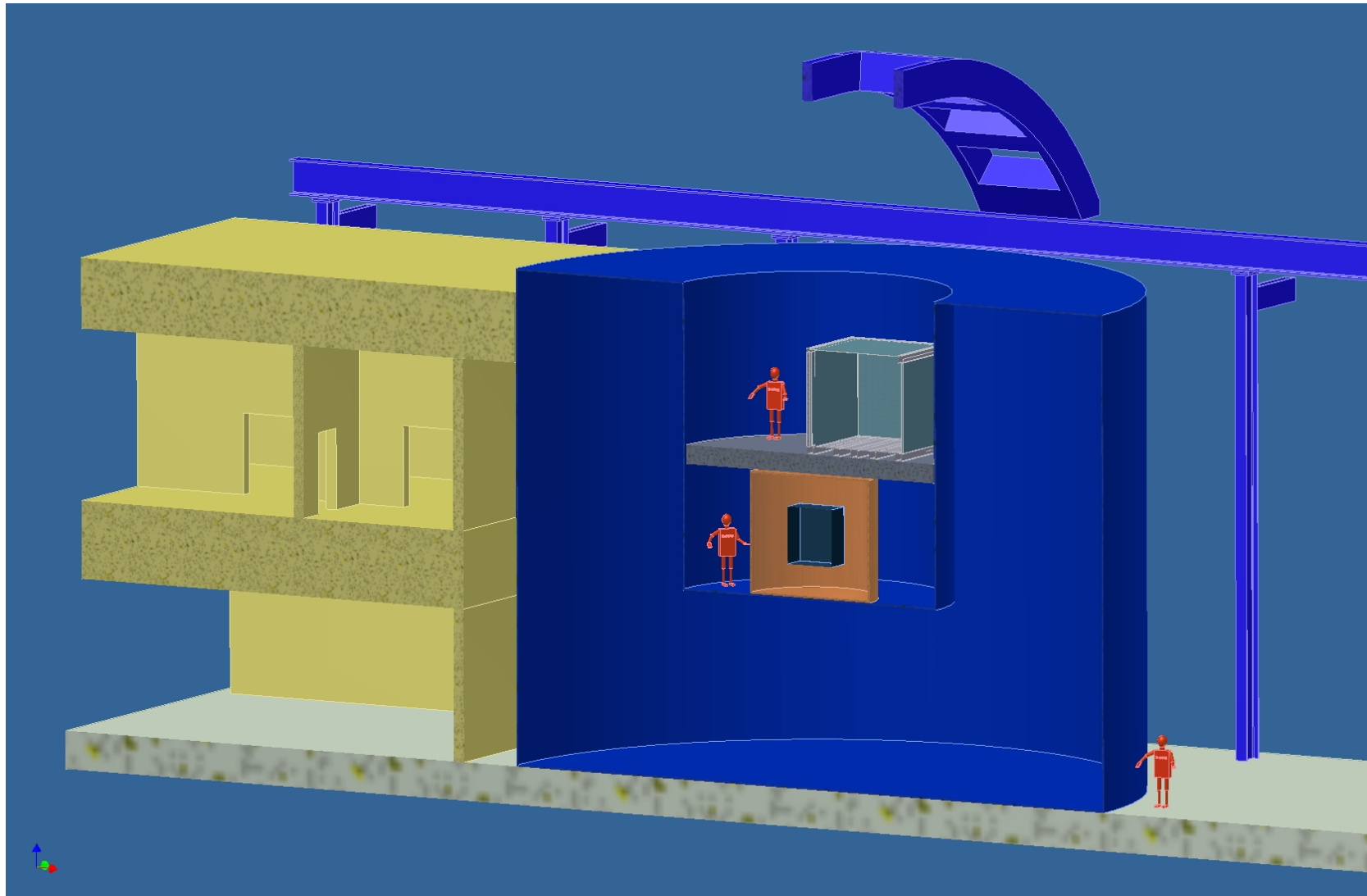
Possibly go to SNOlab

*Start in 2014*

## How does DUSEL S4 fit in with the rest?



# Outline of a LXe multi-ton EXO with Ba tagging in a DUSEL module





# Conclusions

*Over its glorious history neutrino physics has provided plenty of surprises and has required forays in many different areas of science and technology*

*EXO really belongs to this tradition!*

Isotope enrichment at an unprecedented scale (for science) is a reality

Data taking is approaching for EXO-200

Ba tagging for EXO is using bag of tricks borrowed from nuclear and particle physics, AMO and surface science

EXO is ramping up GXe R&D to arrive to the technology decision point with all of the information needed

